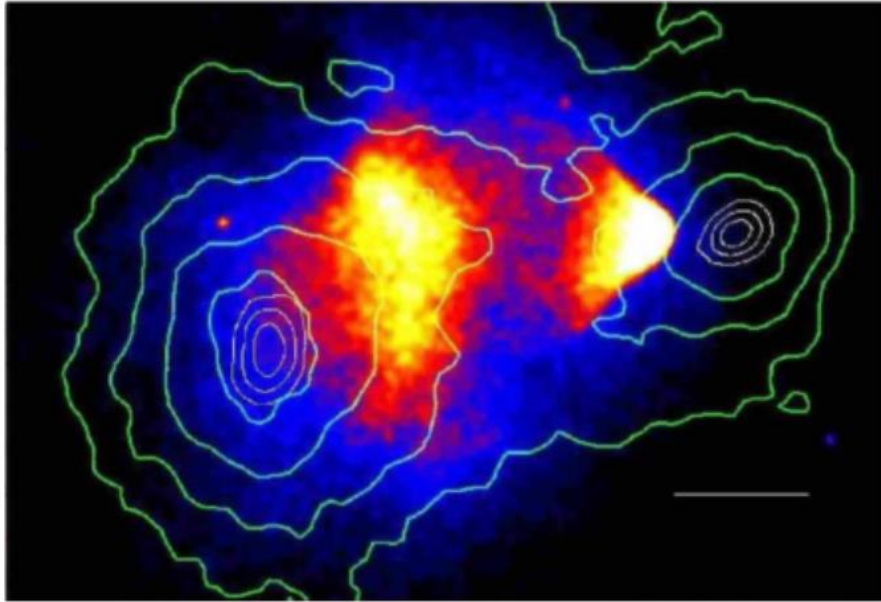


Annual report 2023: Ewa L. Łokas

Published refereed papers:

- Łokas, E. L. "Merging galaxy clusters in IllustrisTNG" 2023, A&A, 673, A131
- Łokas, E. L. "Preprocessing in small groups: Three simulated galaxies interacting prior to cluster infall" 2023, A&A, 678, A147
- Martínez-García, A. M., del Pino, A., Łokas, E. L., van der Marel, R. P., Aparicio, A. "Internal kinematics of dwarf satellites of MW/M31-like galaxies in TNG50" 2023, MNRAS, 526, 3589

Bullet cluster



1E 0657-56 - first clear example of a merging galaxy cluster showing a bow shock associated with a subcluster that recently passed through a bigger cluster

- It was claimed to provide good evidence for the presence of dark matter in galaxy clusters
- The distribution of the total mass inferred from gravitational lensing was found to coincide with the one of galaxies but not with the gas
- The inferred velocity of the bullet ~ 4500 km/s was difficult to reconcile with LCDM theory

