## THE ORIGINAL COMPOSITION OF THE GAS FORMING FIRST-POPULATION STARS IN GLOBULAR CLUSTERS

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Università degli Studi di Padova

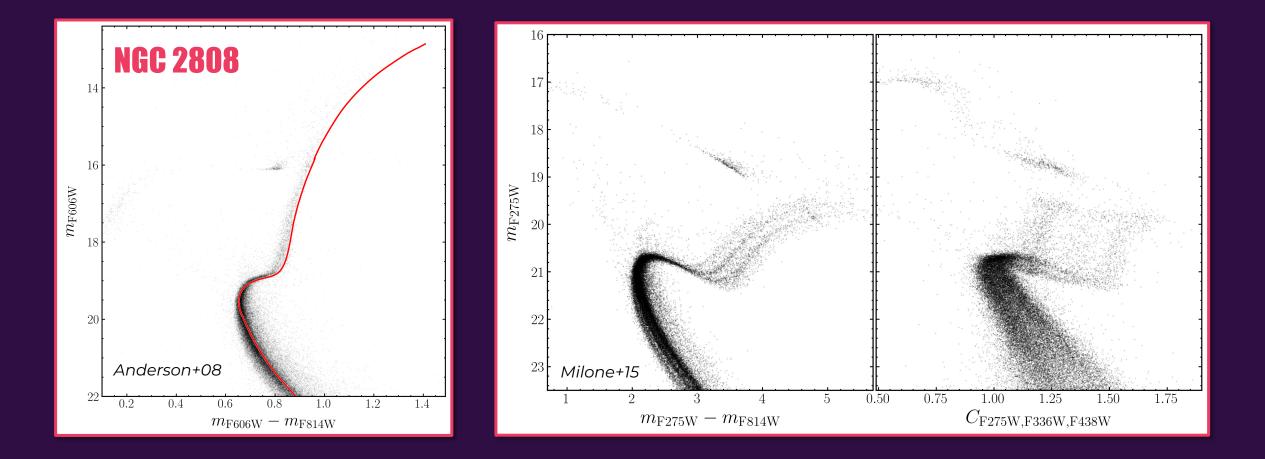


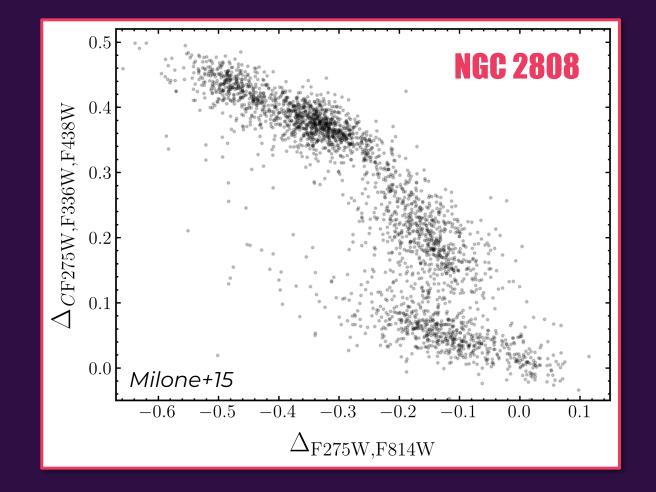


## **THE ORIGINAL COMPOSITION OF PRIMORDIAL CLOUDS**







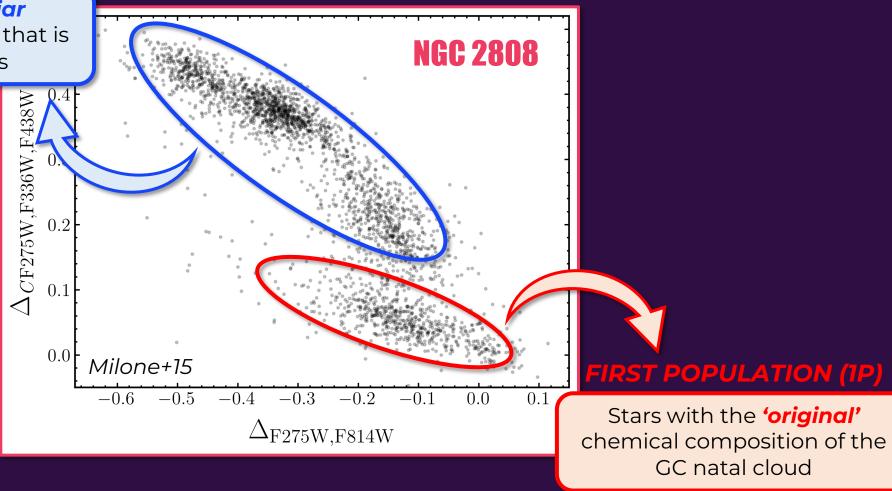


**MODEST 24** 



#### **SECOND POPULATION (2P)**

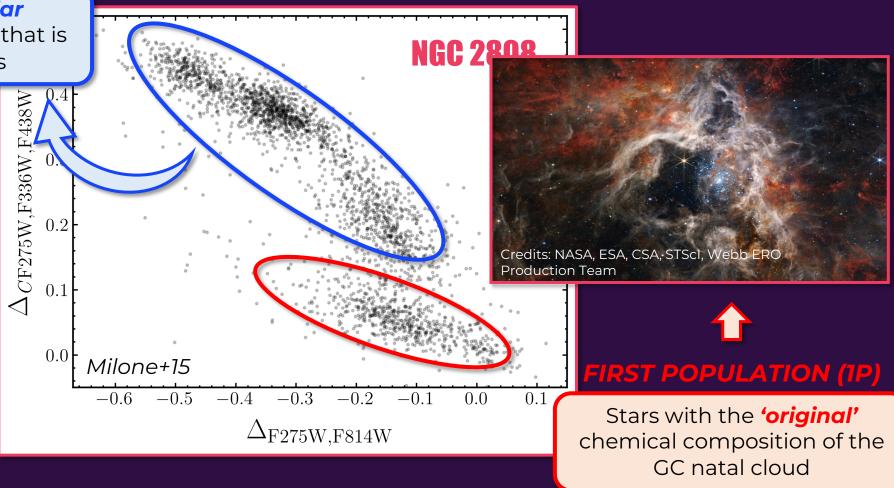
Stars with a **peculiar** chemical composition that is found only in GCs





