

# CAMK Annual Meeting

**Marzena Śniegowska<sup>1,2</sup>**

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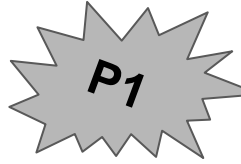
**with collaboration:**

Bożena Czerny<sup>3</sup>, Agnieszka Janiuk<sup>3</sup>, Mikołaj Grzędzielski<sup>3</sup>  
3. Center for Theoretical Physics PAS

**31.01-02.02.2024**

# This year I focused on flares in AGN:

## ★ Changing-Look (CL) AGN and accretion disk instabilities



- Seminars: *The challenges of the modelling Changing Look Active Galactic Nuclei phenomena*: Nicolaus Copernicus Astronomical Center in Warsaw, National Centre for Nuclear Research
- Contributed talk, The Time Domain in the era of Astronomical Big Data, Mitzpe Ramon, Israel
- Contributed talk, The Restless Nature of AGN: 10 years later, Napoli, Italy

### **Off-topic project:**

## ★ VLT spectropolarimetry of NLSy1 galaxies

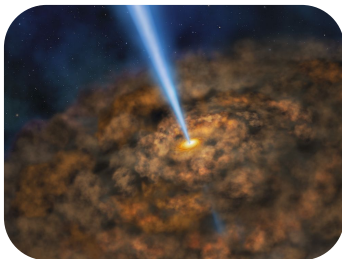
- Do Narrow-Line Seyfert 1 galaxies have smaller BH masses? Spectropolarimetric reverification, 10-14 July 2023, EAS Cracow



# Possible scenarios of CL AGN

Variation of obscuration  
(Extrinsic)

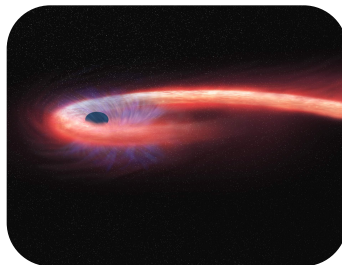
Inconsistent with broad line  
dimming and timescales



Credit: NASA/SOFIA/Lynette Cook

Tidal disruption event  
(Intrinsic)

The flares are longer-lived  
than the typical TDE



Credit: NASA / CXC / M. Weiss

Changes in accretion rate  
(Intrinsic)

