

Paweł Moskalik: 2023

Papers published in 2023

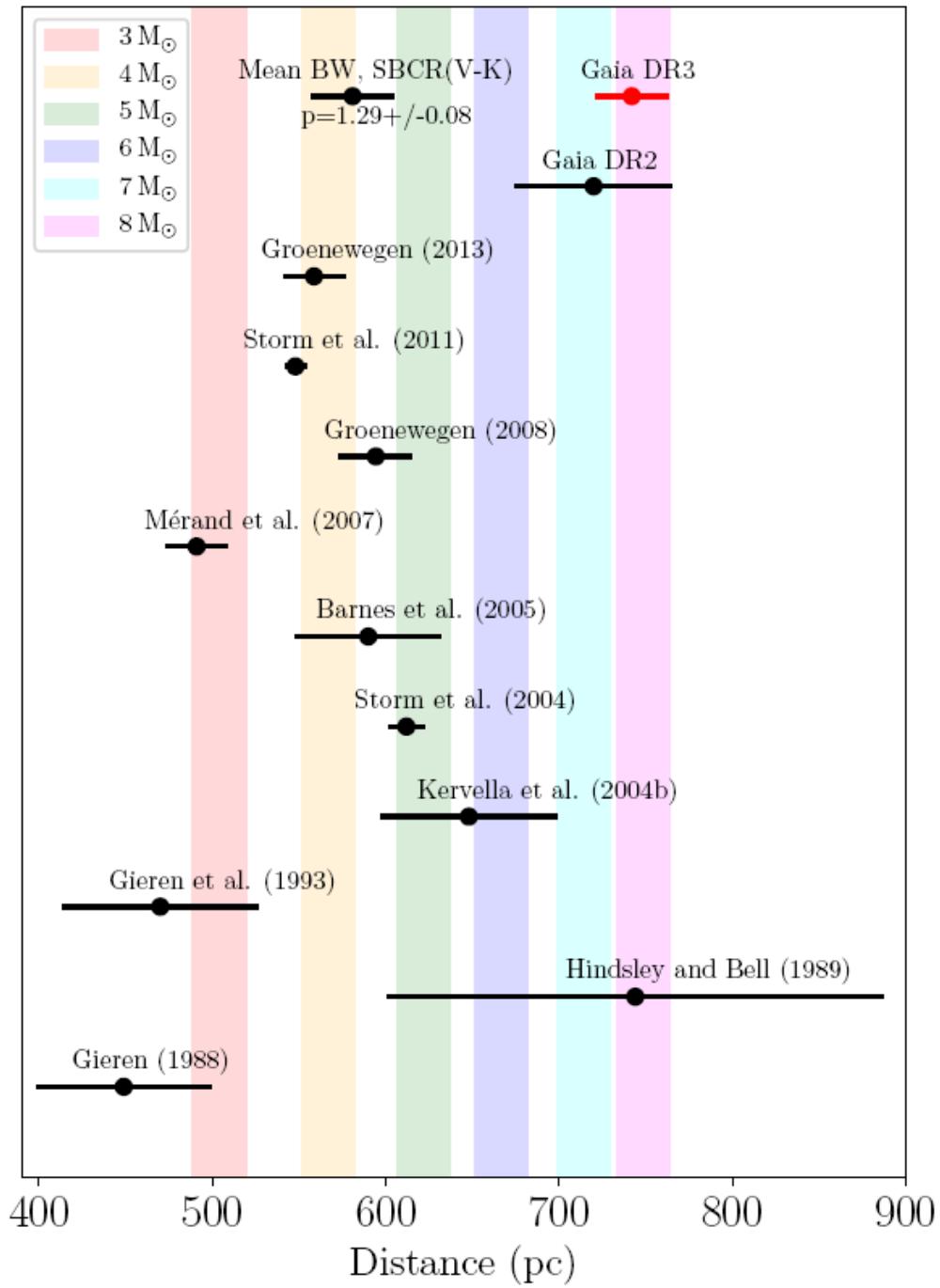
- „**Type II and Anomalous Cepheids in the *Kepler* K2 Mission**”, M. I. Jurkovic, E. Plachy, L. Molnár, M. A. T. Groenewegen, A. Bódi, P. Moskalik & R. Szabó, 2023, MNRAS, 518, 642.
- „**Metallicity Estimation of MW, SMC and LMC Classical Cepheids from the shape of the V- and I-band Light Curves**”, V. Hocdé, R. Smolec, P. Moskalik, O. Ziolkowska & R. Singh Rathour, 2023, A&A, 671, A157.