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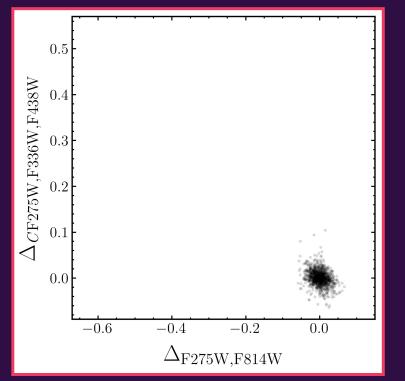
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**1P** stars are **not** consistent with a **simple stellar population** as they exhibit extended or bimodal sequences in the ChM

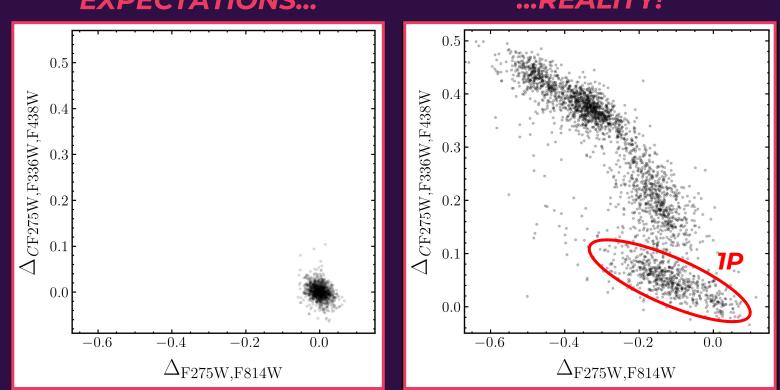




#### **EXPECTATIONS...**

**IP** stars are **not** consistent with a **simple stellar population** as they exhibit extended or bimodal sequences in the ChM



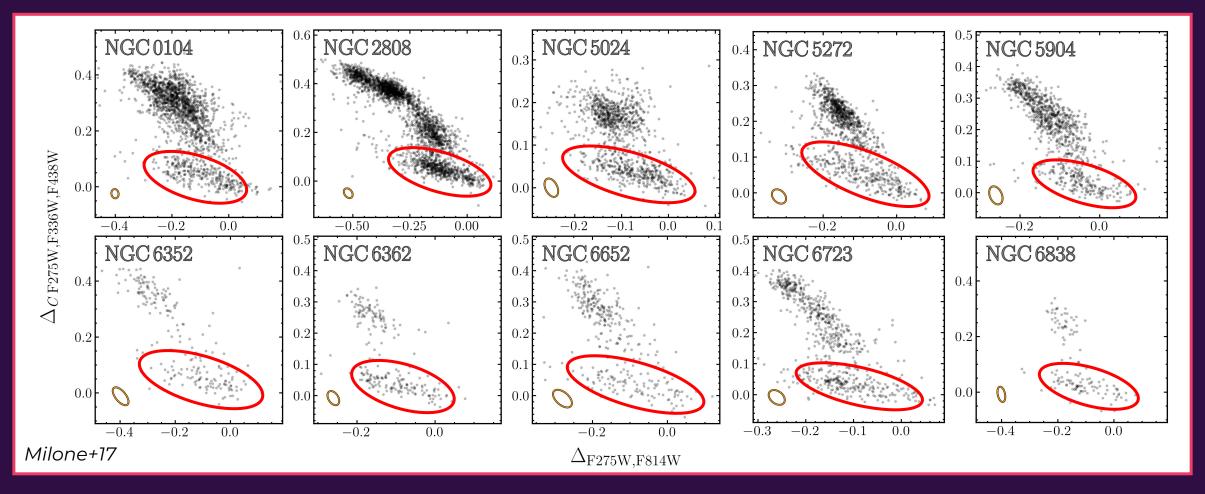


**EXPECTATIONS...** 

...REALITY!

