Group 2 & Group 3 activities

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20-12-2023



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Group 3 Electronics and Data Acquisition and Processing -Members



Mariusz Suchenek (leader)







Marcin Ziembicki (postDoc) Marek Cieślar (postDoc)





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Edit Fenyvesi (postDoc)

Group 2 & Group 3 - Activities

Measurement campaigns:

- Seismic campaign in Sardinia (ET, 2022)
- Infrasound campaign in Sardinia (ET 2022/2023)
- VIRGO seismic and infrasound sensors installation and maintenance (VIRGO)
- Maastricht ETpathfinder seismic and infrasound preparation (ET installation in 2024)
- Seismic and infrasound measurements in LNGS, Laboratori Nazionali del Gran Sasso (ET 2023)

Data analysis and modeling:

- Detection gravitational wave signals from binary black with deep learning - W. Alhassan
- Modeling quartz resonators M. Pietrzak
- Data aquisition from a photomultipliers M. Ziembicki







VIRGO seismic and infrasound sensors

- We are responsible for over 120 seismic and infrasound sensors installed in VIRGO buildings (North, West and Central)
- Sensors create a network which mitigates Newtonian Noise and increases sensitivity of gravitational wave detector
- We use a modified IT infrastructure, modified with sensor synchronization and power supply





The infrasound and seismic installation

- Characterization of the infrasound and seismic field underground, a place similar to the destination site for the telescope and other experiments
- Seismic and infrasound noise are the components of the Newtonian noise
- 22 days of data, 06.03-28.03 2023 ۲
- Data stored at intra.astrocent.camk.edu.pl available for all ۲
- Sensors placed in Hall 1 with existing experiments ۲



The infrasound and seismic installation

 sensors placed on a concrete foundation prepared for the darkwave experiment

Seismic spectra

Seismic spectra - 22 days

Seismic spectra - 10.03.2023 (24 hours)

• for 2 hours (8.00-10.00) significantly higher

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Typical microquake

• RAW data, velocity

Typical data (high pass filtered)

Quake frequency

- Expect a strongest quake with peak maximum velocity over 1 s of 1 mm/s, in 10 years
- 491 hours / 24 = 20,45 days

Sound spectra - averaged 22 days

Acoustic response of the hall

• hall dimensions, approx 100 m x 20 m x 25 m - acoustic resonator

v ~ 350 m/s,

Estimation:

f = 10 Hz λ = 35 m (λ /2 = 17.5 m)

f = 3.5 Hz λ = 100 m (λ /2 = 50 m), f = 1.75 Hz λ = 200 m (λ /2 = 100 m)

Sound spectra - 10.03.2023 (24 hours)

Newtonian noise based on LNGS measurements

- Acoustic NN is not negligible
- Assume a geometry of detector: hall 100 m x 20 m x 25 m
- Noise level in tunnels is too high
- Lower limit from ambient noise can also affect ET sensitivity

Conclusions

 Infrasound and seismic analysis (Newtonian noise = seismic + infrasound noise)

Seismic:

- Very low seismic noise
- We estimate that within 10 years the quake will be a maximum of 1 mm/s

Infrasound:

- Significiant infrasound noise
- A large hall where waves with a frequency of around 2 to 20 Hz freely propagate into the hall

The infrasound installation in Sardinia

- Characterization of the infrasound field in the mine, a place similar to the destination site for the telescope
- Multipoint infrasound measurements
- Infrasound are one of the components of the Newtonian noise
- More than 9 months of data (since 21st November 2022)
- Data stored at etrepo.df.unipi.it and info. also on wiki.et-gw.eu

Two types of infrasound sensors (infrasound microphones):

- GRAS 47AC 1/2"
- Astrocent microphones

GRAS 47AC 1/2" CCP Infra-Sound Microphone Set

Condenser microphone set for infra-sound measurements in open acoustic fields

- frequency range: **0.09 Hz** to 10 kHz
- dynamic range: 20 dB(A) to 148 dB
- sensitivity: 8 mV/Pa

https://www.grasacoustics.com/products/special-microphone/infra-sound-microphones/product/712-47ac

Infrasound microphones - Astrocent

- developed by the Astrocent team
- low-cost infrasound microphone ~100 EU
- frequency range from 0.1 Hz to 120 Hz (AC47 Gras from 0.09 Hz to 20 kHz)
- sensitivity: 2.5 mV/Pa
- distortion max. 1.5 dB

Improved Infrasound microphones - Astrocent

- developed by Astrocent team
- low-cost infrasound microphone ~100 EU
- frequency range from 0.04 Hz to 120 Hz (AC47 Gras from 0.09Hz to 20 kHz)
- sensitivity: 46 mV/Pa
- distortion max. 1 dB

Sos Enattos mine The mine consists of tunnels with caverns

Caverns:

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ST0 - Surface Station

- power supply from solar panels
- 1x GRAS 47AC 1/2", 1x microphone developed by Astrocent
- data synchronized by GPS, 1 pps signal (1 Hz)

Stations

ST1, ST3 - in every station 3x infrasound microphones

data synchronized by NTP serwer via

Internet

- 1x Gras + 1x Astrocent microphone
 (inside the cavern)
- 1x Astrocent (outside the cavern)

Sardinia - Sos Enattos Mine

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Sardinia - Sos Enattos Mine SOE3

Inside the cavity

1x Gras - red

1x Astrocent - blue

Outside the cavity 1x Astrocent - green

10th December

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Coherence ST3

Inside cavity microphones

Gras-Astrocent

Inside and outside cavity

Astrocent-Astrocent

Coherence ST1 (inside) & ST3 (inside)

ST1 & ST3 Gras

Coherence STO (surface) & ST1/3 (inside) ST0 & ST1 Gras

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Conclusions

- Sosenattos 6x infrasound microphones inside the mine, 2 outside (installation since 21st November 2022)
- distributed infrasound measurements
- ST1, ST3 3x infrasound microphones 1x Gras + 1x Astrocent microphone (inside the cavern) + 1x Astrocent (outside the cavern)
- damping of the wall between the tunnel and the cavern about factor 10x (amplitude, 100x PSD)
- very high coherence between microphones below 0.6 Hz

Thank you for your attention Mariusz Suchenek msuchenek@camk.edu.pl

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Power spectrum & coherence

Power spectrum

$$S_{xx,f} = \frac{2dt^2}{T} X_f X_f^*$$
$$S_{xy,f} = \frac{2dt^2}{T} X_f Y_f^*$$

X, Y - spectrum of signal

T - total time of recording

dt - samplig time

Coherence

$$K_{xy,f} = \frac{||}{\sqrt{< S_{yy,f}>}}$$