

First dark matter search results from the LZ experiment

LIDINE 2022

Warsaw, Poland

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On behalf of the LZ collaboration

21 September 2022



- Black Hills State University
- Brandeis University
- Brookhaven National Laboratory
- Brown University
- Center for Underground Physics
- Edinburgh University
- Fermi National Accelerator Lab.
- Imperial College London
- Lawrence Berkeley National Lab.
- Lawrence Livermore National Lab.
- LIP Coimbra
- Northwestern University
- Pennsylvania State University
- Royal Holloway University of London
- SLAC National Accelerator Lab.
- South Dakota School of Mines & Tech
- South Dakota Science & Technology Authority
- STFC Rutherford Appleton Lab.
- Texas A&M University
- University of Albany, SUNY
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- University of Wisconsin, Madison



LZ Collaboration Meeting – September 8–11, 2021



Science and
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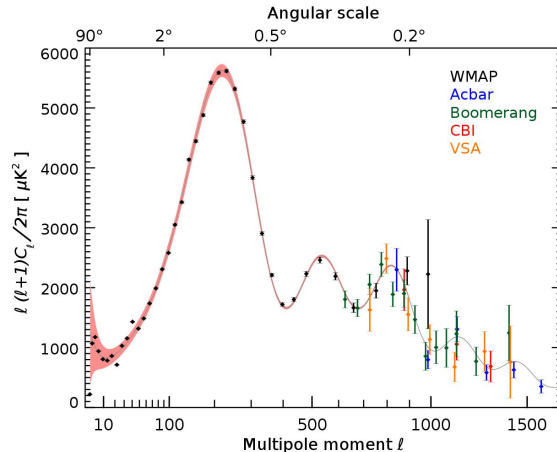


ibs Institute for
Basic Science

- Dark Matter Detector - particularly sensitive to WIMP candidates
- Dark Matter evidence from rotation curves, bullet cluster, CMB ...
- Nature of such a substance is inferred from observation



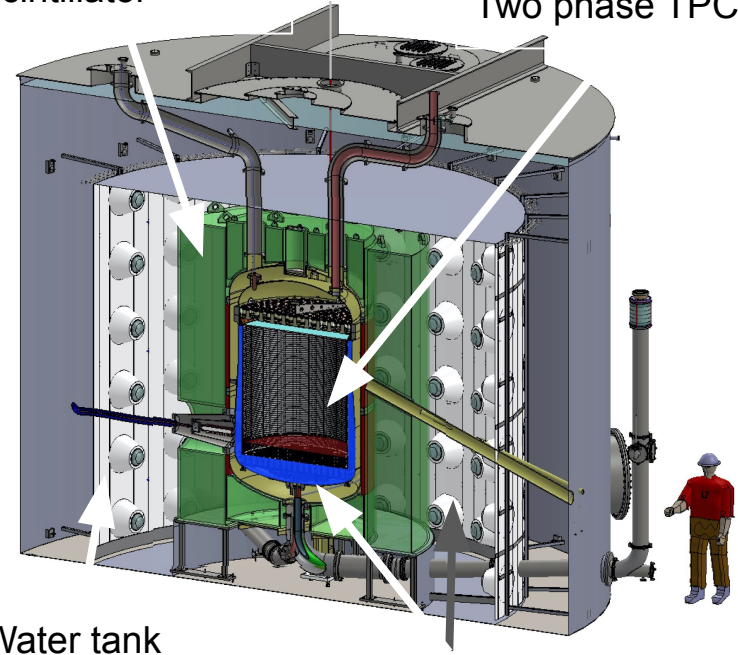
NASA/CXC/M. Weiss



NASA/WMAP Science Team

Gadolinium doped liquid scintillator

Two phase TPC



Water tank

Liquid xenon 'skin'

Principles of detection:

- Interactions produce primary scintillation light (S1) and ionisation electrons
- Secondary scintillation signal (S2) from drifted ionisation charge

Position reconstruction - hit pattern and drift time

- Enables separation of single and multiple scatters

Particle ID from S1/S2 ratio

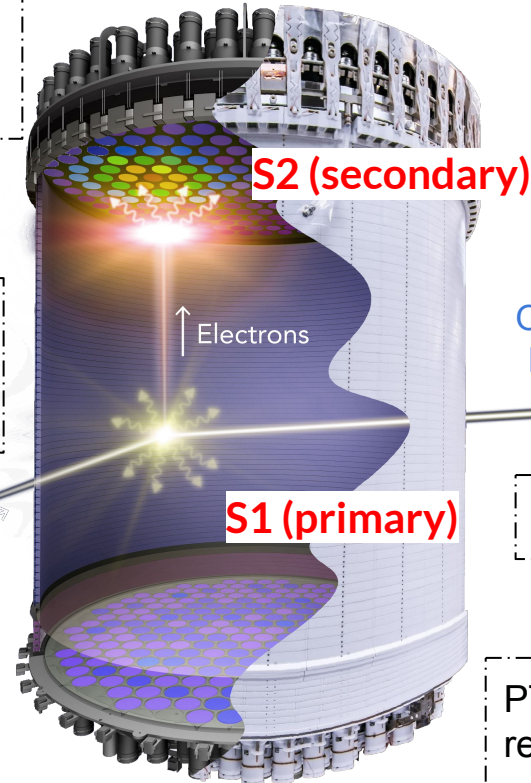
- Nuclear recoils (signal) and electronic recoils (background)

494 3 inch photomultiplier tubes

Ti field shaping rings to increase field uniformity

~1.5m height and diameter

Incoming Particle



4 wire mesh grids, bespoke woven define the drift and extraction fields

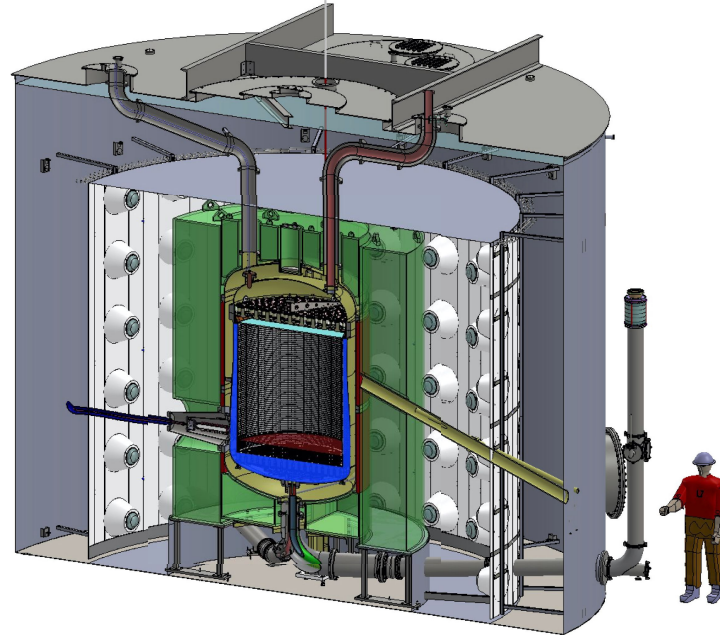
Outgoing Particle

7 tonnes of LXe

PTFE for efficient reflective ability

Xenon 'Skin'