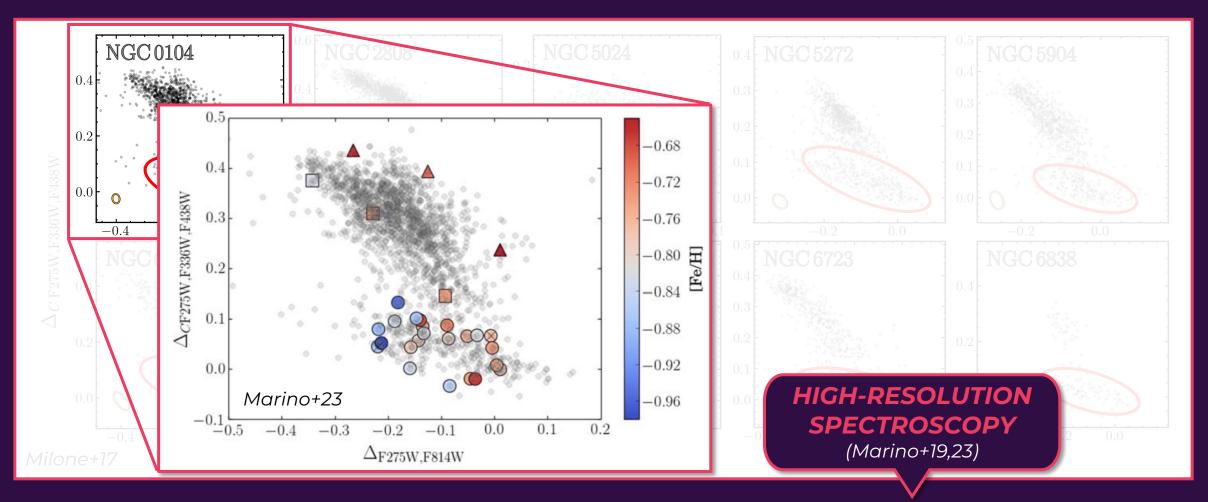
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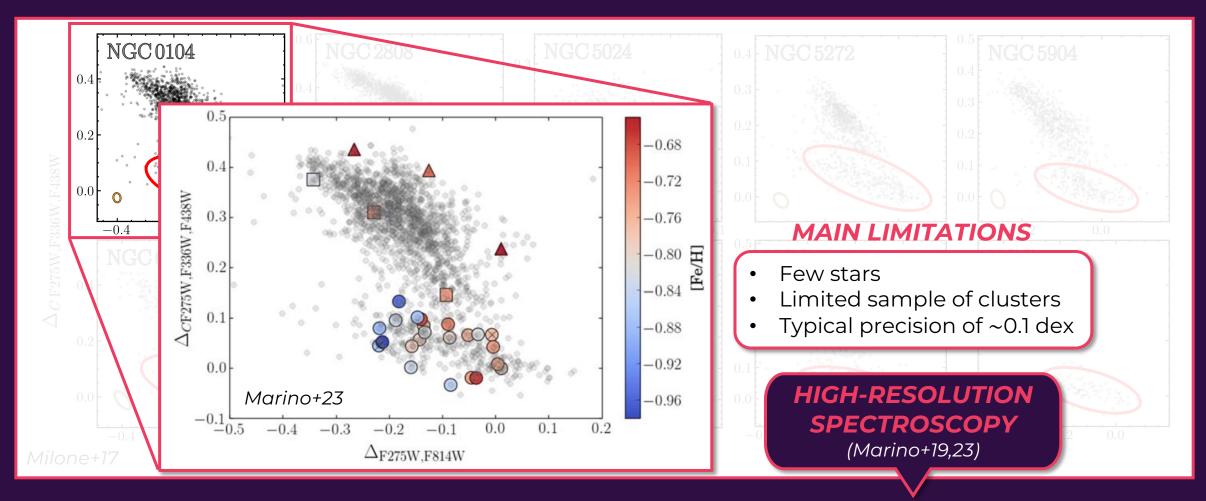
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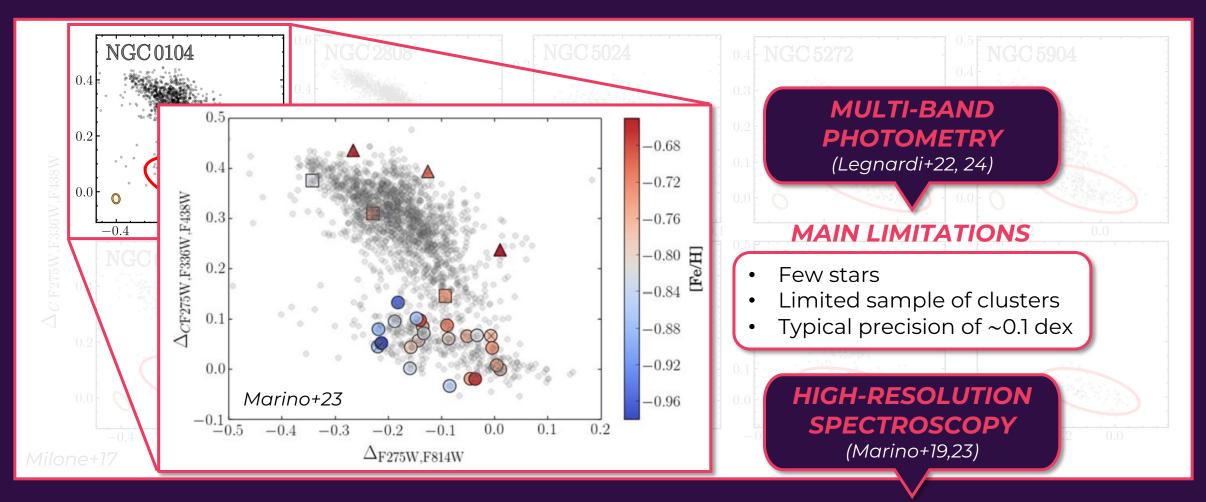
The extended 1P sequence is associated with *metallicity variations* inherited from the pristine material out of which they formed





