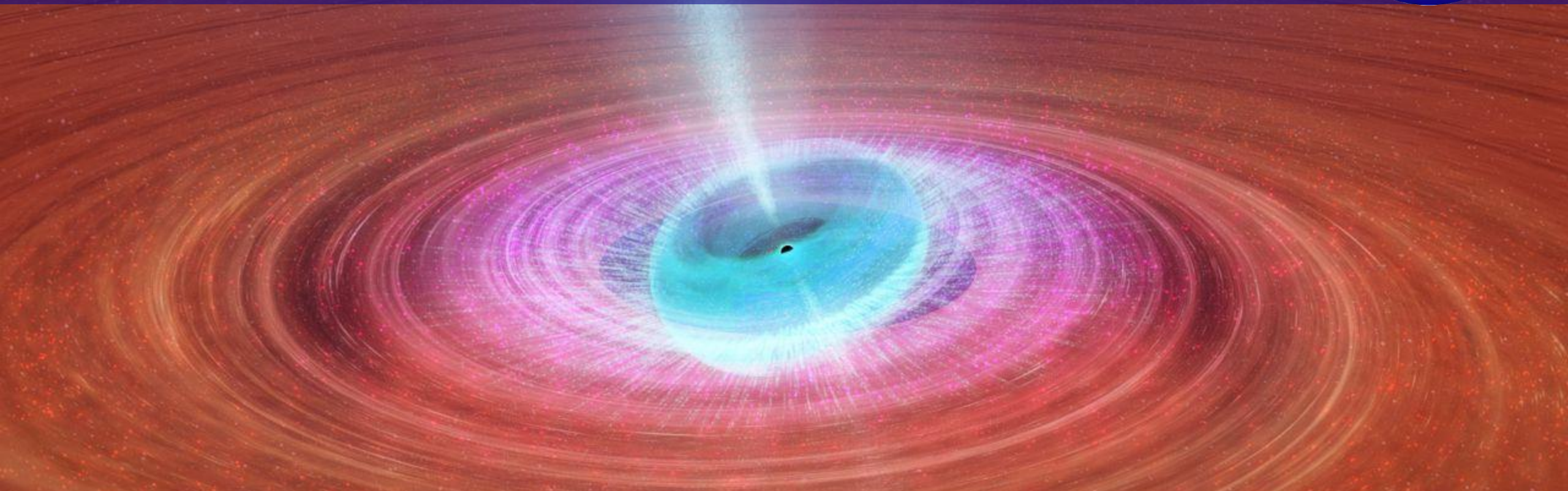


Picosecond Intensity Interferometry to resolve accreting sources



G.Koziol, PhD student, supervisor: R.Walter

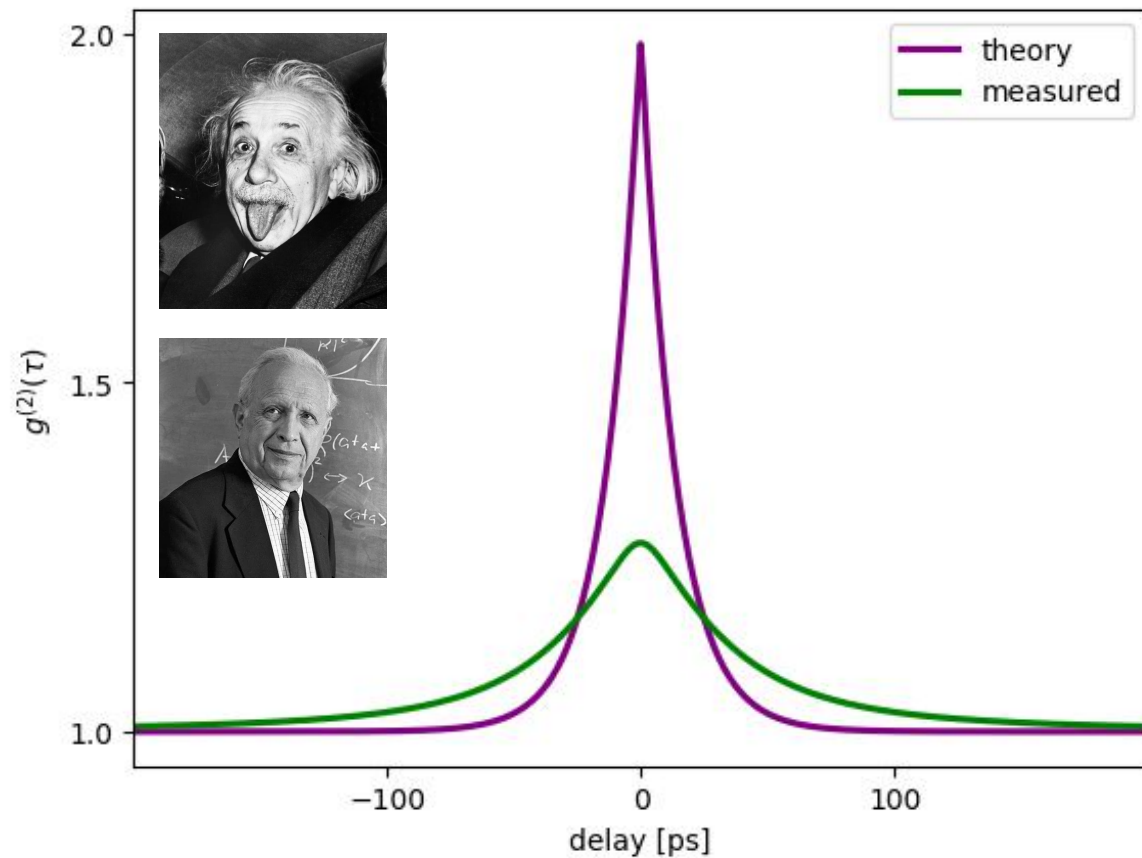
History

2

