STAR CLUSTERS IN THE DISK OF ANDROMEDA

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 5×10^{4} 0.005 **Our aim** was to investigate the properties of star cluster <100 Mvr 100-400 Mvr population in the M31 galaxy. 0.004 <u></u>≥0.003 M_c/M_{\odot} We analyzed 2660 star clusters in M31 using integrated <u>\$</u>0.002 aperture HST photometry and stochastic cluster models. 0.001 Number of clusters in the following age intervals: 726 (<100 Myr), 1141 (100–400 Myr), 494 (400–1200 Myr). 0.000 10 16 19 10 r, kpc r, kpc 7.0 <100 Myr a) We determined: 100 6.0 function) Characteristic cluster mass (Schechter decreases radially from the M31 center. ([⊙]W/W)6ol A "ring" of star formation at 15 kpc is prominent only for younger (<100 Myr) clusters. 100-400 Myr b) 100 15 Clusters form in the same star forming regions for extended periods – indicating their slow dispersion. A strong star-forming episode in the M31 NE part occurred ~200 Myr ago. 10

log(t/yr)

X. kpc