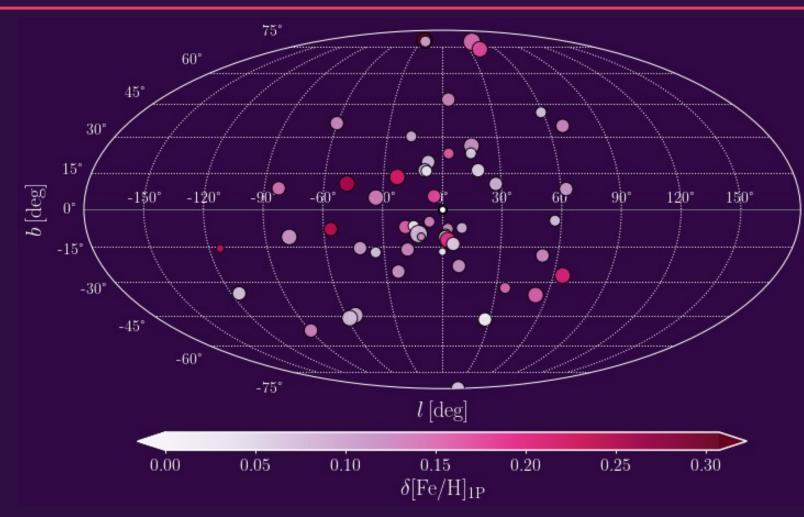
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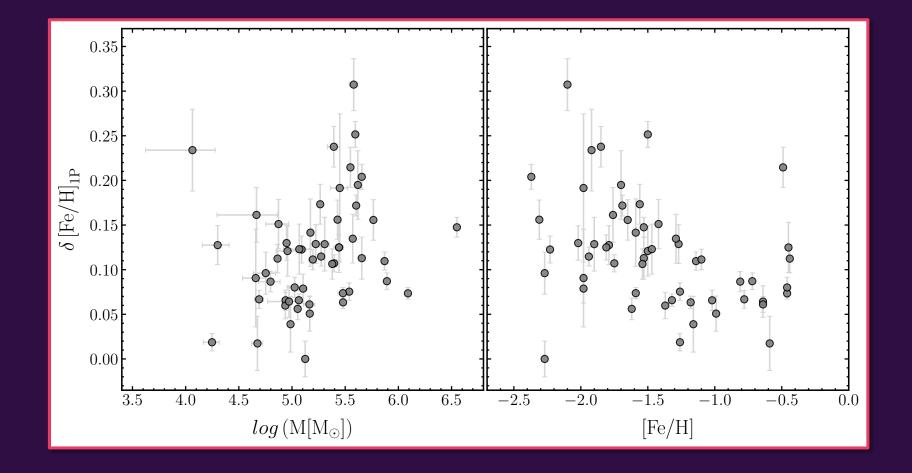




Results based on Legnardi+22

The analysis of ChMs in a large sample of **55 Galactic GCs** demonstrates that [Fe/H] variations exhibit a *high level of variability* within the Galaxy





The analysis of ChMs in a large sample of **55 Galactic GCs** demonstrates that [Fe/H] variations correlate with **cluster mass** and **metallicity** 





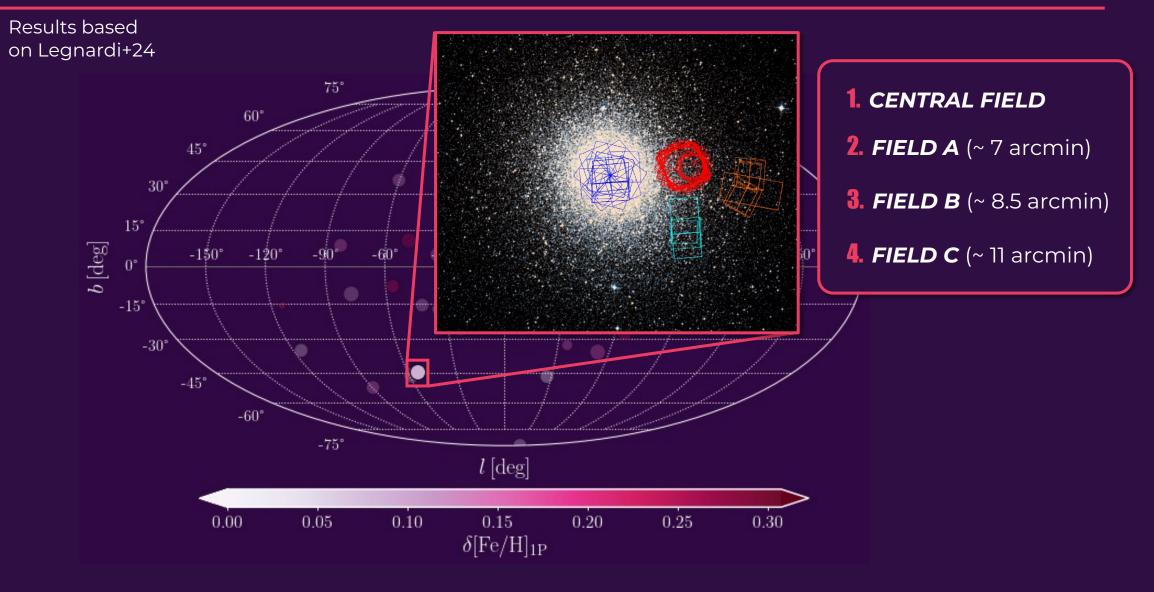
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The metal content of 1G stars has been examined in the **central regions** of **old** Galactic GCs with **multiple stellar populations** 



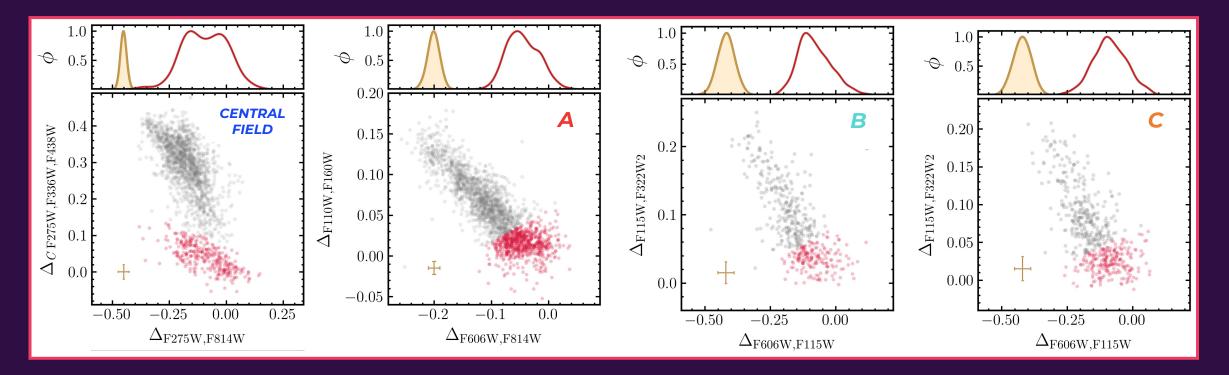
## AN IN-DEPTH ANALYSIS OF 1P STARS IN 47 TUCANAE