$$g^{(2)}(\tau, d) \propto \langle I_1(t) I_2(t + \tau) \rangle$$

$$\vdots$$

$$= 1 + |\mathcal{F}(f)|^2 \cdot |V(d)|^2$$



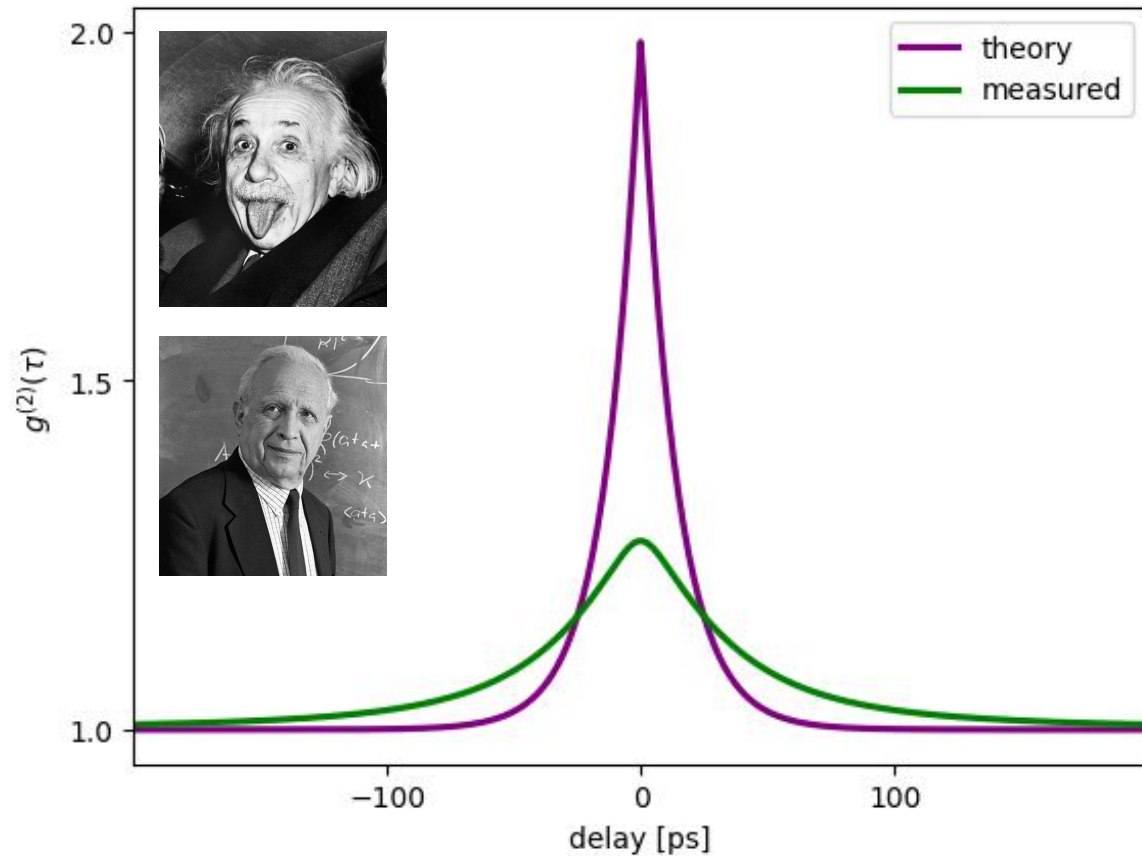
- First experimental proof : R.Hanbury Brown & R.Q.Twiss (1956)
- Measurement of angular diameter of Vega