Paweł Moskalik: 2023

Papers submitted in 2023

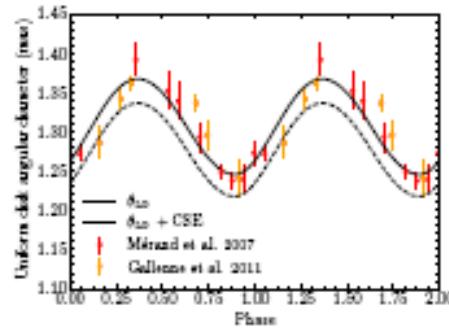
- „**Pulsation Modelling of the Cepheid Y Ophiuchi with RSP/MESA. Impact of Circumstellar Envelope and a High Projection Factor on Baade-Wesselink Method**”, V. Hocdé, R. Smolec, P. Moskalik, R. Singh Rathour & O. Ziółkowska, 2023, A&A (accepted).
- „**Double-Mode RR Lyrae Stars Observed by K2: Analysis of High-Precision Kepler Photometry**”, J. M. Nemec, A. F. Linnell Nemec, P. Moskalik, L. Molnár, E. Plachy, R. Szabó & K. Kolenberg, MNRAS (submitted).
- „**Precise Fourier Parameters of Cepheid Radial Velocity Curves**”, V. Hocdé, P. Moskalik, N. A. Gorynya, R. Smolec, R. Singh Rathour & O. Ziółkowska 2023, A&A (submitted).

Modelling of the Cepheid Y Ophiuchi

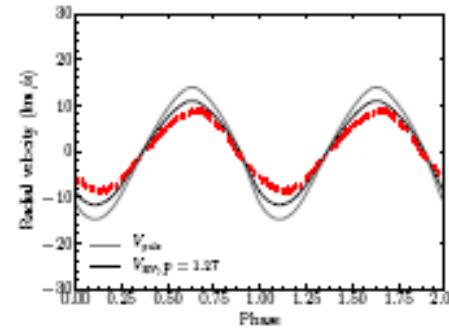


Gaia DR3 : $d = 742 \pm 21 \text{ pc}$

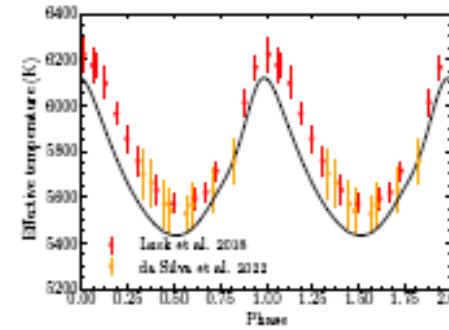
Mean BW : $d = 581 \pm 24 \text{ pc}$



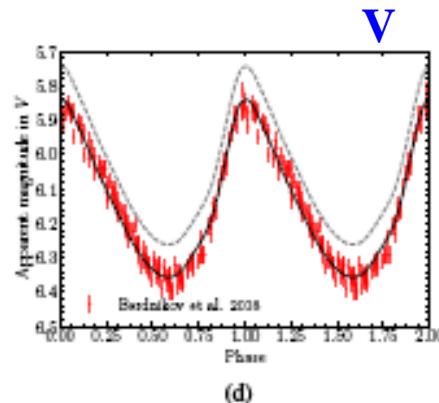
(a)