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Results based on Legnardi+24

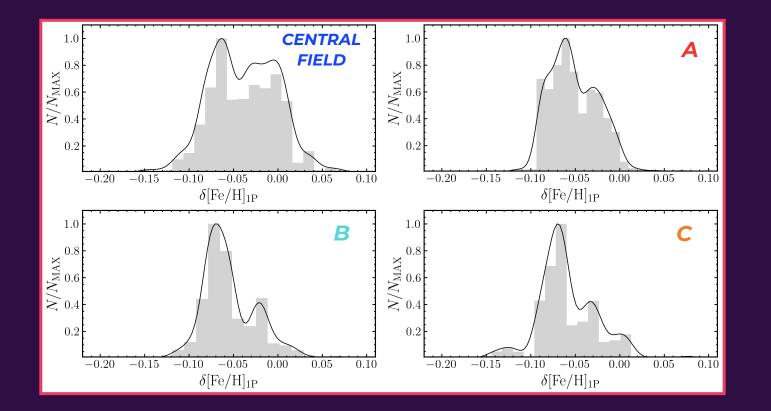


In all ChMs, 1P stars define an **extended** sequence in the direction of the x-axis, which is associated with **metallicity variations** within the primordial cloud





Results based on Legnardi+24



#### Metallicity distributions

The [Fe/H] distributions span similar intervals of  $\sim 0.1$  dex, ranging from  $\sim -0.15$  to  $\sim 0.05$ 

#### **Radial distribution**

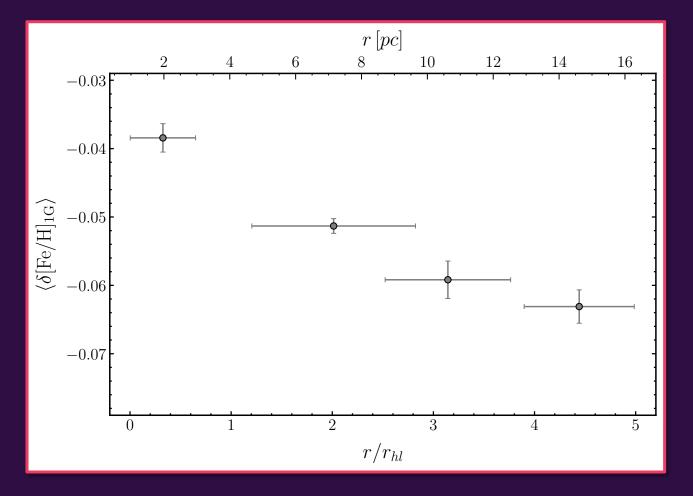
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ARE THE 1G METALLICITY SPREADS EXCLUSIVE TO OLD GALACTIC GCs WITH MULTIPLE STELLAR POPULATIONS?

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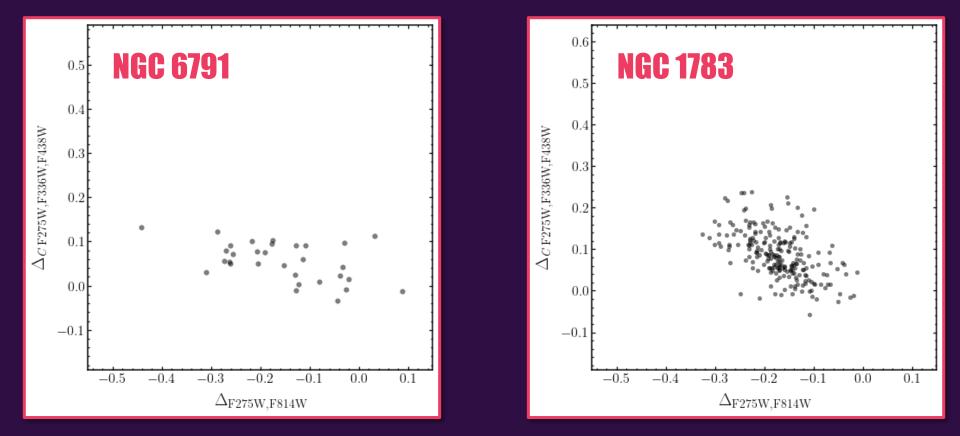
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## SIMPLE POPULATION CLUSTERS: NGC 6791 AND NGC 1783

Results based on Legnardi+24



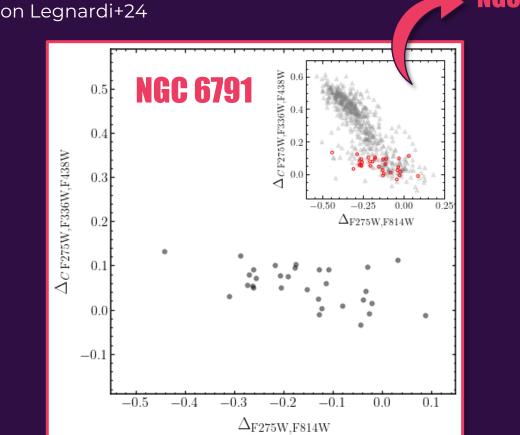
The ChMs of both clusters display **one sequence** of stars, resembling the 1G sequence of GCs with multiple populations