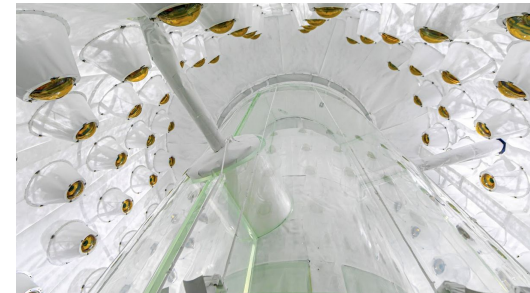
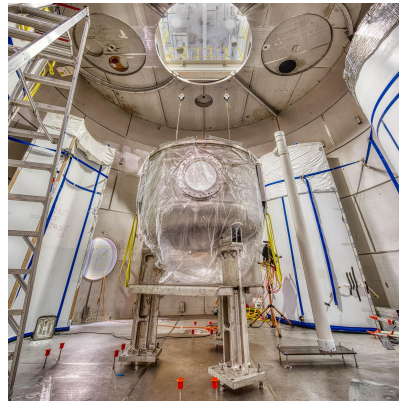
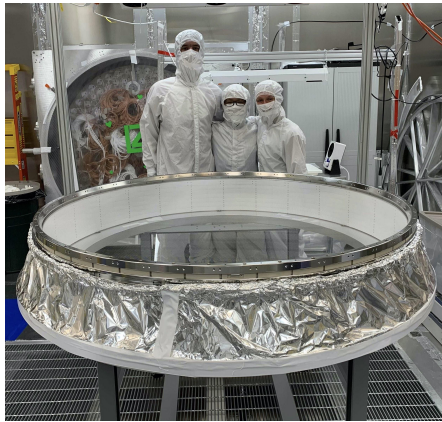
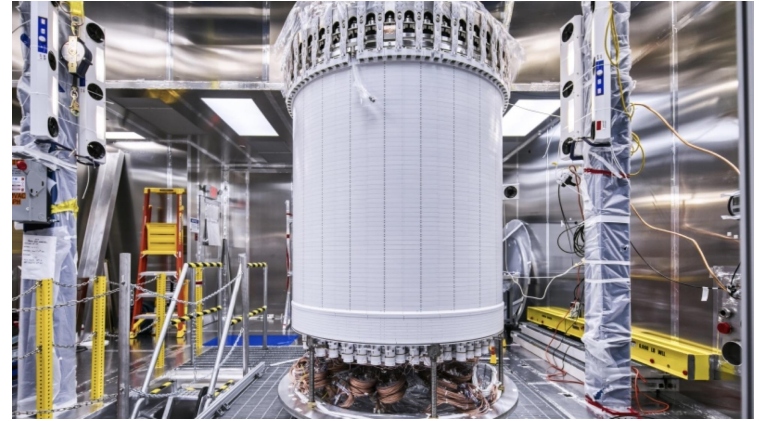
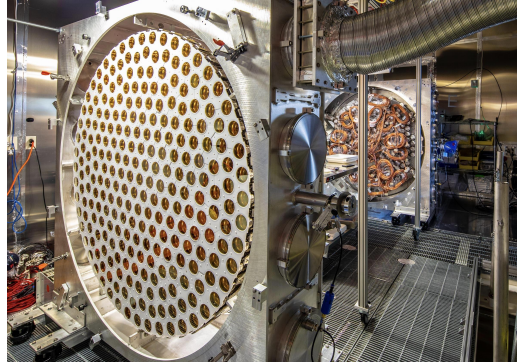
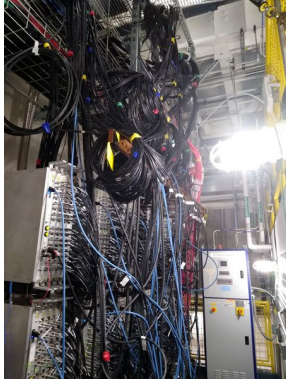
- Surrounds the TPC volume (~2 tonnes)
- Actively instrumented with 1" and 2" PMTs at the top and bottom - active veto.
- Optically isolated from other detectors
- Anti-coincident gamma ray detector



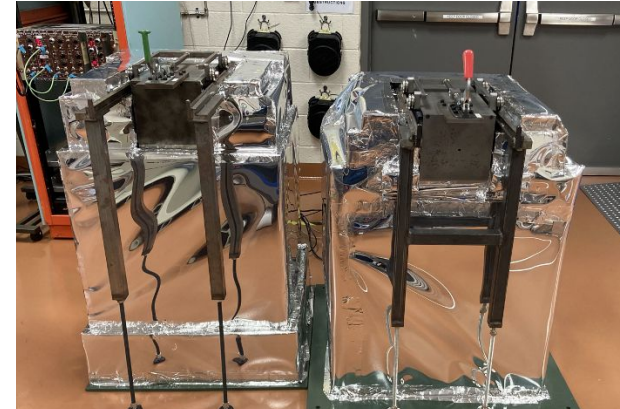
Outer Detector

- Gadolinium liquid scintillator surrounding detector.
- Located within an instrumented water tank (8" PMTs)
- Anti-coincident neutron and gamma detector