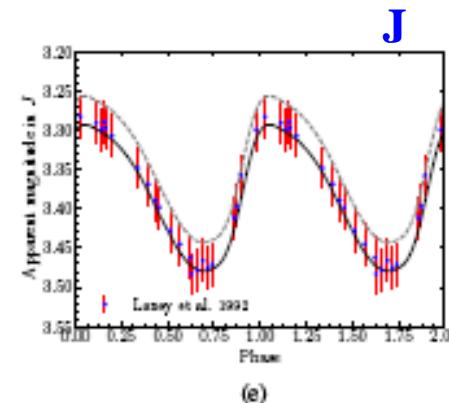
(b)



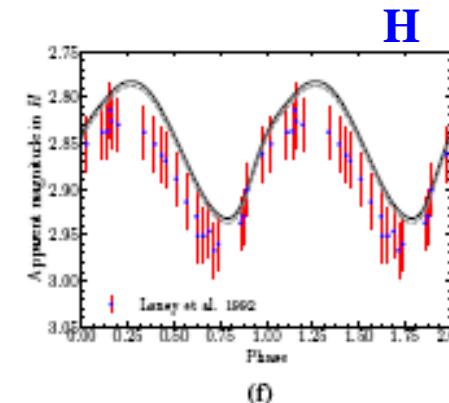
(c)



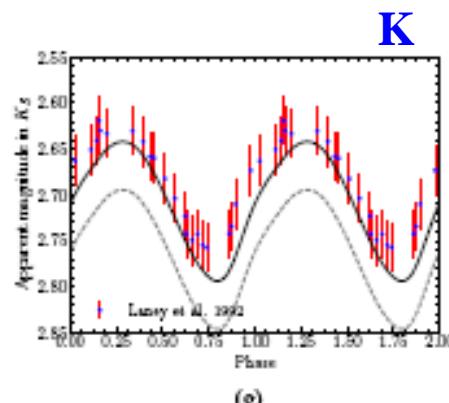
(d)



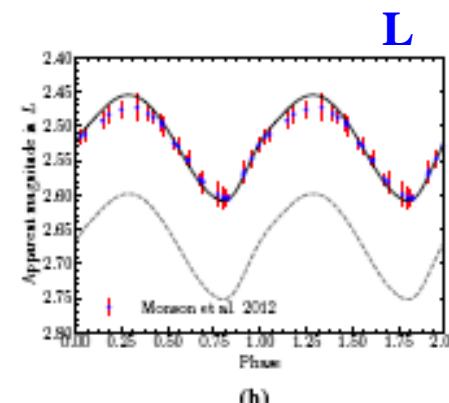
(e)



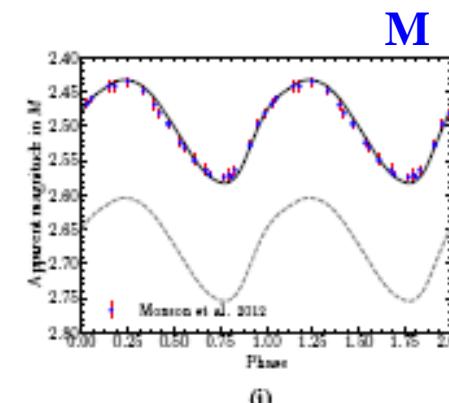
(f)



(g)



(h)



(i)