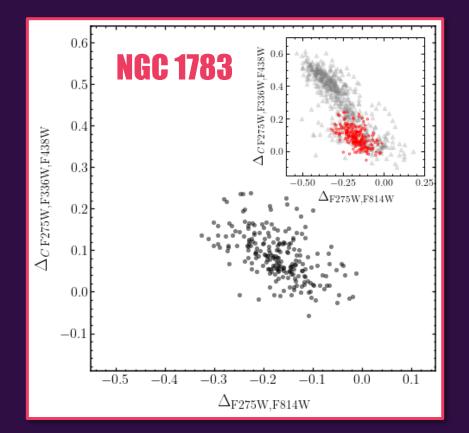


**Results** based

## SIMPLE POPULATION CLUSTERS: NGC 6791 AND NGC 1783



**NGC 6624:** GC with ~ [Fe/H]



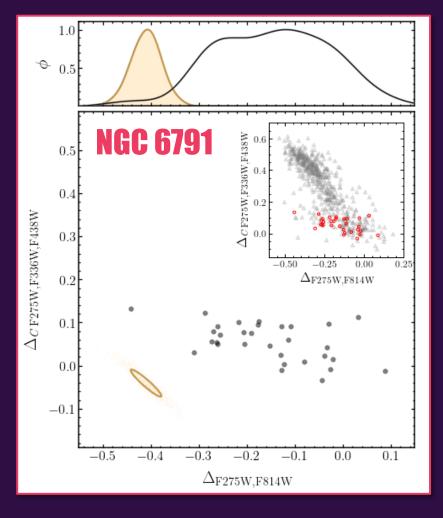
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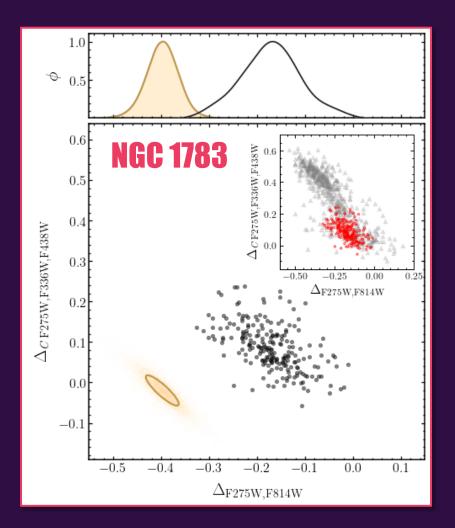




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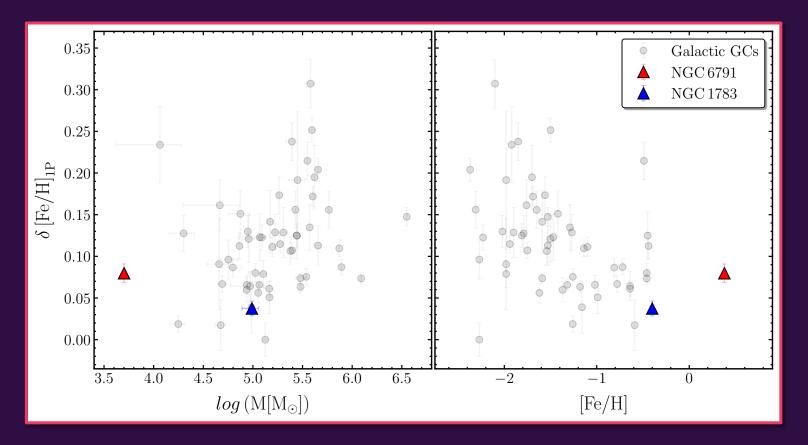






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Metallicity variations of NGC 6791 and NGC 1783 are **comparable** to those observed in **metal-rich GCs** with **similar masses** 



### SUMMARY

### **47 TUCANAE**

#### **GALACTIC GCs**

- The ChM analysis of 55 Galactic
  GCs revealed metallicity
  spreads in the formation gas of
  IP stars.
- Metallicity variations are *highly variable* within the Galaxy,
  ranging from ~ 0.00 to ~ 0.30 dex.
- Metallicity variations correlate with cluster *mass* and *metallicity*.



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- The cluster center shows *higher* metallicity spreads than the outer regions.
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#### **MORE DETAILS**

- **Legnardi et al. 2022**, MNRAS, 513, 735
- **2. Legnardi et al. 2024**, A&A, 687, A160



http://progetti.dfa.unipd.it/ GALFOR



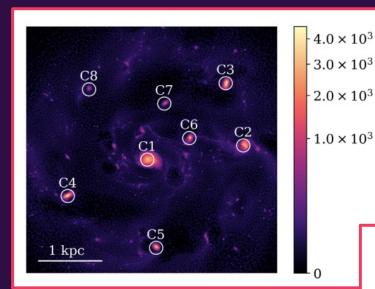
## **METALLICITY VARIATIONS IN GALACTIC GCS**