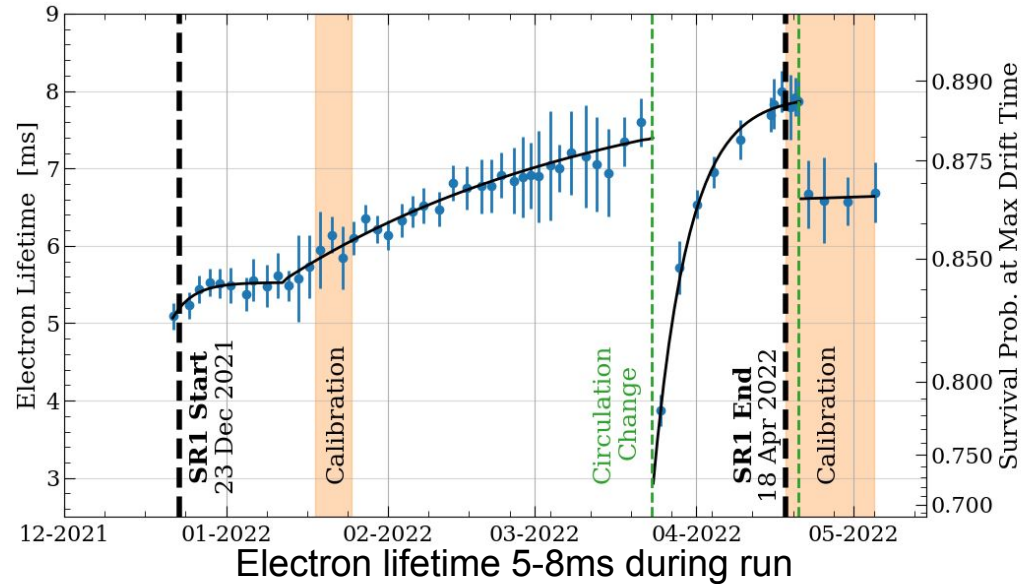
Vetos providing ~88% neutron detection efficiency



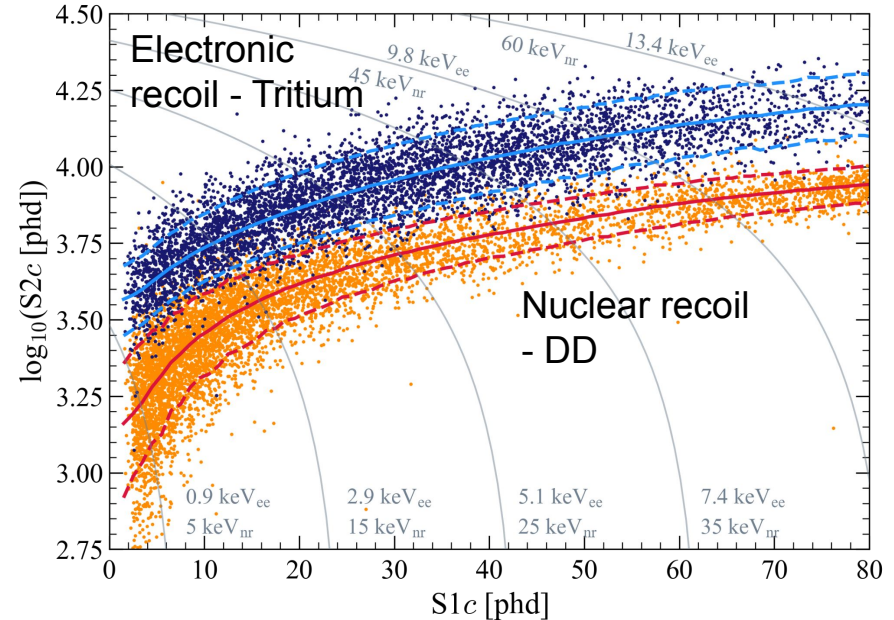
- Material selection
 - Material screening for U,Th,K,Co determined material selection.
 - Screening in Ge facilities (Boulby (UK), SURF (US)), ICP-MS, Radon
- Shielding
 - LZ is 4850 ft (~1 mile) underground - muon flux reduced by over a million.
 - Water tank and Xe self shielding
- Cleanliness/Rn daughter
 - Parts were cleaned and construction performed in a cleanroom
 - Requirements: Dust $<500 \text{ ng/cm}^2$ on all LXe wetted surfaces
 - Rn-daughter plate-out on TPC walls $<0.5 \text{ mBq/m}^2$
- Xenon contaminant removal
 - $^{85}\text{Kr}/^{39}\text{Ar}$ Removal
 - Purification underground with a warm getter



- Goals:
 1. Demonstrate physics capabilities of the detector (not blinded)
 2. Perform competitively with other similar experiments
- Key information:
 1. Livetime 60 days (Data from 23rd Dec 2021 to 12th May 2022)
 2. PMTs: >97% operational throughout run
 3. Liquid temperature: 174.1 K (0.02%)
 4. Gas pressure: 1.791 bar(a) (0.2%)
 5. Gas circulation: 3.3t/day
 6. Drift field: 193 V/cm (32 kV cathode, uniform to 4% in fiducial volume)
 7. Extraction field: 7.3 kV/cm in gas (8 kV gate-anode ΔV)



- Many sources:
 - ^{83m}Kr : monoenergetic ERs, 32.1 keV and 9.4 keV
 - ^{131m}Xe : monoenergetic ER, 164 keV
 - **CH_3T (tritium): beta spectrum - Q-value: 18.6 keV**
 - **Deuterium-deuterium (DD): triggered 2.45 MeV neutrons**
 - Activation lines
 - AmLi: continuum neutrons, isotropic
 - Alphas
 - And more (^{220}Rn , YBe, ^{252}Cf , ^{22}Na , ^{228}Th , etc)
- Some uses:
 - Tune the position reconstruction algorithm in horizontal plane
 - Flat fielding of S1 and S2 signals
 - NEST (<https://nest.physics.ucdavis.edu/>) model tuned to the response of tritium and validated on DD



- Light gain $g1$: 0.114 ± 0.002 phd/photon
- Charge gain $g2$: 47.1 ± 1.1 phd/electron
- Single electron size: 58.5 phd

Electronic Recoils

Dissolved contaminants (beta decay):

- ^{214}Pb (^{222}Rn daughter)
- ^{212}Pb (^{220}Rn daughter)
- ^{85}Kr

Dissolved contaminants (electron capture):

- ^{37}Ar
- ^{127}Xe
- ^{124}Xe (double e-capture)