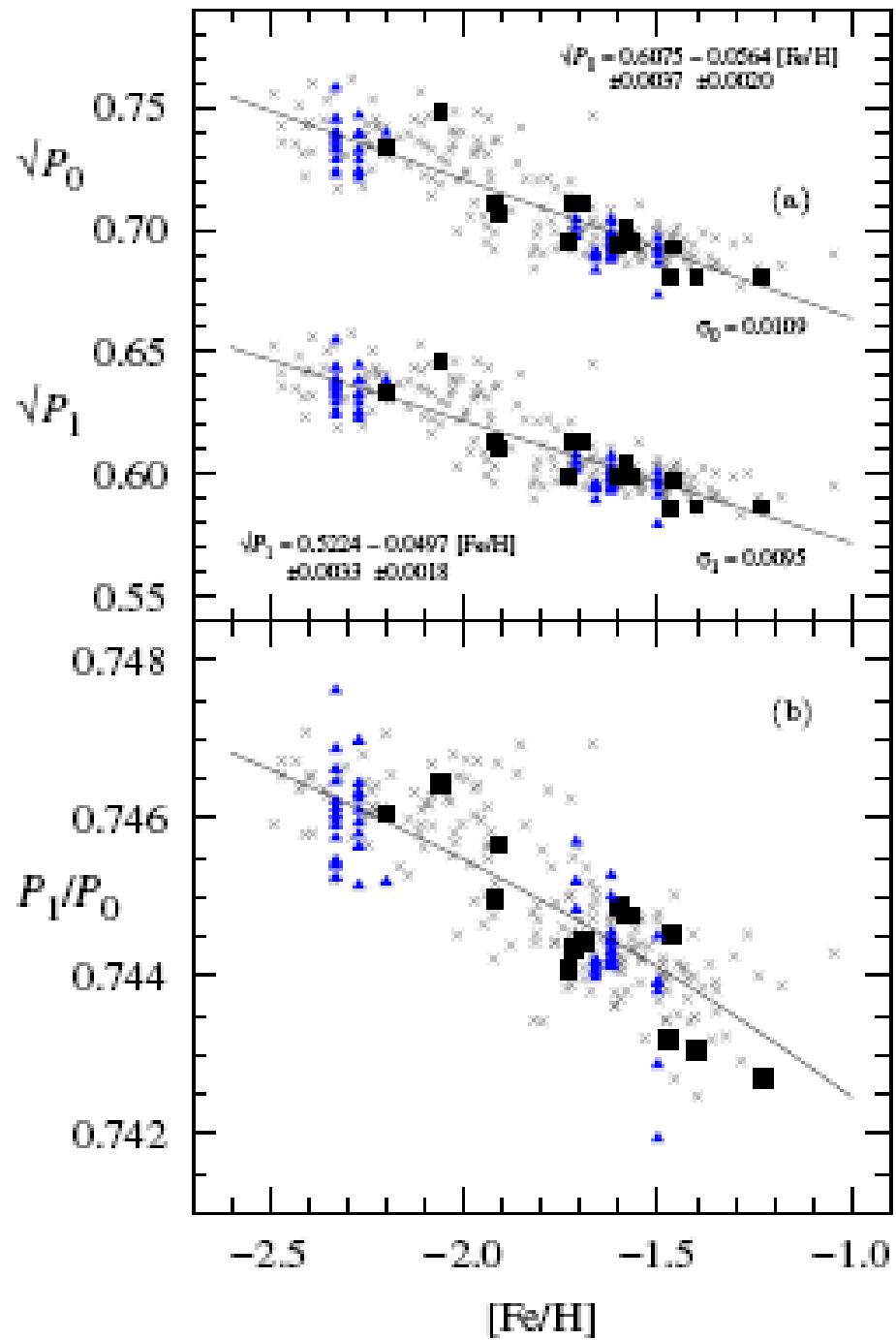
Best fit

$$\mathbf{M} = 8 \mathbf{M}_{\odot} \quad \mathbf{T}_{\text{eff}} = 5725 \text{ K}$$

$$\mathbf{d} = 748 \text{ pc}$$

Fig. 2: Best result of the non-linear analysis with RSP/MESA for $8 M_{\odot}$ model. Uniform disk angular diameter, radial velocity curve and effective temperature are displayed in (a), (b) and (c) respectively. The photometric panels indicate the apparent magnitudes in (d) V-band, (e) J-band (f) H-band (g) K_s-band (h) L-band (i) M-band. In the angular diameter and photometric panels, thick black line and dashed grey lines are RSP models with and without CSE models respectively.