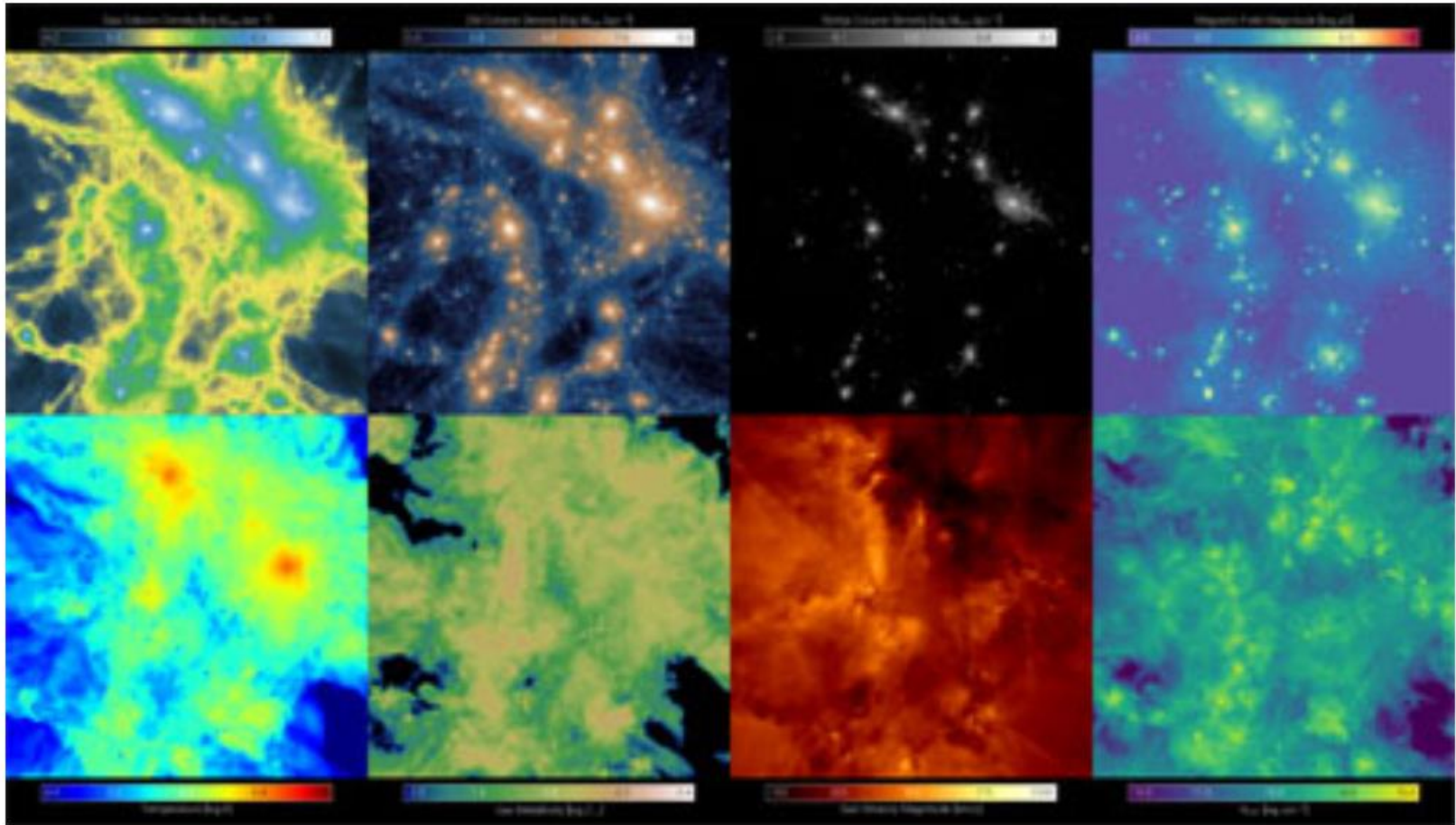
IllustrisTNG simulations

gas density

dark matter density

stellar mass

magnetic field strength



gas temperature

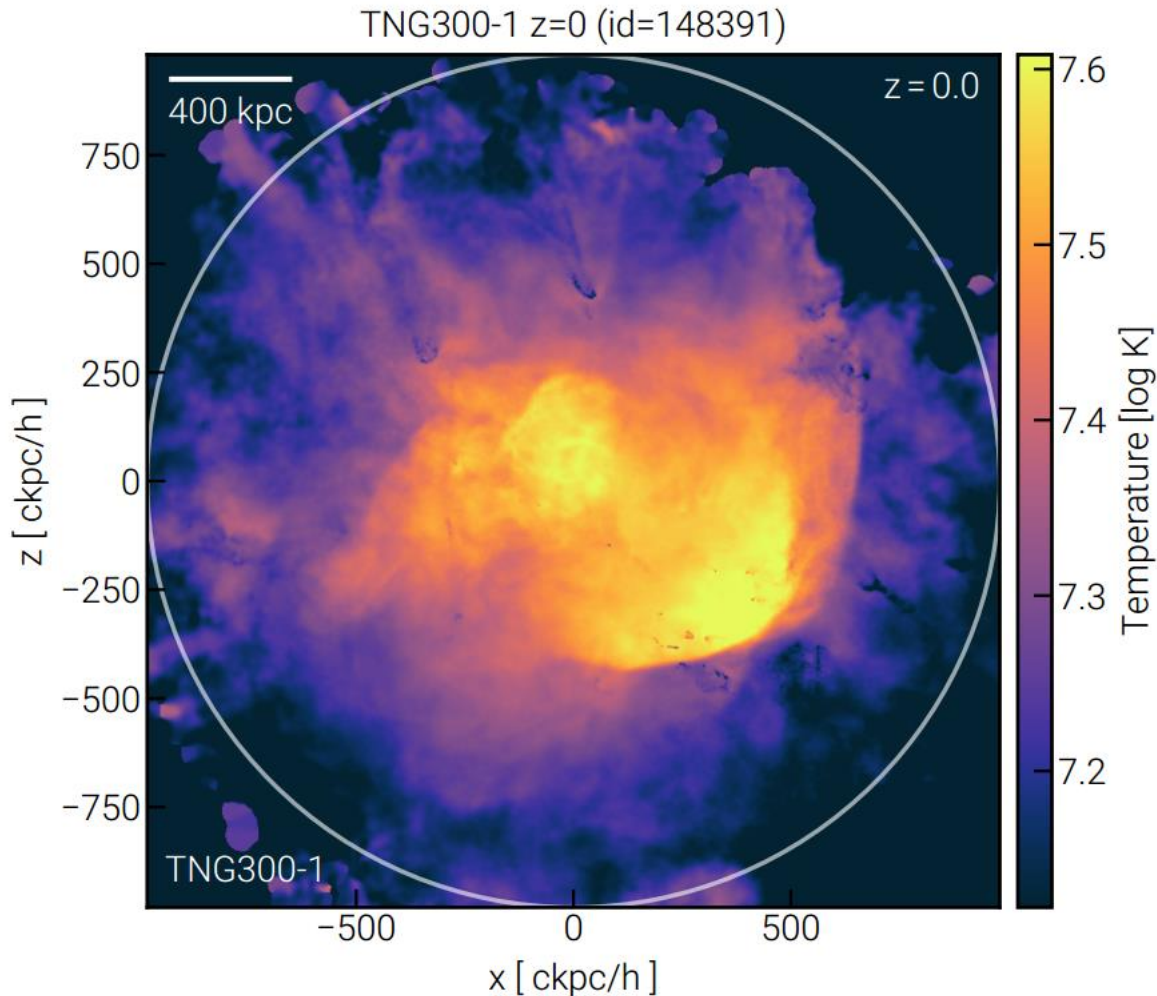
gas metallicity