Gamma emitters in detector materials

^{60}Co , ^{40}K , ^{238}U chain, ^{232}Th chain

Nuclear Recoils

- Neutron emission from spontaneous fission and (α ,n)
- ^8B solar neutrinos

Expected numbers of events:

Electronic recoils: $276 + [0, 291]$ from ^{37}Ar

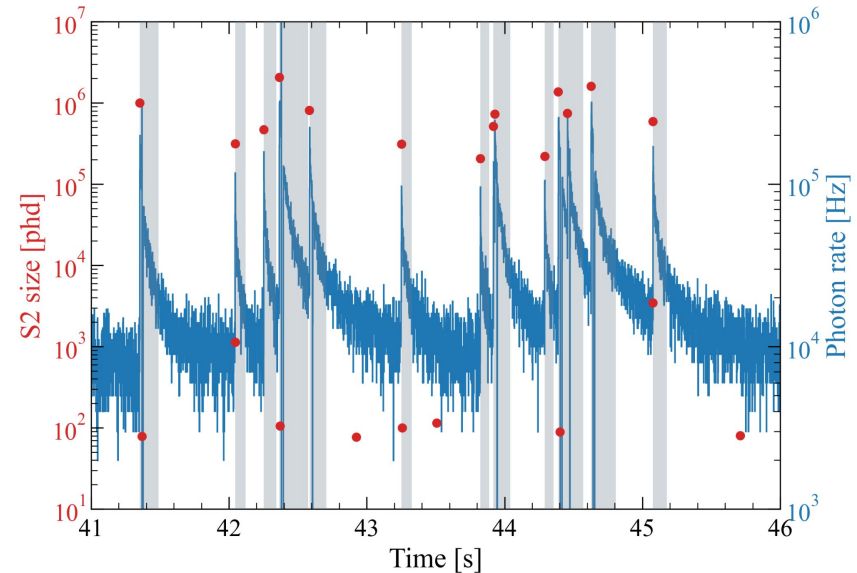
Nuclear recoils: 0.15

There is contamination from spurious signals:

- Electron and photon rates are higher after larger S2s
- Isolated S1 and S2 pulses from charge or light insensitive regions
- Gas events

Solution - data quality cuts:

- **Pulse based cuts**
Targets accidental pileup topologies
Cut mainly on hit pattern and pulse shape
Tuned with tritium and AmLi calibration data
- **Timing cuts**
Removes the high rates of electron and photon emission following S2s



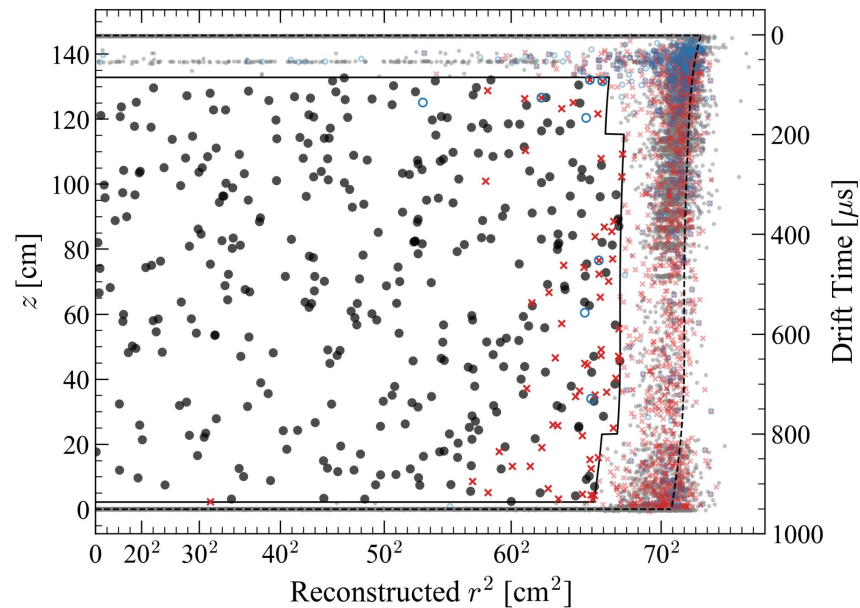
Vetos considered in anti-coincidence with the TPC:

- Prompt OD and Skin cut
 - Used for vetoing gammas
 - Significantly reducing the L,M ^{127}Xe background
- Delayed OD cut
 - Looking for neutron capture on Gd
 - ~ 200 keV threshold
 - In situ expected rate of $^{+0.20}$

Fiducialisation:

- Inner 5.5 tonnes of Xenon considered for the analysis
- Optimise fiducial volume and lower S2 threshold together

- Events surviving all selections
- ✗ Skin-prompt-tagged events
- OD-prompt-tagged events



