Search for Dark Matter with Liquid Argon Detector

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ASTROCENT





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Annual Modulation and DarkSide-50

Results from other experiments

NaI(TI) detector

- DAMA/LIBRA: modulation with proper features at 13.7σCL
- ANAIS112: reject DAMA with $\approx 3\sigma$
- COSINE 100: consistent with both DAMA and the no-modulation case

LXe detector

- XENON100: reject DAMA's modulation in 2-6keV bin
- LUX: 9.2σ tension with the DAMA/LIBRA result
- XMASS: excludes the DAMA/LIBRA allowed region at $\approx 3\sigma$

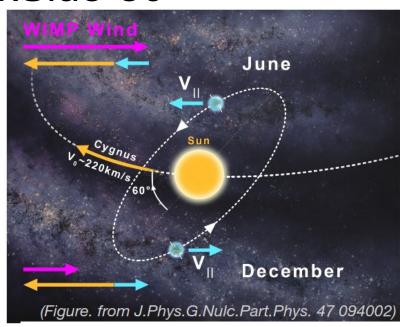
LAr detector

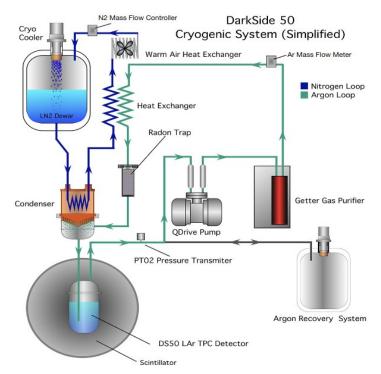
This talk (arXiv:2307.07249)
Search for dark matter annual modulation with DarkSide-50

DarkSide-50 TPC working principle:

Light collected by top and bottom Photomultiplier Tubes (PMT)

- S1 (primary scintillation), is produced in LAr due to excitation and recombination after ionization
- S2 (secondary scintillation) produced in the gas phase by drifted electrons







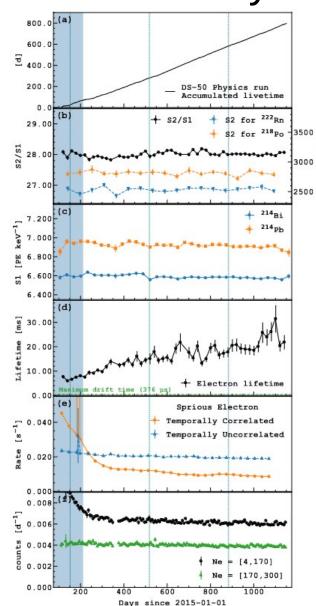
DarkSide-50 Long Term Stability

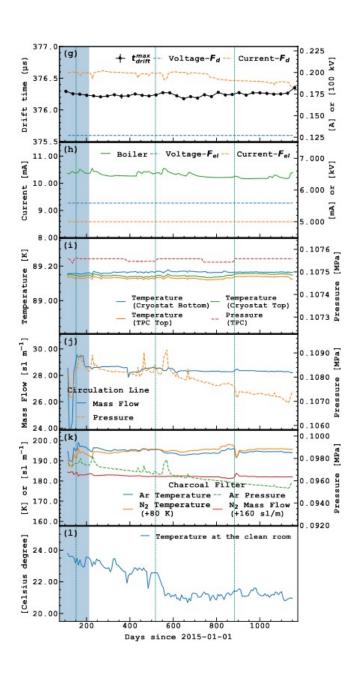
Stability of all 72 slow control parameters was checked:

- Quantitatively
- Lomb-Scargle periodogram
- Correlation with Data

There is <u>no significantly high</u> <u>coefficient</u> and <u>strong periodicity</u> found_for each of energy ranges

The mean number of photoelectrons per ionization electron (g2 [PE/e-]) the drift field (F_d [V/cm]) are stable within 0.5% and 0.01%







Search for Annual Modulation

- Four radioactive isotopes decaying in 3 years are taken into account
- Dark Matter Event rate as a function of time is modeled with a cosine signal:

$$f(t) = \underbrace{A_{\chi} \cos\left(\frac{t - \phi}{\underline{T}/2\pi}\right)}_{\text{Signal Fixed to 1y}} + \underbrace{\sum_{l} \frac{A_{l}}{\tau_{l}} e^{-t/\tau_{l}}}_{\text{Including long-lived isotope}} + \underbrace{C_{,}}_{\text{Including long-lived isotope}}$$