gas velocity

column density of OVI

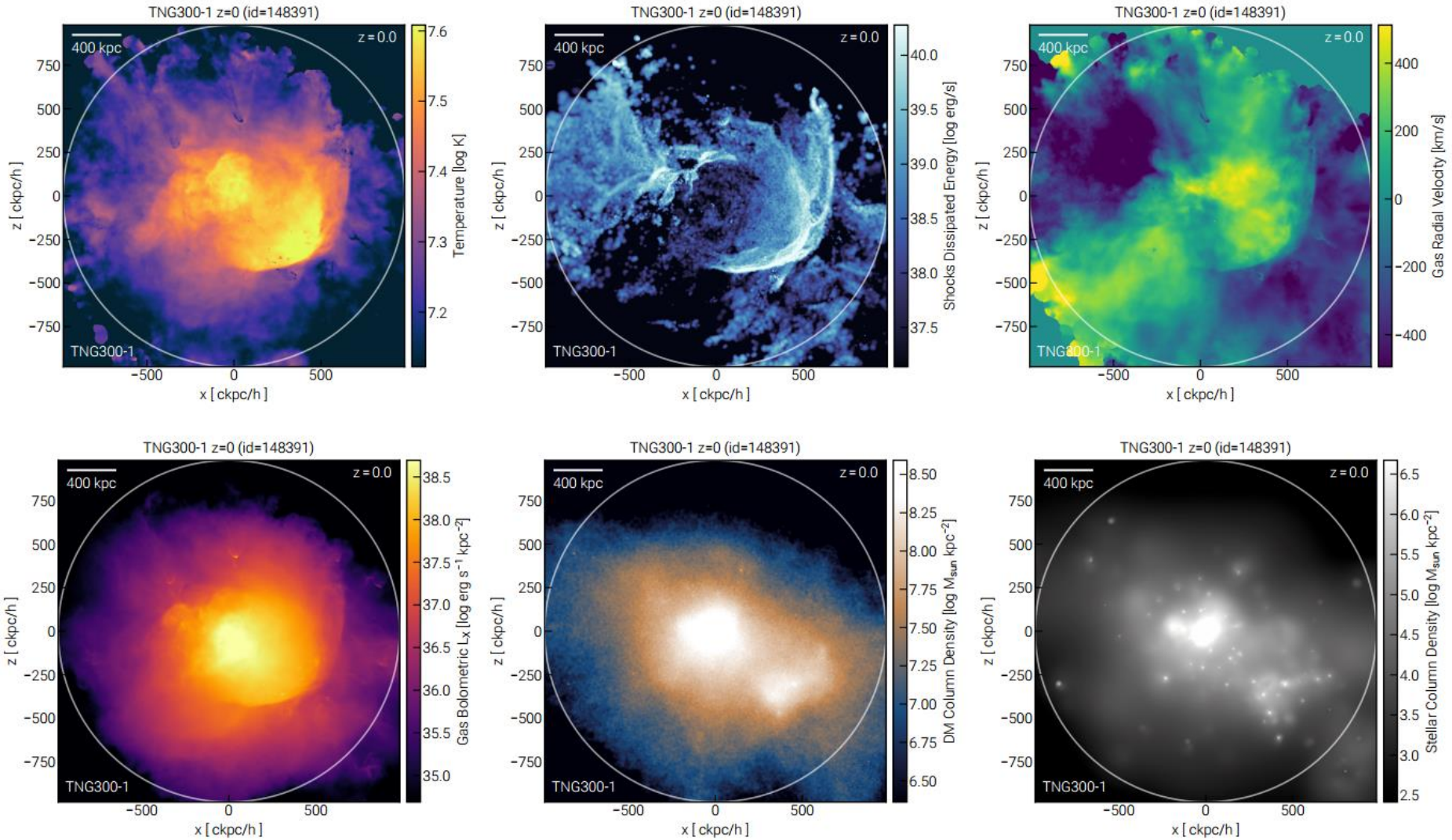
10 Mpc region from IllustrisTNG-100

Example of merging clusters



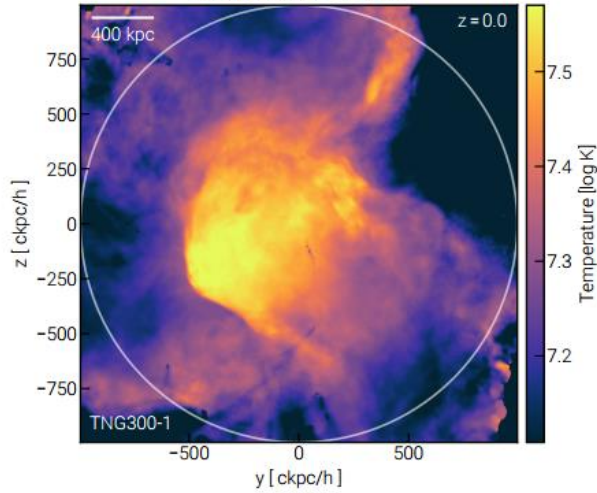
The smaller cluster (of maximum mass $3.1 \times 10^{13} M_{\text{sun}}$) passed close to the center of the bigger one (of mass $3.2 \times 10^{14} M_{\text{sun}}$) 0.5 Gyr ago with a pericenter of about 190 kpc and with velocity of 1670 km/s