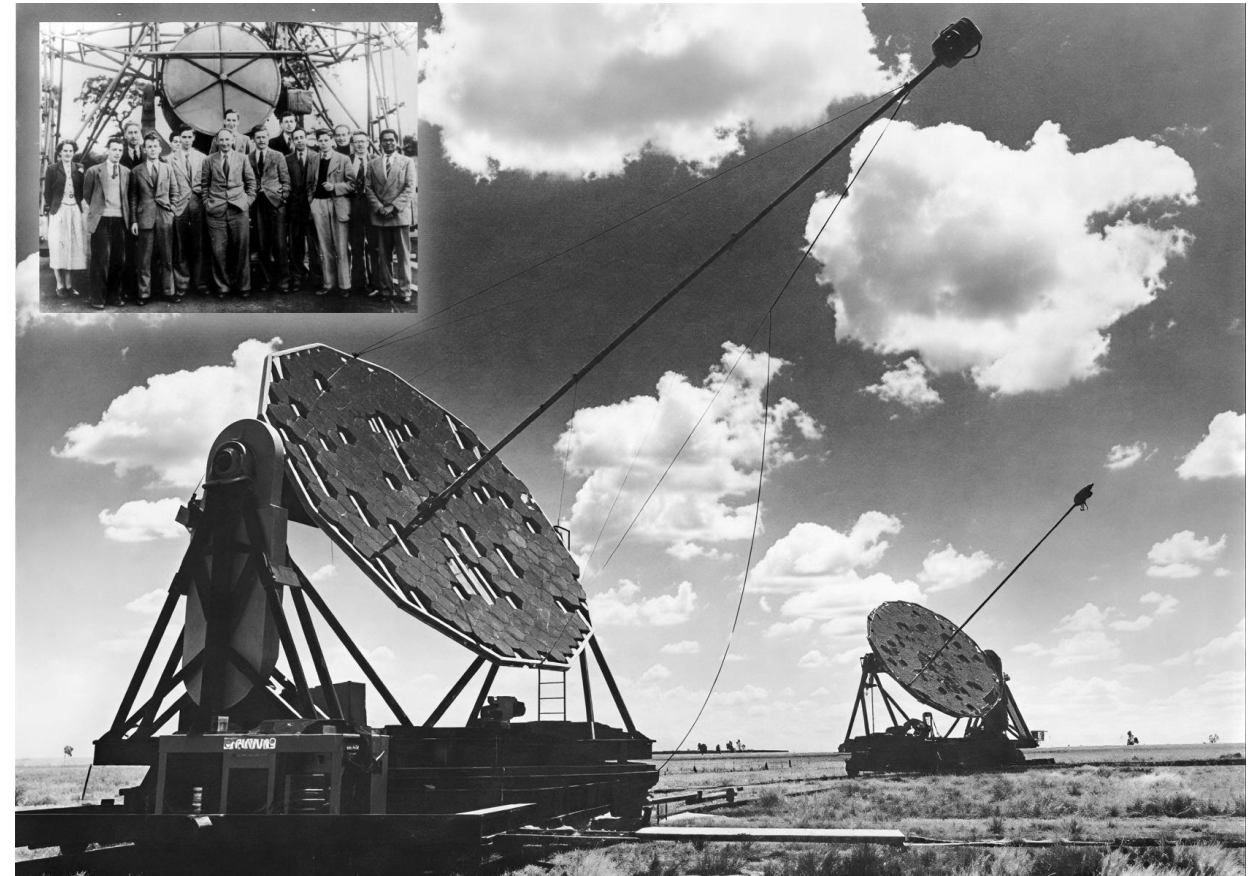
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$$\vdots$$