Cluster ID	$\delta [Fe/H]_{1G}$	Cluster ID	$\delta [Fe/H]_{1G}$
NGC 0104	$0.087 \pm 0.009$	NGC 6304	$0.125 \pm 0.028$
NGC 0288	$0.066 \pm 0.007$	NGC 6341	$0.156 \pm 0.022$
NGC 0362	$0.075 \pm 0.010$	NGC 6352	$0.064 \pm 0.018$
NGC 1261	$0.129 \pm 0.022$	NGC 6362	$0.051 \pm 0.020$
NGC 1851	$0.063 \pm 0.007$	NGC 6366	$0.017 \pm 0.030$
NGC 2298	$0.234 \pm 0.046$	NGC 6397	$0.130 \pm 0.019$
NGC 2808	$0.110 \pm 0.010$	NGC 6441	$0.074 \pm 0.007$
NGC 3201	$0.142 \pm 0.038$	NGC 6496	$0.080 \pm 0.011$
NGC 4590	$0.123 \pm 0.015$	NGC 6535	$0.128 \pm 0.022$
NGC 4833	$0.238 \pm 0.023$	NGC 6541	$0.125 \pm 0.014$
NGC 5024	$0.307 \pm 0.029$	NGC 6584	$0.121 \pm 0.028$
NGC 5053	$0.096 \pm 0.024$	NGC 6624	$0.112 \pm 0.016$
NGC 5139	$0.148 \pm 0.011$	NGC 6637	$0.061 \pm 0.009$
NGC 5272	$0.252 \pm 0.015$	NGC 6652	$0.087 \pm 0.011$
NGC 5286	$0.172 \pm 0.012$	NGC 6656	$0.195 \pm 0.038$
NGC 5466	$0.091 \pm 0.055$	NGC 6681	$0.056 \pm 0.012$
NGC 5897	$0.129 \pm 0.030$	NGC 6717	$0.019 \pm 0.010$
NGC 5904	$0.135 \pm 0.027$	NGC 6723	$0.111 \pm 0.011$
NGC 5927	$0.215 \pm 0.023$	NGC 6752	$0.106 \pm 0.017$
NGC 5986	$0.074 \pm 0.006$	NGC 6779	$0.192 \pm 0.083$
NGC 6093	$0.107 \pm 0.010$	NGC 6809	$0.115 \pm 0.011$
NGC 6101	$0.079 \pm 0.016$	NGC 6838	$0.067 \pm 0.010$
NGC 6121	$0.039 \pm 0.031$	NGC 6934	$0.123 \pm 0.028$
NGC 6144	$0.161 \pm 0.031$	NGC 6981	$0.151 \pm 0.028$
NGC 6171	$0.066 \pm 0.011$	NGC 7078	$0.204 \pm 0.014$
NGC 6205	$0.113 \pm 0.024$	NGC 7089	$0.156 \pm 0.023$
NGC 6218	$0.060 \pm 0.015$	NGC 7099	$0.000 \pm 0.020$
NGC 6254	$0.173 \pm 0.022$		

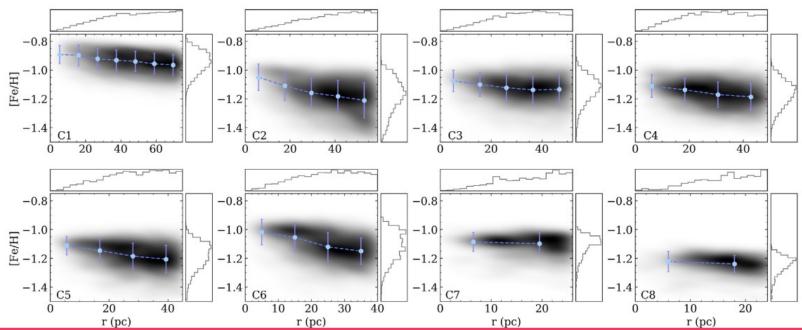
Legnardi+22

## **METALLICITY VARIATIONS IN GALACTIC GCs**

**MODEST 24** 



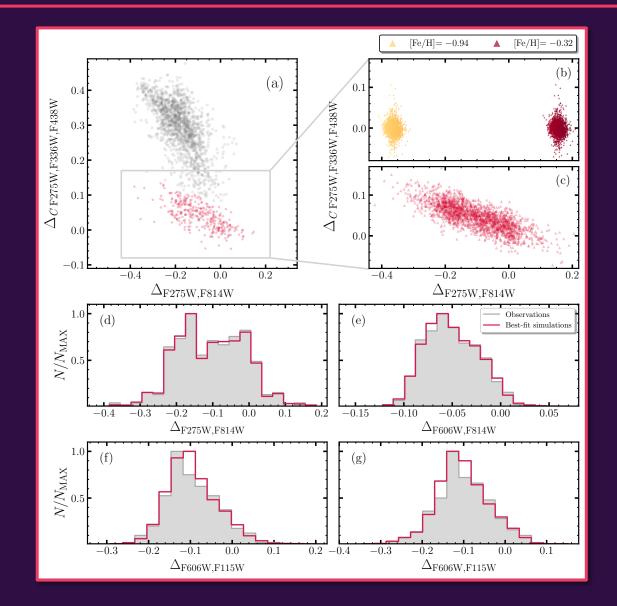




McKenzie+21

· **B2** 

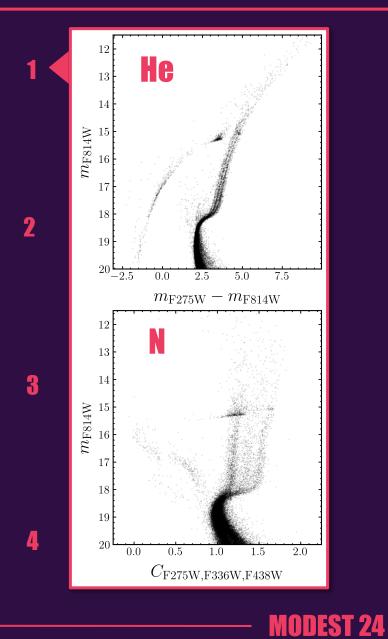
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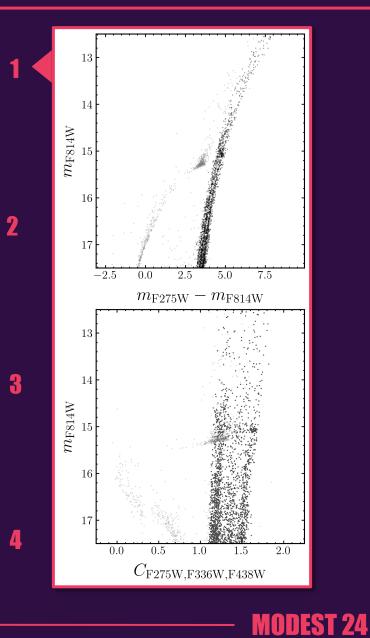
Legnardi+24

