Annual Report 2023 M. R. Schmidt

CAMK PAN, Warszawa, 2024



List of publications

Lithium in red novae and their remnants Kamiński, T. Schmidt, M. Hajduk, M., Kiljan, A., Izviekova I., Frankowski A. 2023, A&A 672, A196

ExoMol line lists - LIV. Empirical line lists for AIH and AID and experimental emission spectroscopy of AID in A^{1} (v = 0, 1, 2),

Yurchenko, S. N.; Szajna, W., Hakalla, R., Semenov, M., Sokolov, A., Tennyson, J., Gamache, R., Pavlenko, Y., Schmidt, M.R. 2024, MNRAS, 527, 9736

A radical transition in the post-main-sequence system U Equulei

Kamiński, T., Schmidt, M. R., Djupvik Anlaug, A., Menten, K. M., Kraus, A., Iłkiewicz, K., Steinmetz, T., Mobeen, Z., Szczerba, R. 2024, A&A

A bipolar structure and shocks surrounding the stellar-merger remnant V1309 Scorpii Steinmetz, T., Kamiński, T., Schmidt, M., Kiljan, A. 2024, A&A

Projects

ESPRESSO/VLT spectra of VY CMa

with T. Kamiński,

Empirical line list of ScO A 2Pi - X 2Sigma system with Y. Pavlenko

Update of line list of TiO B 3Pi - X 3Delta bands with Y. Pavlenko

MNRAS 479, 1401–1411 (2018) Advance Access publication 2018 June 11

ExoMol line lists XXVIII: the rovibronic spectrum of AlH

Sergei N. Yurchenko, Henry Williams, Paul C. Leyland, Lorenzo Lodi and Monthly Notices Jonathan Tennyson* of the

Department of Physics and Astronomy, University College London, London WROWAGRS, TRONOMICAL SOCIETY

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AIH lines in the blue spectrum of Proxima Centauri

Yakiv V. Pavlenko,^{1,2,3,4} Jonathan Tennyson^{,5} Sergei N. Yurchenko^{,5} Mirek R. Schmidt,² Hugh R. A. Jones,⁴ Yuri Lyubchik¹ and A. Suárez Mascareño¹⁰³

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ExoMol line lists – LIV. Empirical line lists for AlH and AlD and experimental emission spectroscopy of AlD in $A^1\Pi$ (v = 0, 1, 2)

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All in Proxima Centauri (M5.5 Ve Teff=2900 K)

ExoMol - "complete" list of lines of AIH (AloHa) extended with radiative dampings



AlH $A^{1}\Pi - X^{1}\Sigma^{+} v', v''=0,0$

HARPS spectrum in black, synthetic one with atomic lines in blue, synthetic AIH in red,



Extreme red supergiant VY CMa **ESPRESSO** spectra - emission component

























Molecular emission originates in the distant part of the circumstellar envelope in the gas of temperature of slightly above 100 K

There is no contradiction with earlier estimations of excitation temperature of 700 K and with large distant for emitting area from the photosphere



Emipirical list of lines of ScO A-X band and its verification in spectra of M-type stars collaboration with Y. Pavlenko (Main Astronomical Observatory, UAS Kiev)



Synthetic spectrum of HD189124 with Exomol TiO, YO and ScO



3000/0.0/0.0

Wavelengh (A)

Spectral analysis of laboratory spectrum of TiO B-X (system gamma')

Bands v=2-1, 3-2, 4-3 (already analysed), 5-4 (probably), 6-5 (extrapolated)