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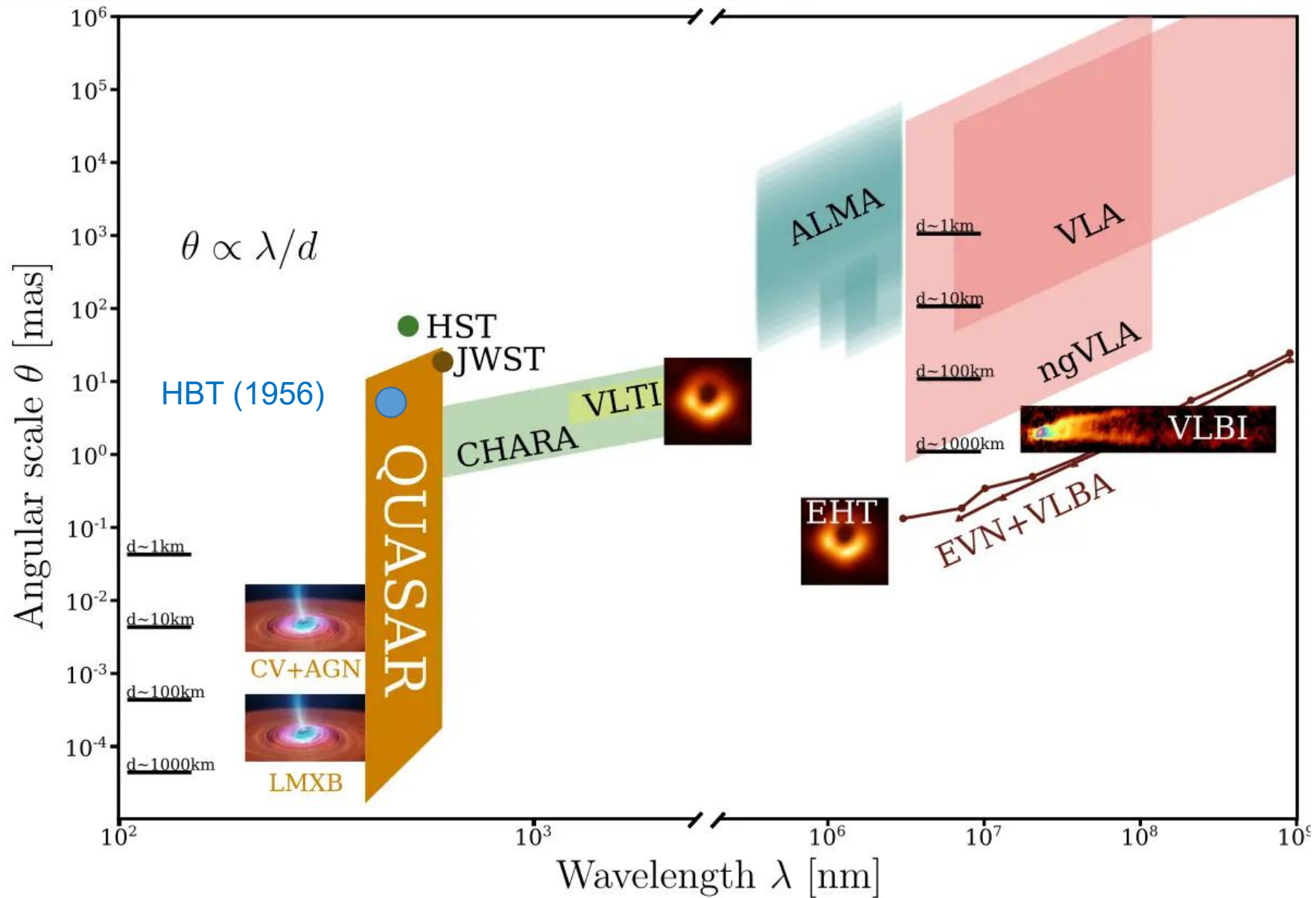