Double-Mode RR Lyrae Stars Observed by K2



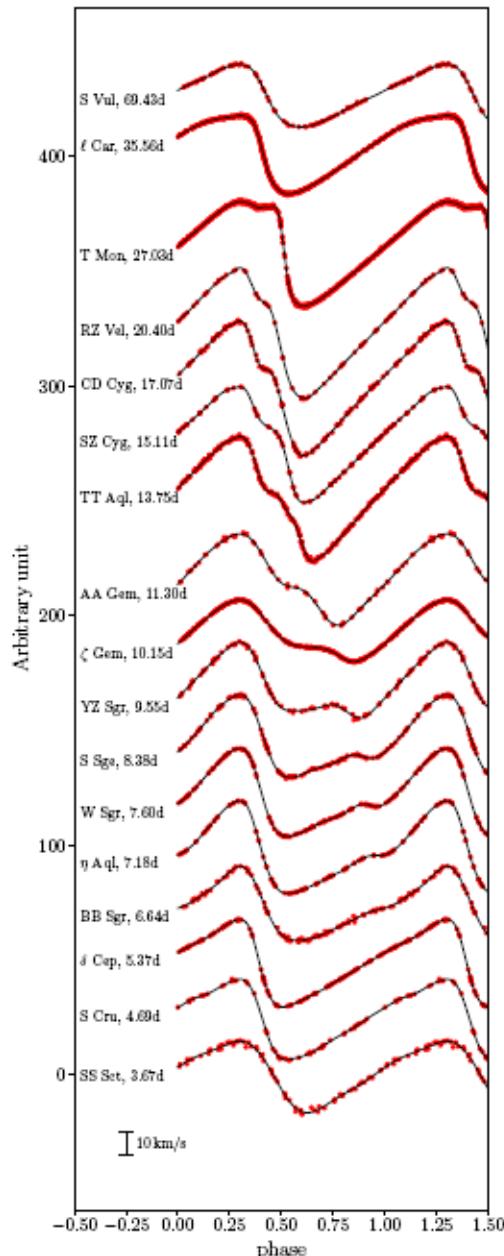
$$[\text{Fe}/\text{H}] = 7.59 - 13.25 \sqrt{P_0},$$

$$[\text{Fe}/\text{H}] = 7.42 - 15.08 \sqrt{P_1},$$

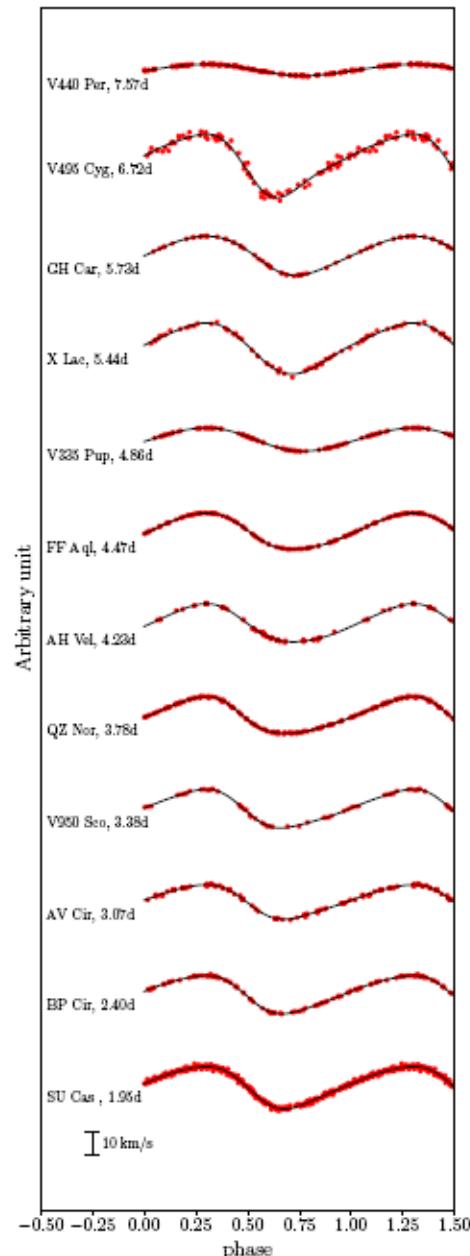
$\sigma = 0.17 \text{ dex}$

$\sigma = 0.17 \text{ dex}$

Precise Fourier Parameters of Cepheid Radial Velocity Curves



(a)



(b)

Fig. 1: Radial velocity curves and Fourier fits for a set of fundamental-mode (a) and first-overtone Cepheids (b). The scale is indicated in the left bottom corner of the plots.