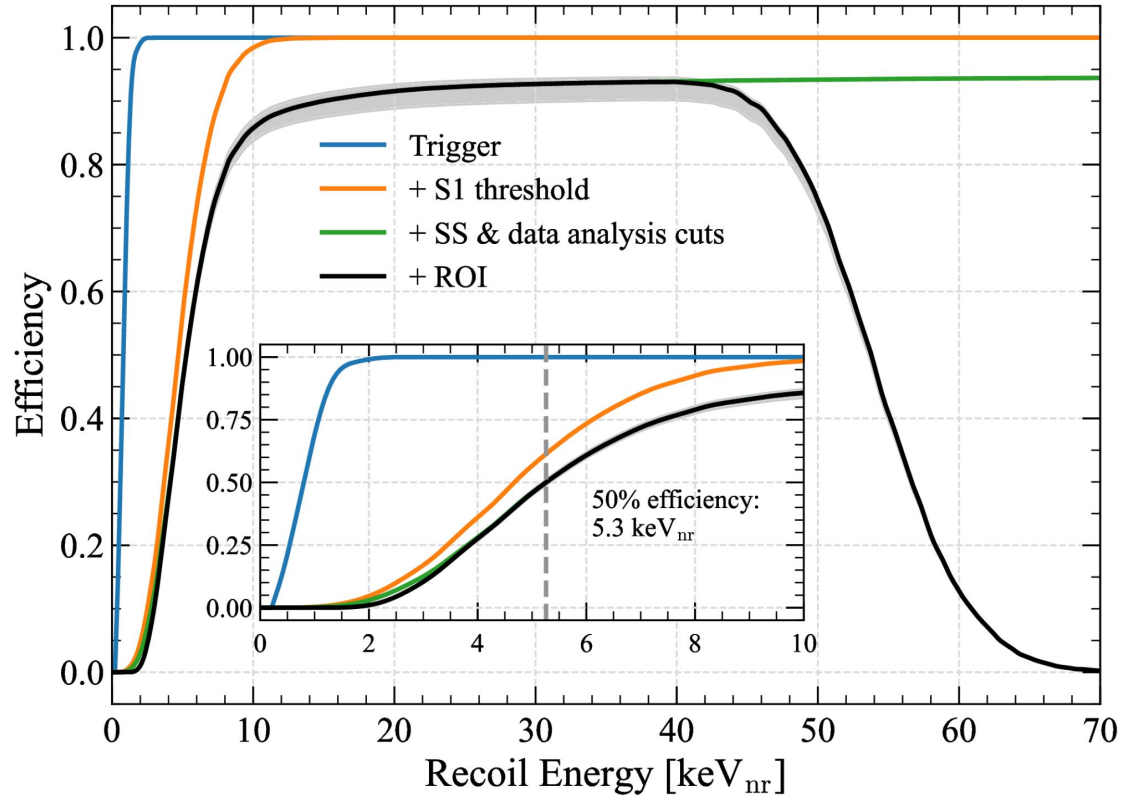


Trigger - all events triggered

S1 threshold - apply S1 three fold coincidence requirement

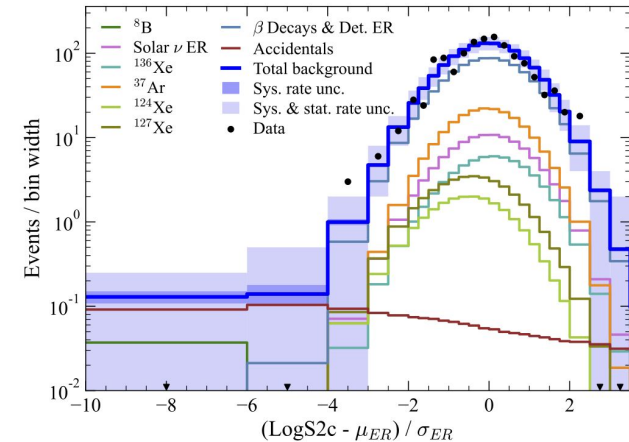
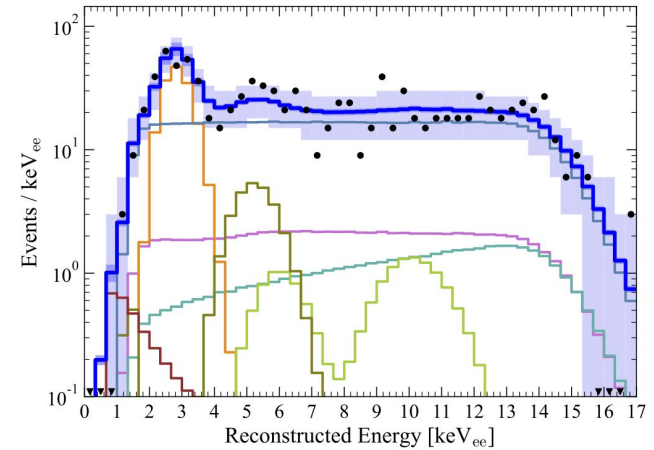
SS & data analysis cuts - single scatter cut and application of cuts previously considered

ROI - Considering events in the energy range shown, cut applied in S1, S2 space



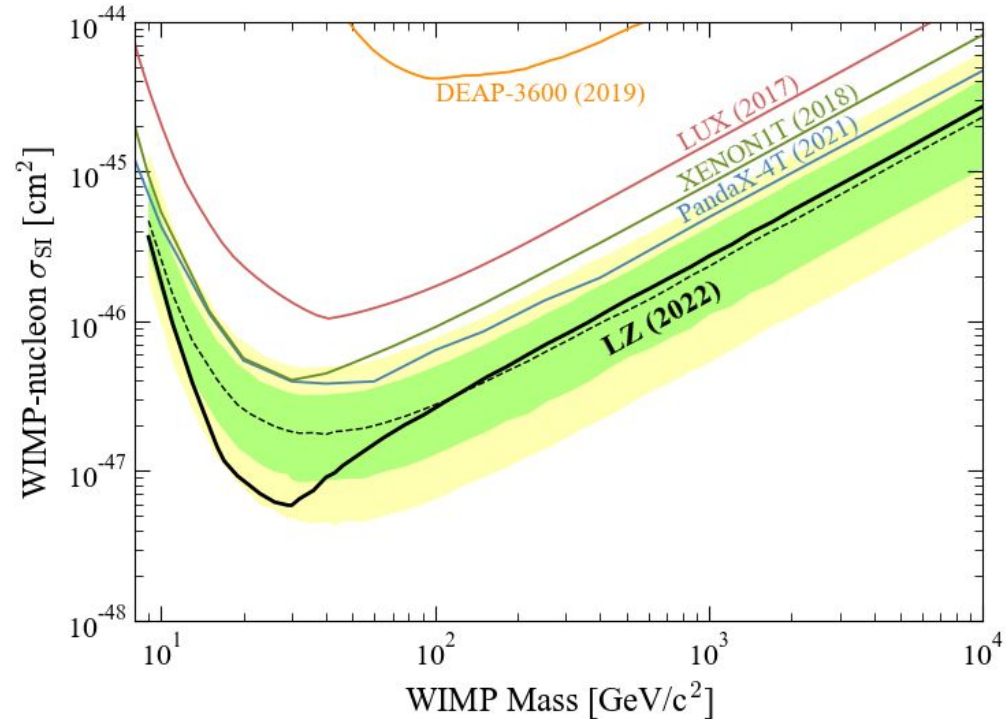
Source	Expected Events	Best Fit
β decays + Det. ER	218 ± 36	222 ± 16
ν ER	27.3 ± 1.6	27.3 ± 1.6
^{127}Xe	9.2 ± 0.8	9.3 ± 0.8
^{124}Xe	5.0 ± 1.4	5.2 ± 1.4
^{136}Xe	15.2 ± 2.4	15.3 ± 2.4
^8B CE ν NS	0.15 ± 0.01	0.15 ± 0.01
Accidentals	1.2 ± 0.3	1.2 ± 0.3
Subtotal	276 ± 36	281 ± 16
^{37}Ar	[0, 291]	$52.1^{+9.6}_{-8.9}$
Detector neutrons	$0.0^{+0.2}$	$0.0^{+0.2}$
30 GeV/ c^2 WIMP	–	$0.0^{+0.6}$
Total	–	333 ± 17