- First experimental proof : R.Hanbury Brown & R.Q.Twiss (1956)
- Measurement of angular diameter of Vega



QUASAR project

4

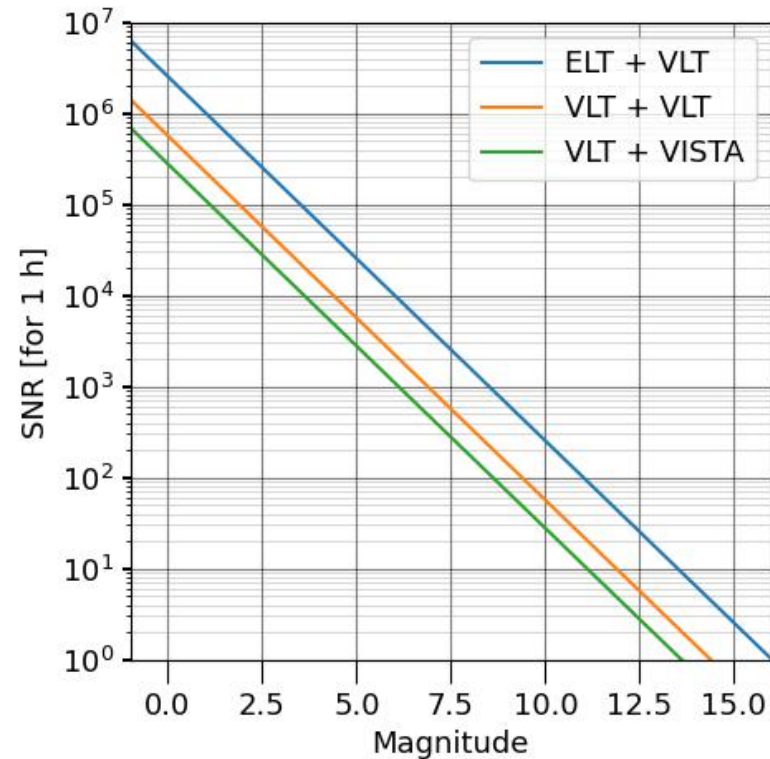
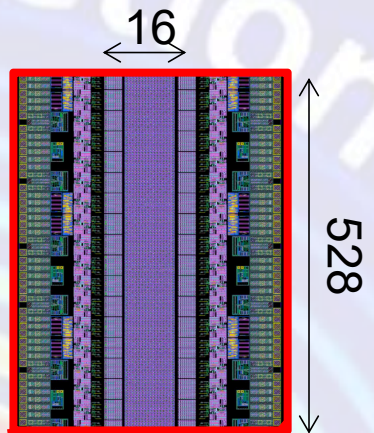


