COSMIC MAGNETIC FIELDS KRZYSZTOF NALEWAJKO, CAMK PAN

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Introduction

ORGANIZATION

- Registration open until Sunday, March 6th.
- Tuesdays from March 1st to June 21st 14 meetings. No lecture on April 19th, May 3rd, May 17th.
- First hour (11:15 12:00) on phenomenology (observational evidence), Second hour (12:15 13:00) on theory.
- Lecture materials available after each lecture [<u>https://events.camk.edu.pl/event/37/</u>]:

 slides by end of Tuesday,
 - notes by end of Thursday.
- 2 lectures (June 7th, 14th) reserved for presentations, final lecture (June 21st) for the exam.
- Credit points: problems (max 50), presentation (max 50), exam (max 50), individual interview.
- Problems: assigned regularly, 1-page reports by next lecture, please don't clone them.
- Presentations: journal club of classic papers (suggestions welcome).
- Grades: no pass (0-59), 3 (60-79), 4 (80-99), 5 (100-150).

questions, comments, suggestions: <u>knalew@camk.edu.pl</u>

SCOPE OF THIS LECTURE

- phenomenology: Earth, planets, Sun, stars, pulsars/ magnetars, galaxies, AGN (relativistic jets), extragalactic
- theory: electrodynamics, MHD, waves, instabilities, shocks, reconnection, particle acceleration, radiation, dynamo

BASIC QUESTIONS

- What are magnetic fields?
- Where in the Universe do we find magnetic fields?
- How do we know that magnetic fields are there?
- What is the role of magnetic fields in the Universe?
- How important are they?

COSMIC RADIATION BACKGROUND



GAMMA-RAY SKY FROM FERMI/LAT (5 YEARS, > GEV)



HILLAS PLOT



cgs units $1 \text{ G} = 10^{-4} \text{ T}$