Best fits at all masses are for zero WIMP events



90% CL upper limit on WIMP-nucleon
 σ_{SI} is $5.9 \times 10^{-48} \text{ cm}^2$ at $30 \text{ GeV}/c^2$
WIMP mass

- Limit constructed using frequentist profile likelihood ratio (PLR)
- Two sided
- Signal rate constrained to be positive only



Results here are only a small fraction of the total exposure of the experiment
(~17x more exposure)

LZ is multi-physics device:

- Solar axions and axion like particles
- Neutrinoless double beta decay ($^{136}\text{Xe}/^{134}\text{Xe}$)
- Low mass dark matter (S2 only, Migdal)
- Higher mass dark matter (EFT)
- Etc.

Leading Xenon Researchers unite to build next-generation Dark Matter Detector

SURF is distributing this press release on behalf of the DARWIN and LZ collaborations

July 20, 2021

- Consortium formed between LZ, XENON, DARWIN
- Successful XLZD meeting 27-29 June 2022 at Karlsruhe Institute of Technology
- <https://xlzd.org/>
- [White paper \(2203.02309\)](#)

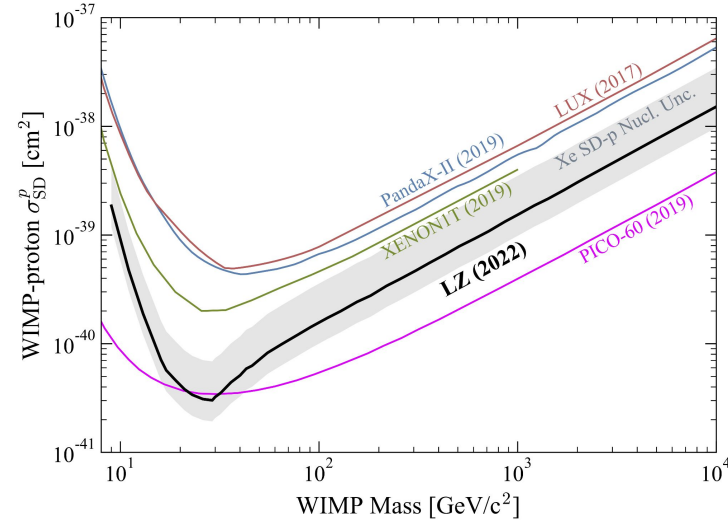
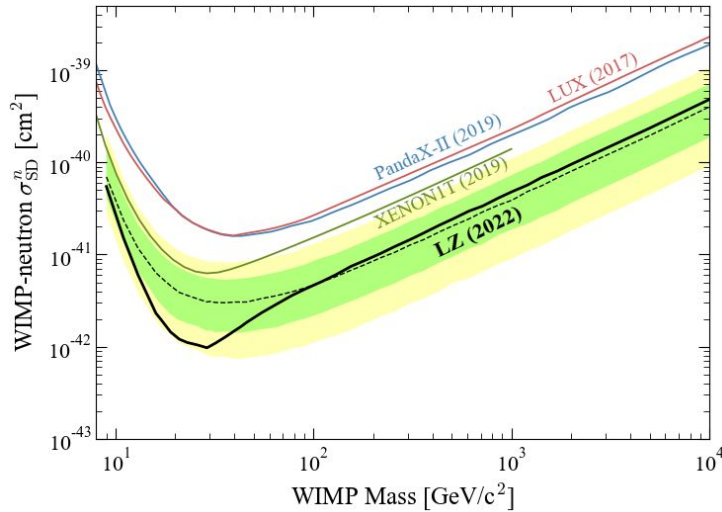


LZ is now the most sensitive experiment in the world in the considered mass range. See the paper: [arXiv:2207.03764](https://arxiv.org/abs/2207.03764)

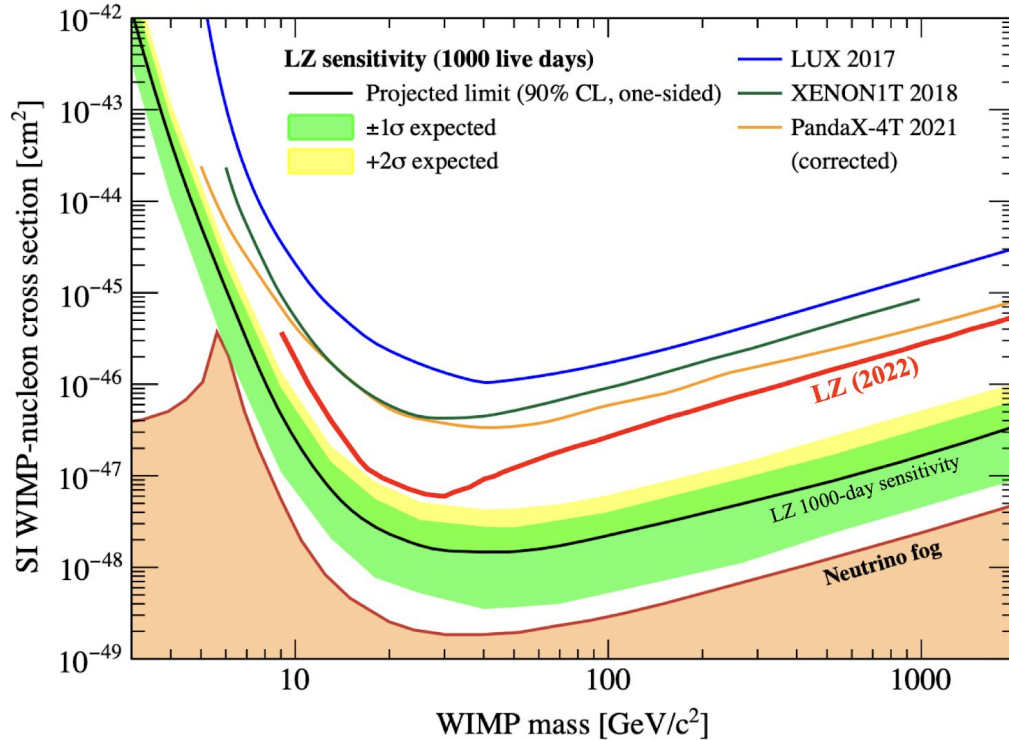
The detector is performing well, and is producing high quality data

Exciting times still lie ahead...

