Scintillation Light Detection Performance for the DUNE ND-LAr 2x2 Modules

Anja Gauch for the DUNE Collaboration LIDINE 2022 September 21-23

Deep Underground Neutrino Experiment - DUNE



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- International experiment for neutrino science
- Intended to determine CP violating phase in leptonic sector and neutrino mass hierarchy
- Two neutrino detectors placed 1300 km apart in the world's most intense 1.2 MW neutrino beam
- Beam is upgradable to 2.4MW







ND-LAr



DUNE ND-LAr



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- 5 by 7 optical separated modules
- Each 1x1x3 m³
- Need modularity to increase:
 - Precision
 - Light spatial resolution
 - Associate unique charge attachment to each interaction (Neutrino pile-up)



Beam spill in ND-LAr. Highlights the challenge of associating fast neutron induced energy deposits to a neutrino vertex



A. Abed Abud et al. [DUNE], instruments 5, no.4, 31 (2021) doi:10.3390/instruments5040031 [arXiv:2103.13910 [physics.ins-det]]. Anja Gauch

Dune Near Detector Modules

- Currently, prototyping the DUNE ND-LAr Modules
- ProtoDUNE-ND is a 2x2 ND-LAr prototype
- Module-0 is the first of four modules in the $2x^2$ ND-LAr



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Role of ProtoDUNE-ND in DUNE



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- Demonstrate modular TPC performance in Fermilab's NuMI neutrino beam.
- Test includes:

P. Adamson, K. Anderson, M. Andrews, R. Andrews, I. Anghel, D. Augustine, A. Aurisano, S. Avv aku-mov, D. S. Ayres and B. Baller, et al. Nucl. Instrum. Meth. A 806, 279-306 (2016) doi:10.101 6/j.nima.2015.08.063 [arXiv:1507.06690 [physics.acc-ph]].

- Charge and light system
- Data acquisition system (DAQ)
- Charge reconstruction with pixel-based readout
- Combined charge and light reconstruction
- Physics performance with neutrino interactions



A. Abed Abud, 2021, Instruments 5, no.4, 31

Light Readout System (LRS) in $u^{\scriptscriptstyle b}$ UNIVERSITÄT **ProtoDUNE-ND Modules**

- 24 LCMs and 8 ACL per Module ٠
- SiPM-based detectors for efficient collection of single UV photons with large surface coverage (6 SiPM per ACL 2 SiPM per LCM)
- Placed inside the field-shaping structure to • increase light yield and decrease dead space
- LRS provides 30% optical detector coverage •
- Photon detection efficiency of 0.6% for LCM, for ACL 0.2%
- Photomultiplier Tubes (PMT) are too big for • modular TPCs like DUNF ND-I ar
- ACL has better spatial resolution, LCM has better timing resolution

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Light Detector for DUNE ND-N. Auger et. al., Instruments 2, no.1, 3 (2018) doi:10.3390/instruments 2010003 [arXiv:1711.11409 [physics.ins-det]].

 ArCLight is based on the ARAPUCA principle of light trapping

Machado, A.; Segreto, E. ARAPUCA a new device for liquid argon scintillation light detec tion. Journal of Instrumentation 2016, 11, C02004–C02004. doi:10.1088/1748-0221/11/0 2/c02004.)



• Light Collection Modules have WLS fibers as bulk structure



Similar working principles:

- Vacuum ultraviolet (VUV) scintillation light transitions when traveling through TPB coated surfaces
- Light enters bulk structure which acts as light trap
- In the bulk structure the photons are measured by SiPMs

For more informations about the ArCLight see Jan Kunzmann's talk

Dive 2 ProtoDUNE-ND Modules u^{\flat} **Tested at Bern** BERN DUNE

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- Module 0:
 - 8 days —
 - 60 million self-triggered cosmic ray-induced events
- Module 1:
 - 3 days —
 - 20 million self-triggered cosmic ray-induced events
- Trigger:
 - **Charge**: self-trigger mode, tile triggers when channel-level charge threshold is exceeded (~100keV)
 - **Light**: LCM provide external trigger to charge readout system (~ 30 photoelectrons (p.e.))
- To merge data, LRS trigger signal is written to the charge readout data stream



Neutrino like interaction in Module-0





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SiPM gain calibration

- Hamamatsu S13360-6050CS SiPMs
- Calibration with LED before taking cosmic data
- Bias voltage of SiPM channel adjusted to obtain uniform gain distribution across the channels



Positions of LEDs on Module



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Time resolution of LCM

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- Cosmic muons traversing the TPC were used to extract the time resolution
 - Waveforms are processed with a Fourier transform to increase ability to measure the front edge.
 - Linear fit to baseline and front edge
 - Crossing point provides a robust single-channel event time
- For large signals, timing resolution approaches ~2 ns

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CIVETime resolution of LCM

- Timing resolution of ~2 ns
- DUNE must be capable with its timing resolution to:
 - identify Michel electrons
 - Study event pile-up in neutrino interactions
- Example of stopping muon and delayed Michel electron

Wave form with sample number of 10ns (Module 0)

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Gamma conversion event, charge pixel view

Gamma conversion event, triggering wave form from SiPM sum channel

- Charge is self triggering, therefore the triggering threshold of the LRS determines which charge events can be matched to light events
- Trigger signal is based on sum channels of the LCM light readout
- Gamma conversion with total energy deposit of about 20MeV
- LCM sum signal peaks at -24000 ADC counts which means that the effective threshold was 1.6MeV
- Given signal clarity (compared to Module 0), operation below threshold of 1MeV possible

Photon detection efficiency (PDE)

- Expected PDE is dependend on the efficiency of:
 - TPB - WLS $Q_{PDE} = \frac{1}{2} \varepsilon_{TPB} T_{425} \varepsilon_{WLS} \varepsilon_{SA} \varepsilon_{coll} \varepsilon_{SiPM}$
 - Average spectral acceptance
 - Reflecting surfaces to deliver photons to the SiPM openings

- SiPMs
- PDE is estimated by comparing the data and Geant4 simulation.

Photon detection efficiency (PDE)

Module 0:

- PDE for data across both TPCs of the module
- ACL 7 missing because of turned of charge tile

Module 1:

- PDE for data across one TPC of the module
- Improvement for the ArCLights when compared to Module 0 due to addition of mirrors on bordering edges.

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Summary

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- DUNE ND-LAr is a modular detector using compact light detectors borrowing principles from ARAPUCA.
- Prototype experiment named ProtoDUNE-ND is in construction and will test LRS under cosmics (Bern) and neutrino interactions (Fermilab)
- LRS has a timing resolution of ~2ns for LCM
- LRS has PDE of 0.6% and 0.2% for LCM and ArCLights, respectively.
- Publication in progress on analyzing LRS data under cosmic at Bern for a single module.

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Back up

ArCLight

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- Low-volume large-area light detector
- Development, production and testing in Bern
- Fully dielectric to be placed in drift field along drift direction
- ArCLight achieves spatial resolution requirement (aim for 5cm)
 - associate a recoil proton with the corresponding neutrino event

ArCLight

- VUV photons 128 nm
- shift to blue 430 nm
- shift to green 490 nm
- Blue photons can penetrate through the dichroic mirror
- green photons can not
- Size: 30 cm x 28 cm x 1cm

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- Front (left) contains 4900 charge-sensitive pixels with 4.43mm pitch, facing the cathode
- Back (right) contains a 10x10 array of LArPix ASICs
- The tile dimensions are 31cm x 32cm
- 8 anode tiles per TPC in ND-LAr 2x2 modules
- Module-0 comprises 78400 instrumented LArTPC pixels

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