Properties of merging clusters

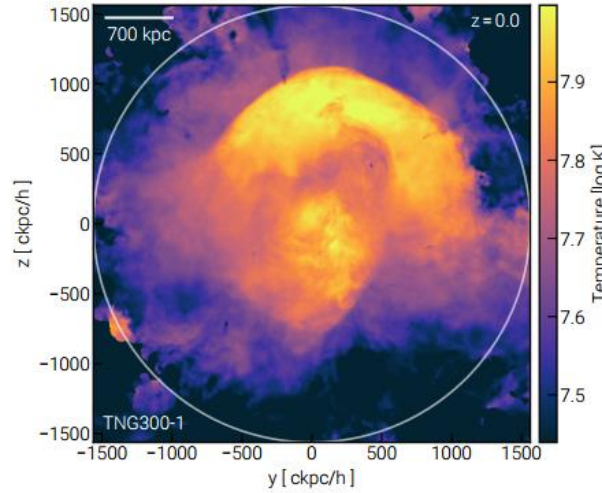


Other examples of merging clusters

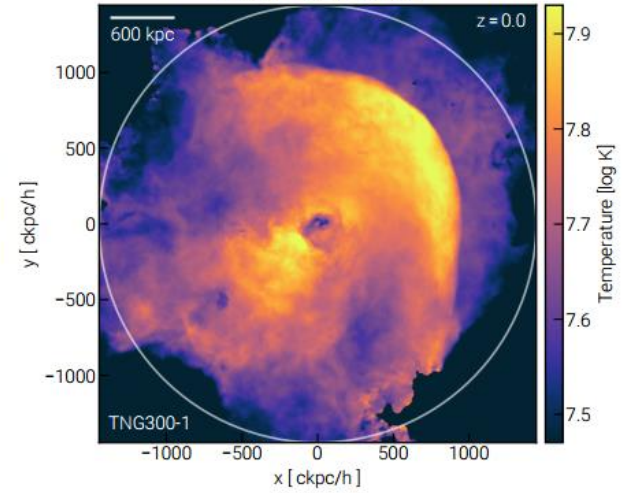
TNG300-1 z=0 (id=150265)



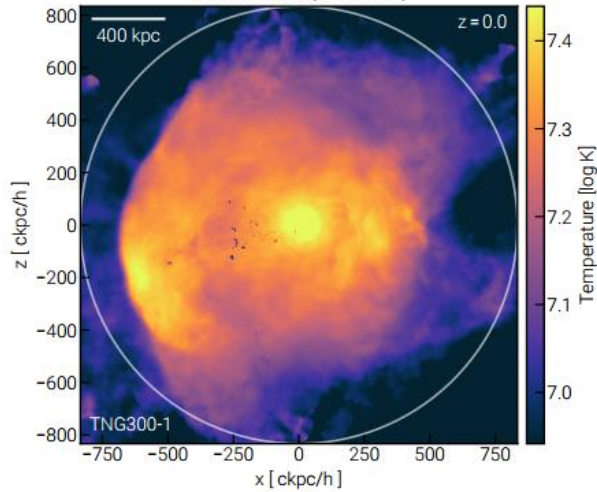
TNG300-1 z=0 (id=11748)



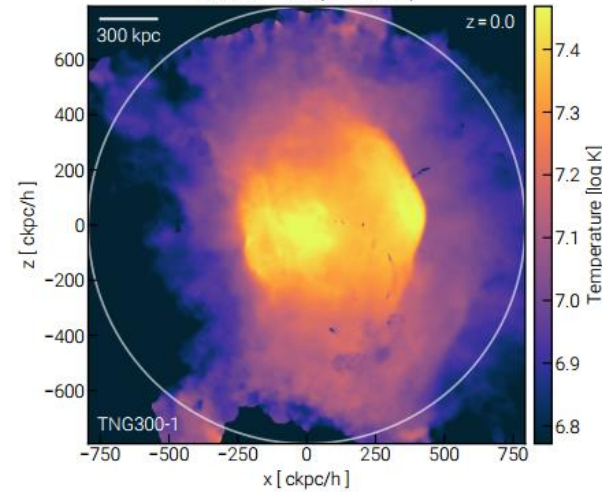
TNG300-1 z=0 (id=17908)