Distinct physical processes  
from the rest of quasars

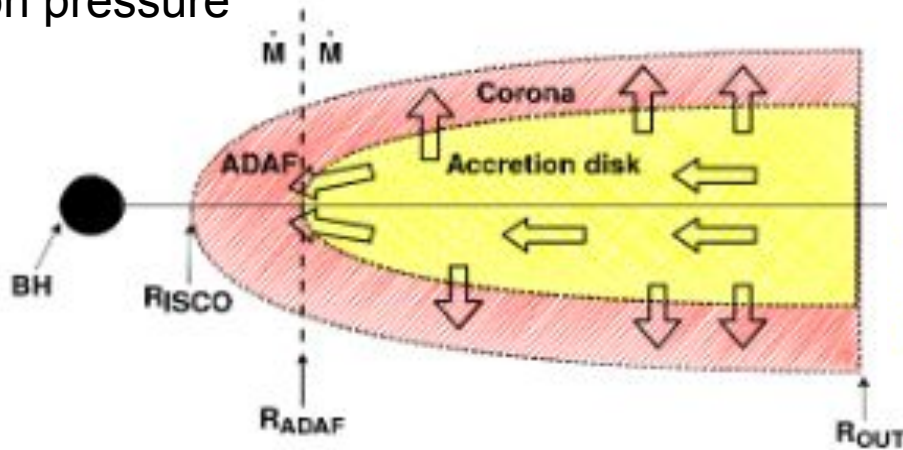


Credit: NASA/JPL-Caltech

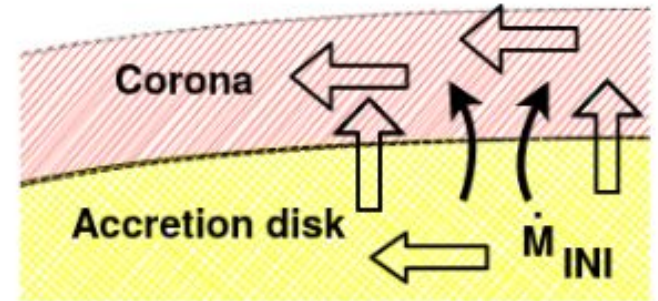
P1

Marzena Śniegowska, Mikołaj Grzędzielski,  
Bożena Czerny, Agnieszka Janiuk

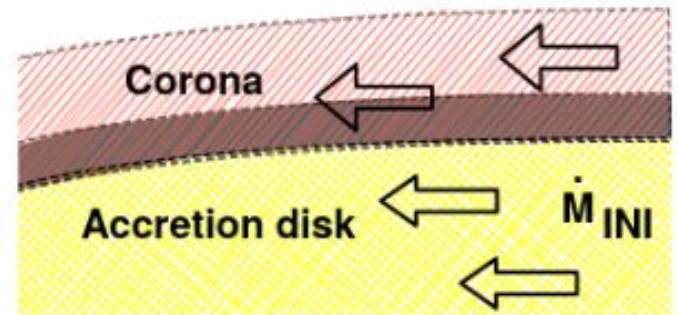
- ★ We model luminosity changes for objects with  $10$ ,  $10^5$ ,  $10^7$  solar masses
- ★ We use the time-dependent evolution of a black hole accretion disk unstable due to the dominant radiation pressure



Model A



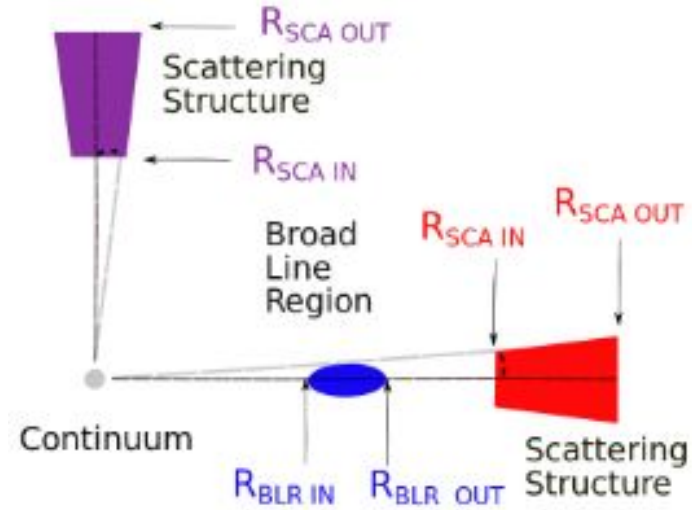
Model B



P2

**Marzena Śniegowska**, Swayamtrupta Panda, Bozena Czerny,  
Djorde Savic, Mary Loli Martinez-Aldama, Paola Marziani,  
Jian-Min Wang, Pu Du, Luka Popovic, Saraf, Chandra Shekhar

- ★ Spectropolarimetric data for 3 NLSy 1 candidates
- ★ We recover the viewing angles
- ★ We confirm the small values of the black hole mass in these sources and their high Eddington nature
- ★ We recover the observed H $\alpha$  line profile both in the natural and polarized light using the STOKES modelling. We recover the polarization fractions of the order of 0.2-0.5%



# To sum up this year:

## Published:

- ★ 'Modified models of radiation pressure instability applied to  $10$ ,  $10^5$ , and  $10^7 M_{\odot}$  accreting black holes', Śniegowska et al. 2023; Astronomy & Astrophysics, Volume 672, id.A19, 21 pp.
- ★ 'Spectropolarimetry and spectral decomposition of high-accreting narrow-line Seyfert 1 galaxies'; Śniegowska et al. 2023; Astronomy & Astrophysics, Volume 678, id.A63, 25 pp.

**GRANTS:** PRELUDIUM 2021/41/N/ST9/02280

Joined SDSS-V collaboration, so more exciting projects ahead of me

Thank you!