Likelihood fit to the model with 7-d time bin:

$$\mathcal{L} = \prod_{i \in t_{\text{bins}}} \mathcal{P}\left(n_i \mid m_i(A_{\chi}, \phi, C, \Theta)\right) \times \prod_{\theta_k \in \Theta} \mathcal{G}(\theta_k \mid \theta_k^0, \Delta \theta_k).$$

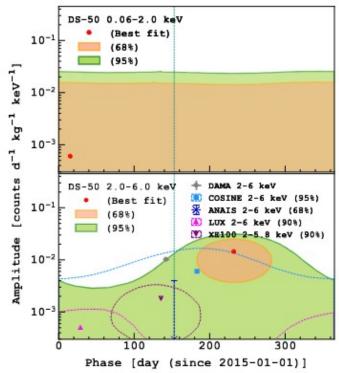
The 1D fit (uppers plots) are **consistent to the background-only model**

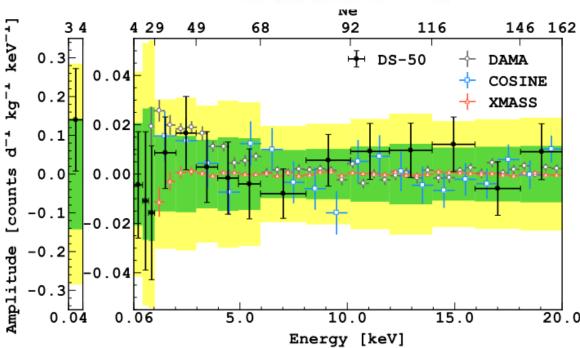
- Neither confirm nor reject the DAMA's observation

The 2D simultaneous fit (bottom plot) in both time and energy bins uses:

$$\mathcal{L} = \prod_{i \in t_{\text{bins}}} \prod_{j \in E_{\text{bins}}} \mathcal{P}\left(n_i^j \mid m_i^j(A_\chi^j, C^j, \tilde{\Theta})\right) \times \prod_{\tilde{\theta}_k \in \tilde{\Theta}} \mathcal{G}(\tilde{\theta}_k \mid \tilde{\theta}_k^0, \Delta \tilde{\theta}_k),$$

- Fixed the phase ϕ (June 2nd) and period T (1-yr)
- Amplitudes of the short-decayed component for each energy bin are correlated







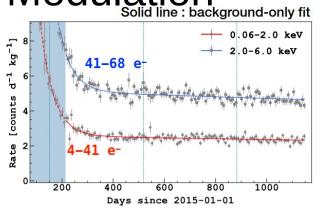
Search for Annual Modulation

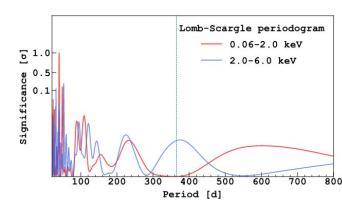
Lomb-Scargle algorithm is applied to look for any periodic signal

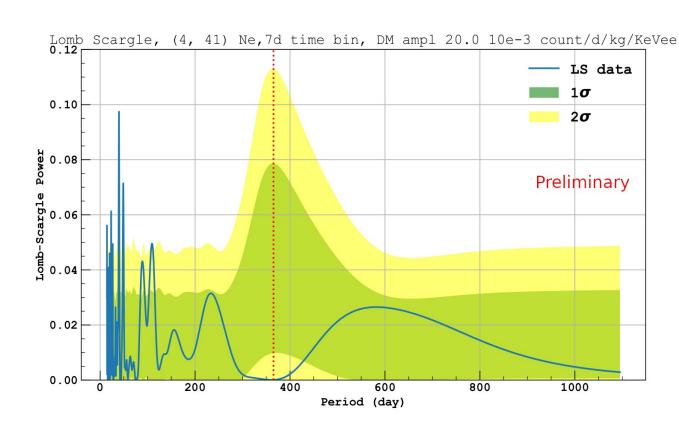
- Residuals of the backgroundonly fit are converted into the frequency space

No significant signal is observed

- Bottom plot shows a Lomb-Scargle periodogram, with Brazilian band corresponding to toy-MC datasets, showing that a median of 1 σ significance for the false alarm probability is obtained with the addition of 0.03 counts/(d kg keV)







Other activities

- DEAP Collaboration meeting, Mexico City
- DarkSide Collaboration meeting, June, LNGS, Italy
- XVIII International Conference on Topics in Astoparticle and Underground Physics (TAUP2023), Vienna Proceeding under review
- Submitting the thesis manuscript

Thank you for your attention.