QUASAR key improvements

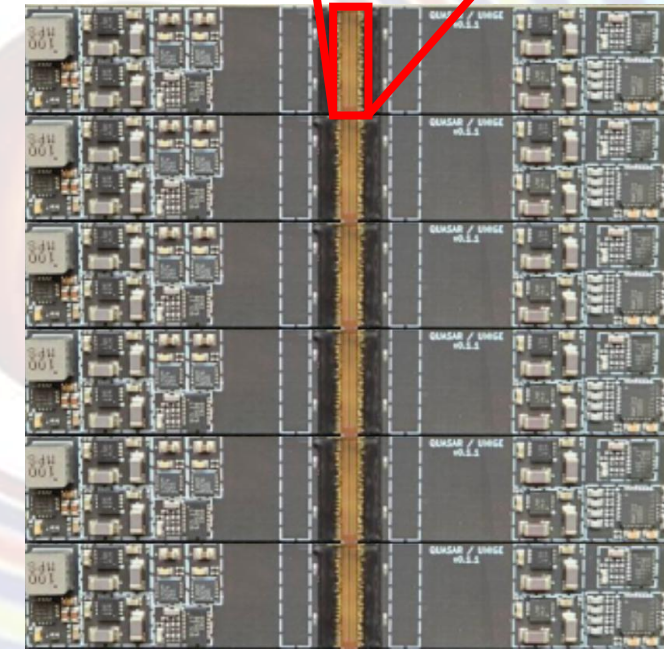
5

$$\text{SNR} \sim \sqrt{A_1 A_2} \sqrt{N_{\text{chan}}} \sqrt{\frac{T_O}{\sigma_t}}$$

- 60% fill factor
- ~3ps LSB
- 40% QE
- 41Gb/s



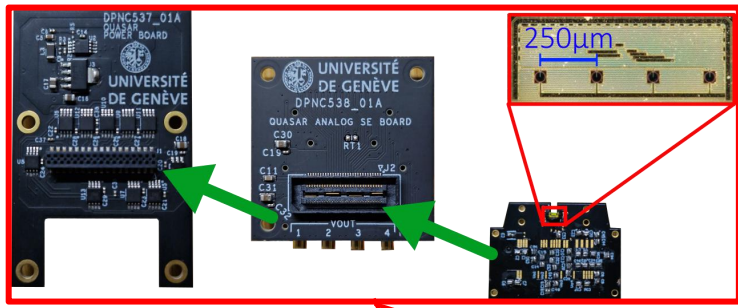
Targets	Size (θ)	Baseline ($\lambda/\tan \theta$)
0.01AU astrometry on α Centauri AB	2 μ deg	10 m
Decretion disks	μ deg	50m
Stars and stellar systems	μ deg to ndeg	50m-km
Micro to nano structures on the Sun	μ deg to ndeg	m-km
White dwarfs	8 ndeg	2.5 km
Accretion disks around white dwarfs	5 ndeg	2 km
Broad line region in quasars, H_0	1-100 ndeg	0.1-20 km
Tidal disruption events	1-100 ndeg	0.1-20 km
Stripped supernova @ 50Mpc	ndeg	20 km
Supermassive star	ndeg	20 km
Accretion disks of quasars	ndeg	20 km
Disk & jet connexions	ndeg	20 km
Binary black-holes	ndeg	20 km
Lensing (differential astrometry: light halo, DM)	0.3 ndeg	60 km
Black-hole horizons	100 pdeg	200 km
Accretion disks around neutron stars	30 pdeg	600 km
Pulsar magnetosphere	pdeg	10'000 km