Backup



Uncertainty band to on nuclear form factor for Xe (*). "Brazil" band omitted for clarity
(*) [P. Klos, J. Menéndez, D. Gazit, and A. Schwenk Phys. Rev. D 88, 083516 \(2013\)](#)



90% CL minimum (one sided) of
 $1.4 \times 10^{-48} \text{ cm}^2$ at $40 \text{ GeV}/c^2$
from Phys. Rev. D 101, 052002 (2020)

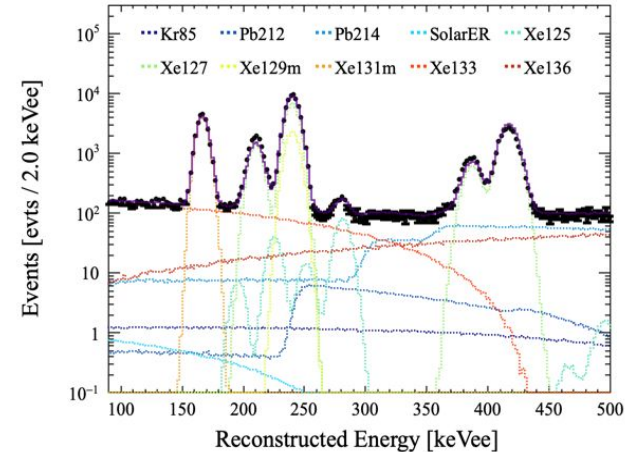
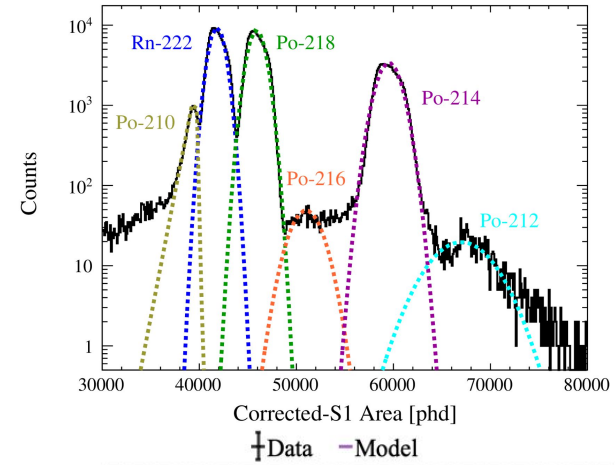
^{214}Pb is dominant electronic recoil background (^{222}Rn decay chain)

Constrain the rate within the WIMP search:

1. Constraining backgrounds in the sideband
2. Alpha tagging

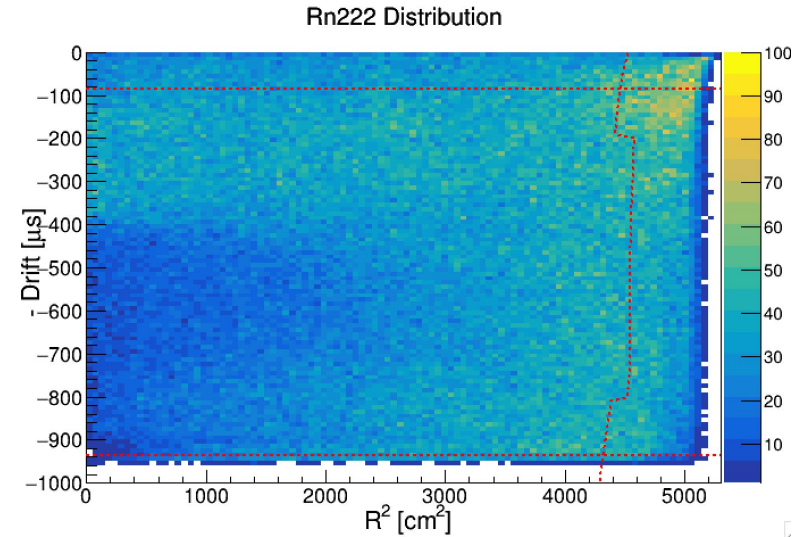
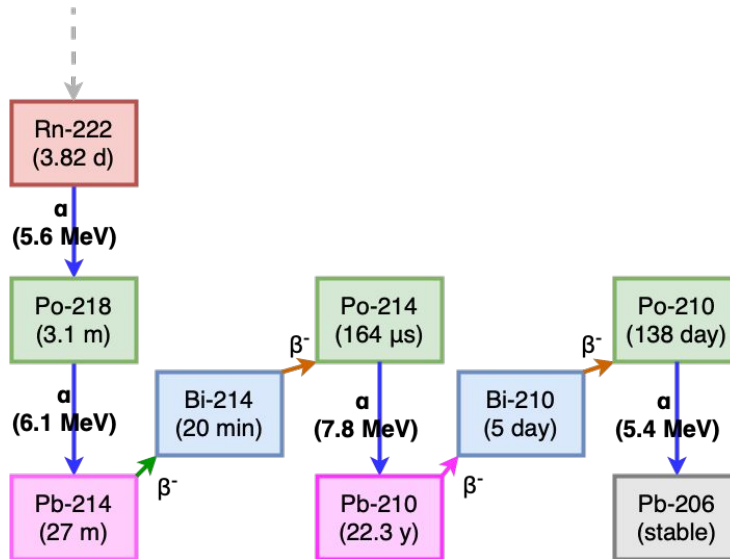
Consistent with assay campaign

Isotope (decay)	Activity [$\mu\text{Bq/kg}$]
^{222}Rn (alpha)	4.37 ± 0.31 (stat)
^{218}Po (alpha)	4.51 ± 0.32 (stat)
^{214}Pb (beta)	3.26 ± 0.13 (stat) ± 0.57 (sys)
^{214}Po (alpha)	2.56 ± 0.21 (stat)



^{222}Rn not uniformly distributed.