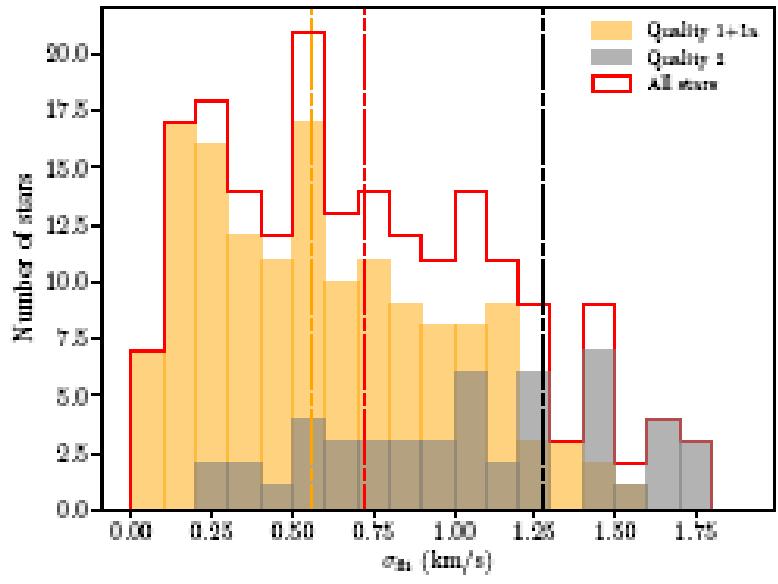
218 single-mode Cepheids

- **178 FU**
- **33 FO**

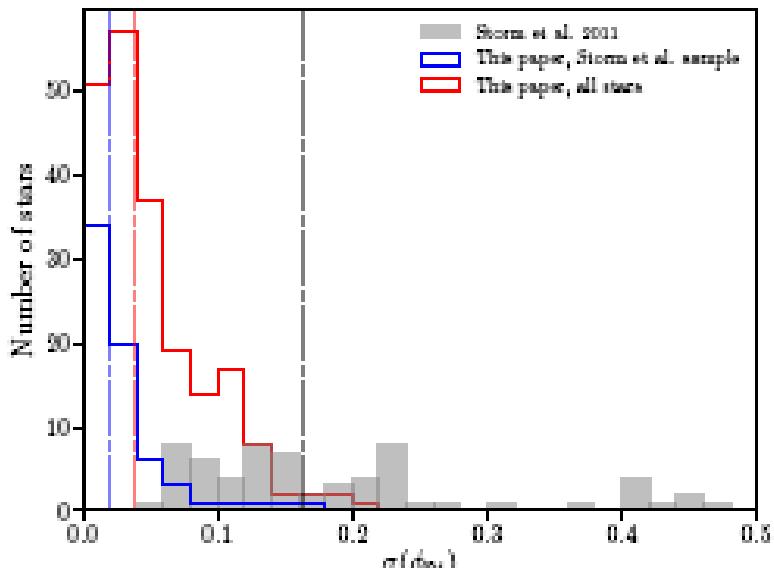
$$\mathbf{V}_r(t) = \mathbf{A}_0 + \sum \mathbf{A}_k \sin(k\omega t + \phi_k)$$

$$\mathbf{R}_{k1} = \mathbf{A}_k / \mathbf{A}_1$$

$$\Phi_{k1} = \phi_k - k\phi_1$$



(a)



(c)

Median σ :

- **Quality 1 (147 stars) : 0.56 km/s**
- **Quality 2 (71 stars) : 1.05 km/s**
- **All (218 stars) : 0.72 km/s**

comparison with Storm et al. (2011)

Number of stars : 76 → 218

Median $\sigma(\phi_{21})$: 0.16 → 0.04 / 0.02