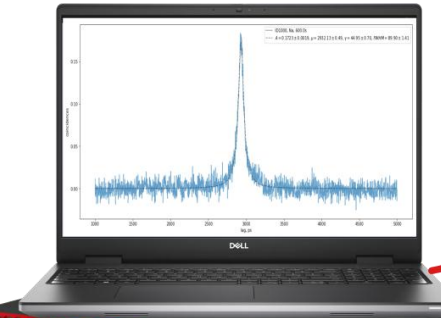
Received, tested now

Current setup

6



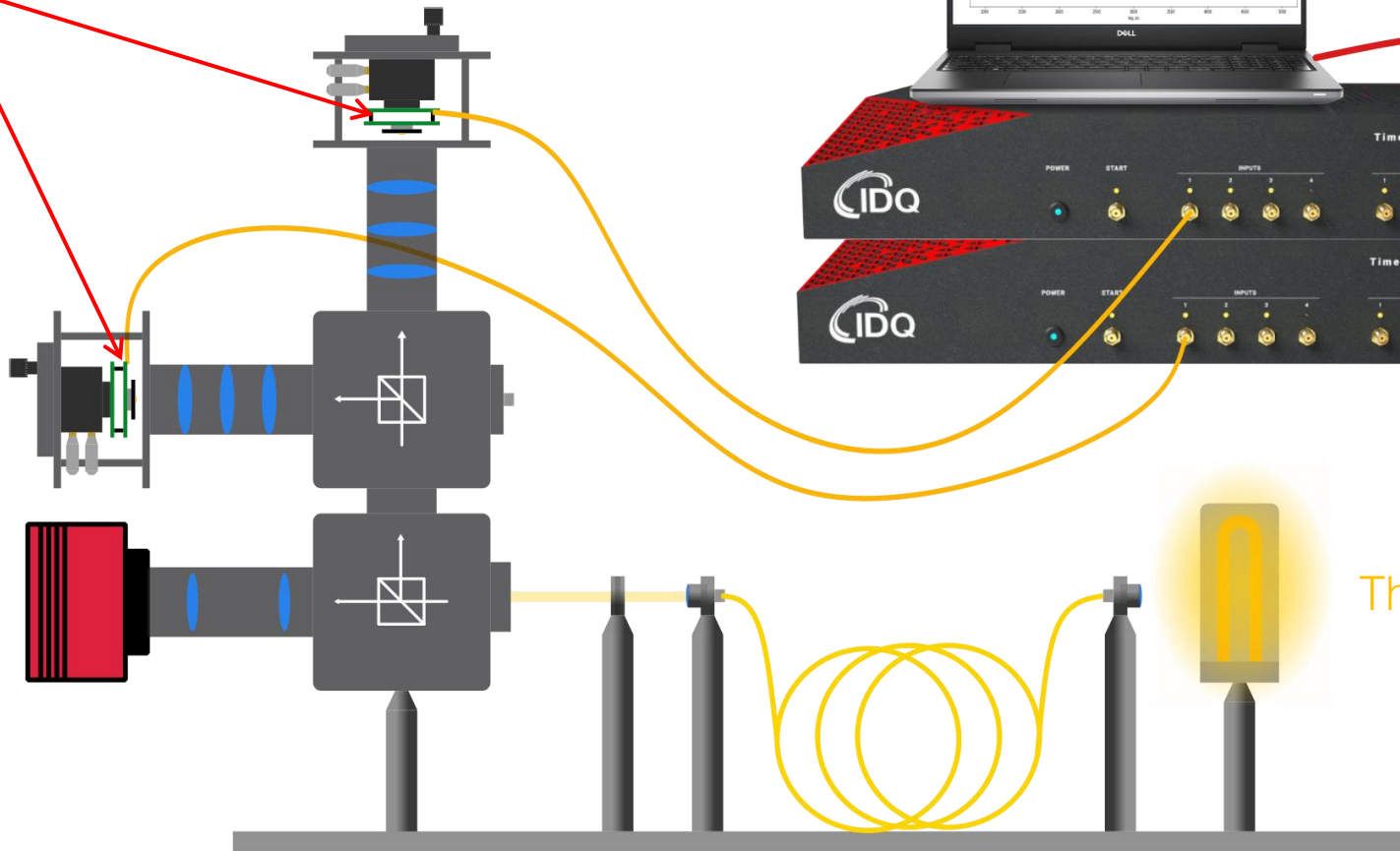
Offline correlations