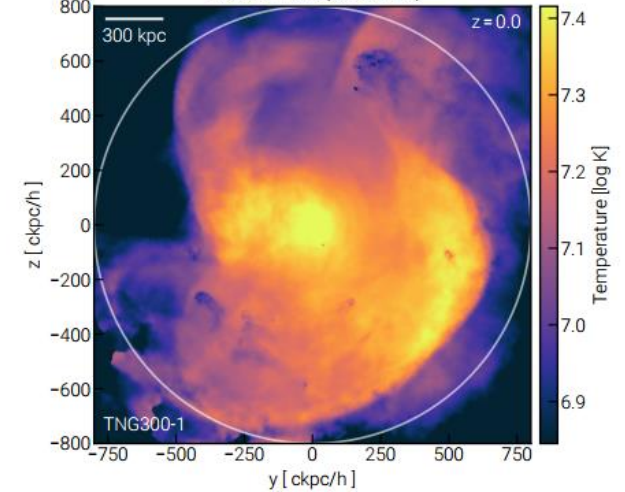
TNG300-1 z=0 (id=265569)



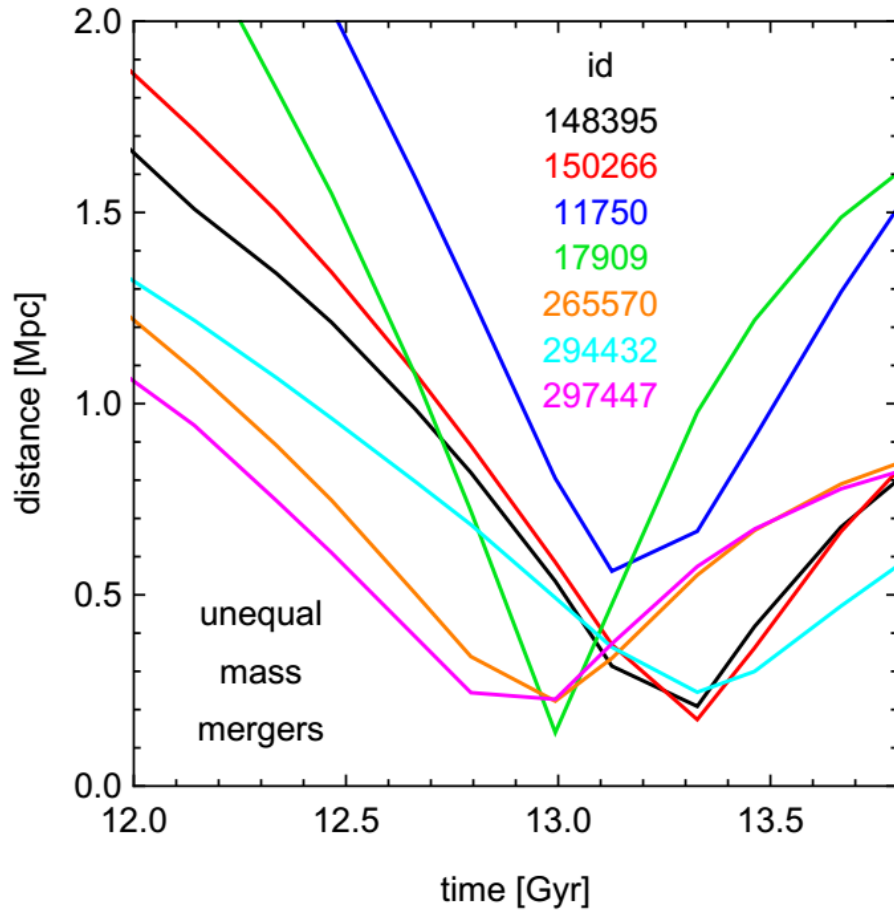
TNG300-1 z=0 (id=294431)



TNG300-1 z=0 (id=297446)



Orbits of subclusters



- The companion cluster passed close to the main one only once, between 0.9 and 0.3 Gyr ago, with the pericenter distance of 100-530 kpc and a velocity of up to 3400 km/s
- The subcluster is typically an order of magnitude smaller in mass than the main cluster before the interaction

Invited talk on tidally induced bars



- Meeting organized to celebrate Lia Athanassoula's career and her pioneering contributions to the field of bar dynamics
- Presentation files are available here:
<https://zenodo.org/communities/bars2023/>