**B3** 

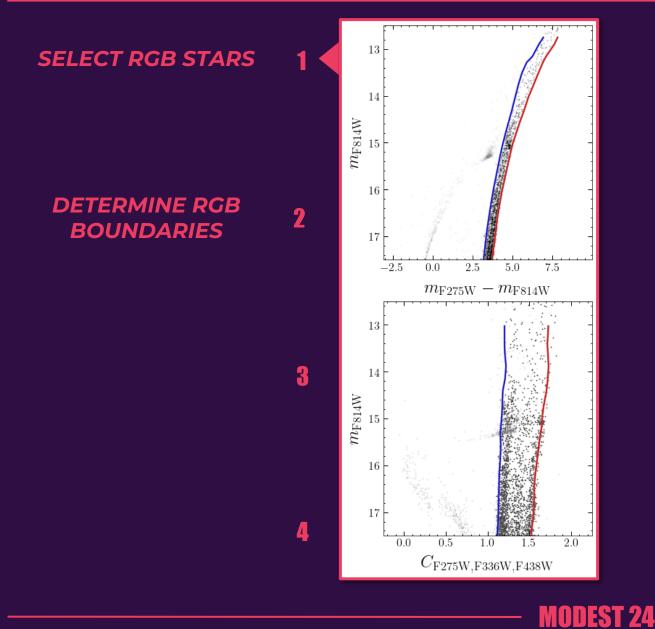


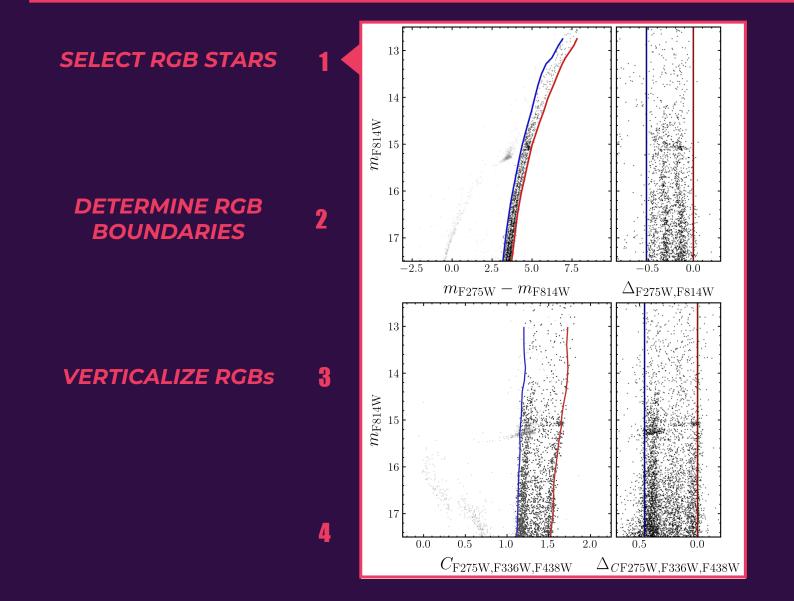


SELECT RGB STARS

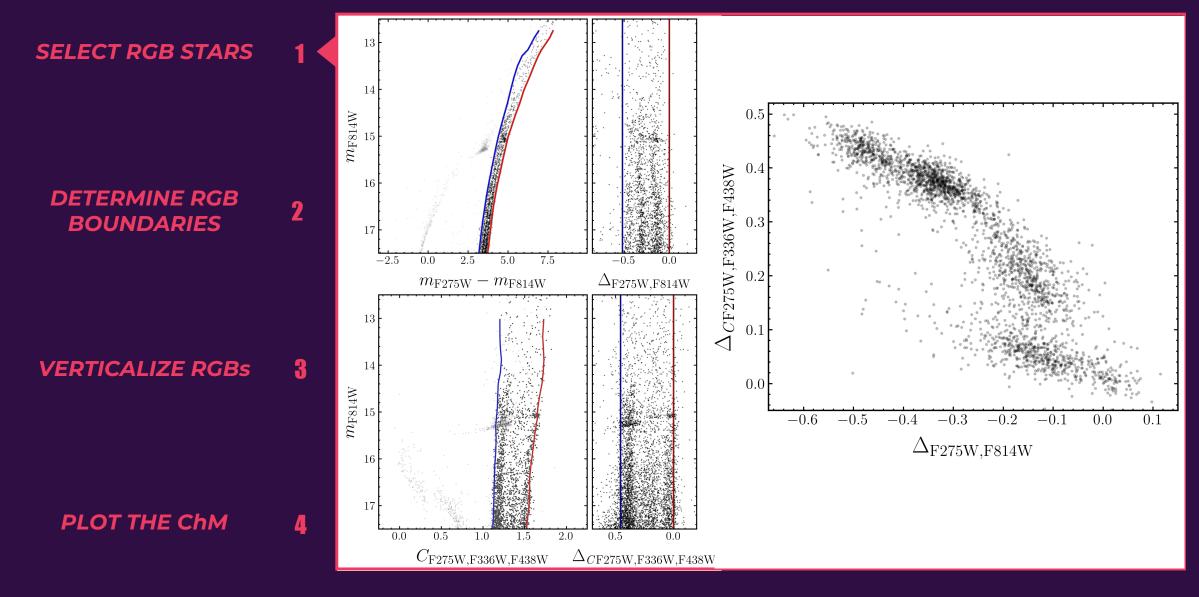


2.0









**MODEST 24**