



Annual Report

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Gonzalo Rojas García <u>rojas@camk.edu.pl</u> Supervisor: Dr Hab. Dariusz Graczyk CAMK Annual Meeting Warsaw, Poland January 31th, 2024

The systems and the Data

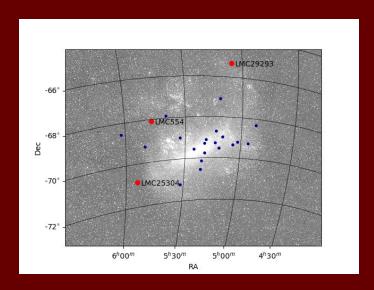
3 Late-type, long-period binaries from LMC

- LMC25304
- LMC29293
- LMC554

Data

- I and V band OGLE IV light curves.
- MIKE, HARPS and UVES spectra.

	LMC29293	LMC25304	LMC554	
R.A. (J2000)	05:05:37.37	05:54:19.44	05:47:08.27	
Decl. (J2000)	-65:17:13.4	-70:50:24.0	-68:04:12.0	
Period (days)	123.17	194.517	133.17	
T ₀ (HJD-2450000)	5573.83022	7146.40980	5565.68789	
Va (mag)	16.904	16.520	16.605	
I^a (mag)	15.971	15.527	16.149	
J^b (mag)	15.365 ± 0.025	14.870 ± 0.026	15.235±0.045	
H^b (mag)	14.895 ± 0.032	14.361 ± 0.029	14.905±0.074	
Kb (mag)	14.818 ± 0.044	14.290 ± 0.038	14.588±0.109	



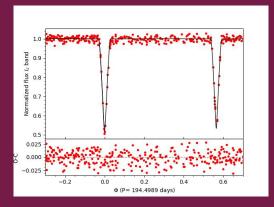
Location of the sources over LMC projection

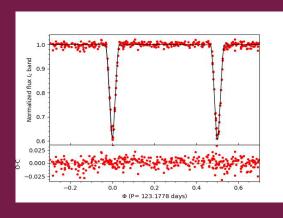
Magnitudes from the OGLE Collection of Variable Stars (Pawlak et al. (2016)) Infrared magnitudes from 2MASS 6X Point Source Working Database (Cutri et al. (2012)