Stratification in LXe flow is a possible tool to reject ^{214}Pb in future



Argon

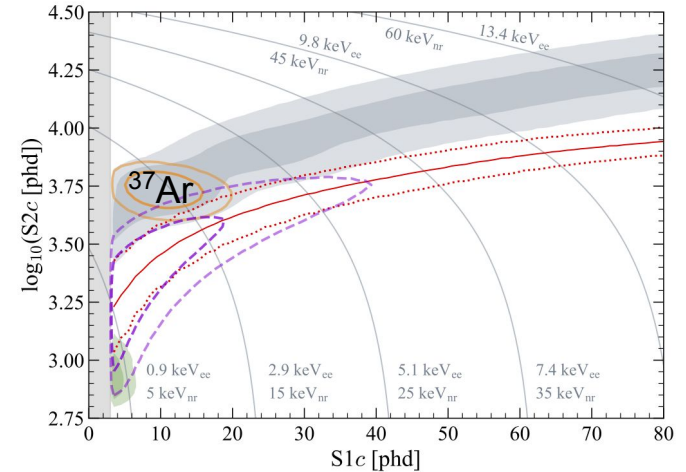
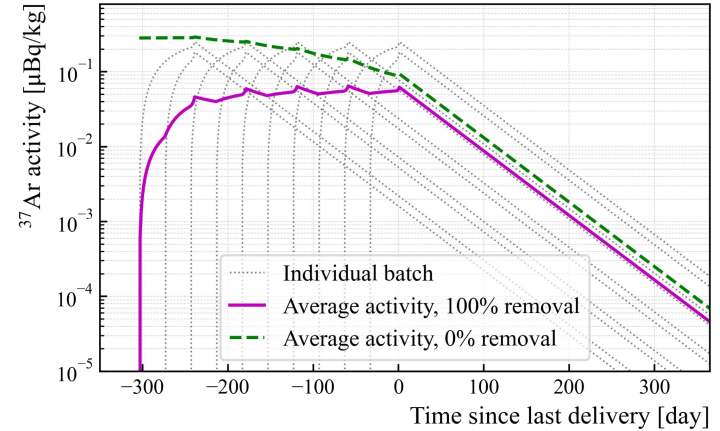


^{37}Ar decays ($T_{1/2} = 35$ d, monoenergetic
2.8 keV ER deposition)

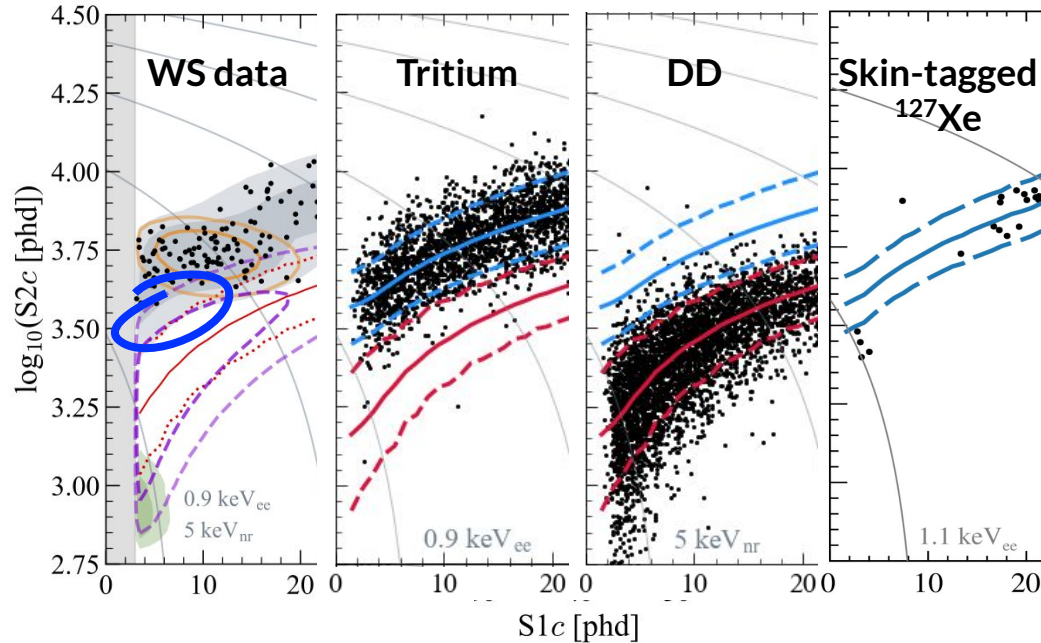
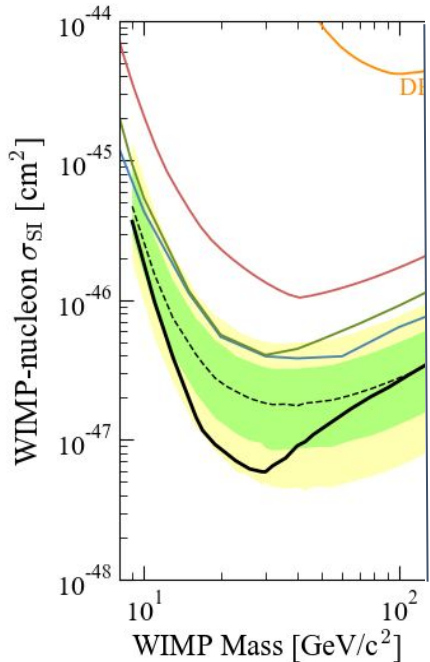
Predominant source of argon in LZ is
through cosmogenic spallation

LZ Collaboration, Phys. Rev. D 105, 082004 (2022), [2201.02858](#)

Activity estimates can be formed showing
approximately 100 decays in data (large
uncertainty)



- Downward fluctuation in the observed upper limit near $30 \text{ GeV}/c^2$ is a result of the deficit of events under the ^{37}Ar population.
- Tritium** data analyzed identically to WS data; well-covered
- DD** data also shows deficit region is well-covered.



- M-shell decays of ^{127}Xe populate near deficit region. Observed rate of M-shell decays with coincident γ -ray tagged by the skin is consistent with expectation, given signal efficiencies.
- Conclusion: Deficit appears consistent with under-fluctuation of background.

Selection description	Events after selection
All triggers	1.1×10^8
Analysis time hold-offs	6.0×10^7
Single scatter	1.0×10^7
Region-of-interest	1.8×10^5
Analysis cuts for accidentals	3.1×10^4
Fiducial volume	416
OD and Skin vetoes	335

Doke Plot - energy calibration