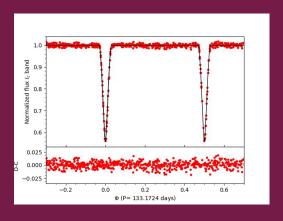
LMC25304

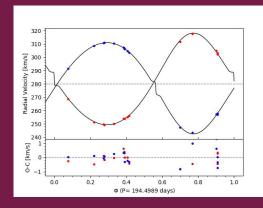
LMC29293

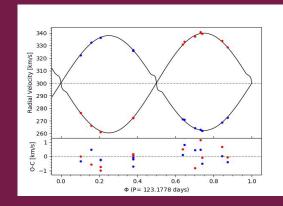
LMC554

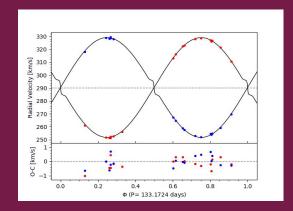








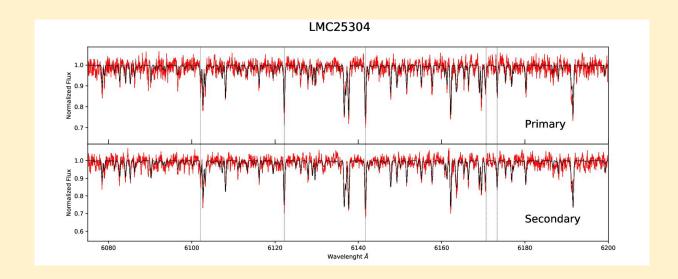




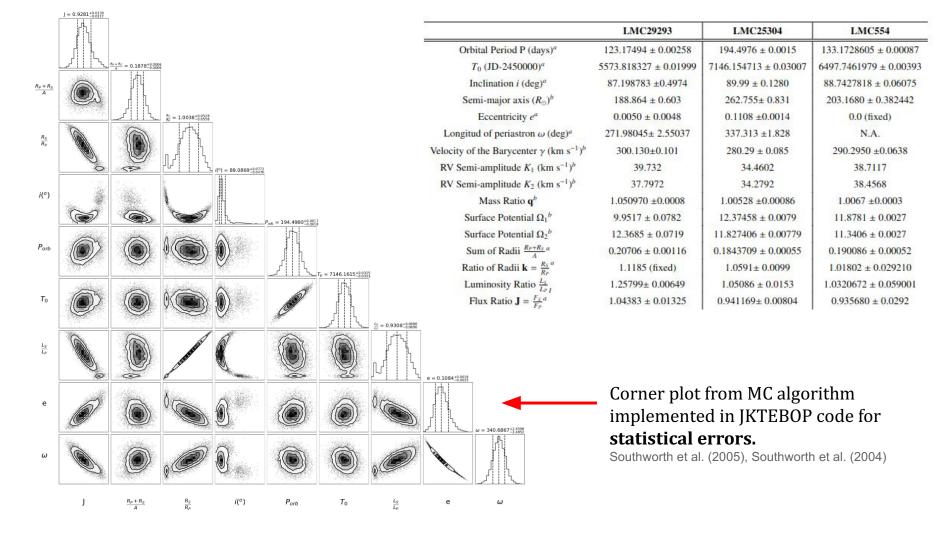
Atmospheric Analysis

- Grid Search in Stellar
 Parameters (GSSP)
 A. Tkachenko 2015
- SynthV LTE-based radiative transfer code by Tsymbal (1996).

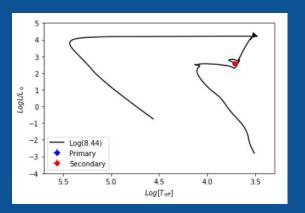
LLModels (Shulyak et al. (2004) MARCS (Gustafsson et al. (2008)

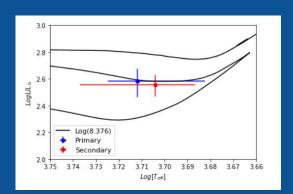


	LMC29293		LMC25304		LMC554	
	Primary	Secondary	Primary	Secondary	Primary	Secondary
T _{eff} (K)	5153+183	5125+272	5155+222	5080 ⁺¹⁸⁶ ₋₁₈₆	5517 ⁺³⁹⁶ ₋₃₉₆	5415 ⁺³⁰¹ ₋₂₅₇
[M/H] (dex)	$-0.42^{+0.16}_{-0.18}$	-0.27 ^{+0.23} -0.23	-0.58 ^{+0.24} -0.23	-0.62 ^{+0.20} -0.19	-0.35 ^{+0.22} _{-0.22}	-0.55 ^{+0.22} _{-0.28}
ξ (km s ⁻¹)	1.36+0.85	3.09+1.5	2.16+0.97	1.81 +0.69	$2.6^{+0.88}_{-0.79}$	2.01 +1.2 -1.0
$v \sin i (\text{km s}^{-1})$	9.334 +3.111	8.421+4.011	7.538 +3.3	8.368 +2.63	7.77 +3.9	8.43+4.0



Evolutionary stage



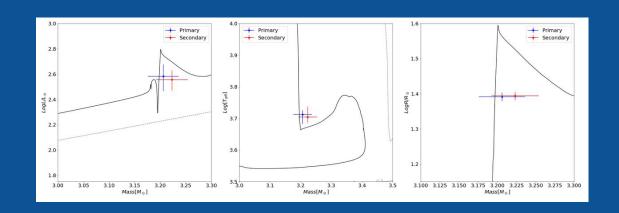


MESA Isochrones and Stellar Tracks (MIST) (Choi et al. (2016) and Dotter (2016))

LMC 25304 237.6 Myrs

LMC 29293 380.1 Myrs

LMC 554 301.9 Myrs



Other Activities

• Observation at Cerro Murphy Observatory with optical 0.8m telescope.

Contributed talk at EAS 2023, Kraków, Poland

 Part of the time domain group of Wide field spectroscopic telescope (WST)

• Orbyts Volunteer, partnered with LICEO MULTICULTURAL CLAUDIO FLORES SOTO, Cerro Alto, Chile.