ID1000 TDC
1ps digital res.
4ps jitter



Sync

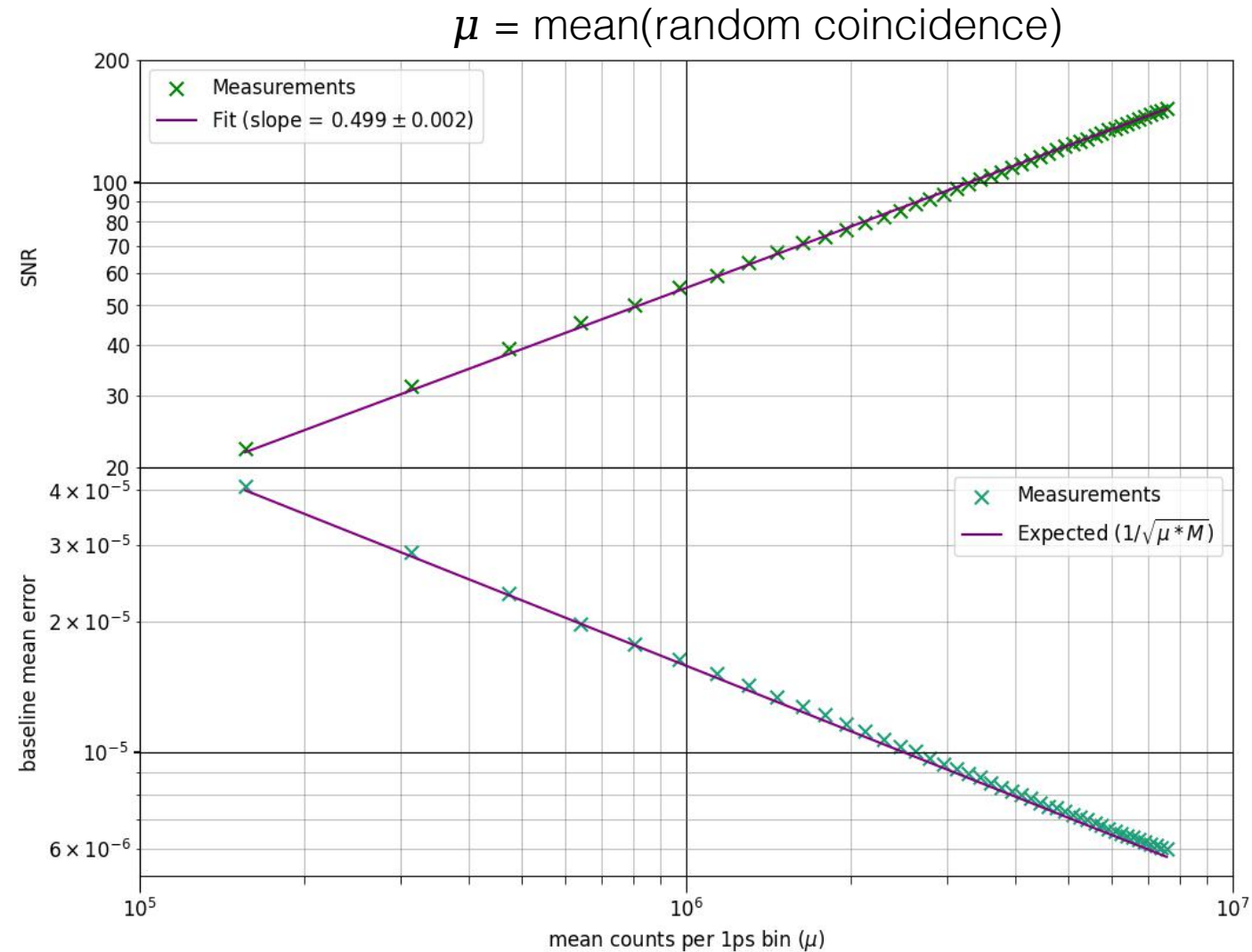


Thermal source + filter

Current status & results: Specs

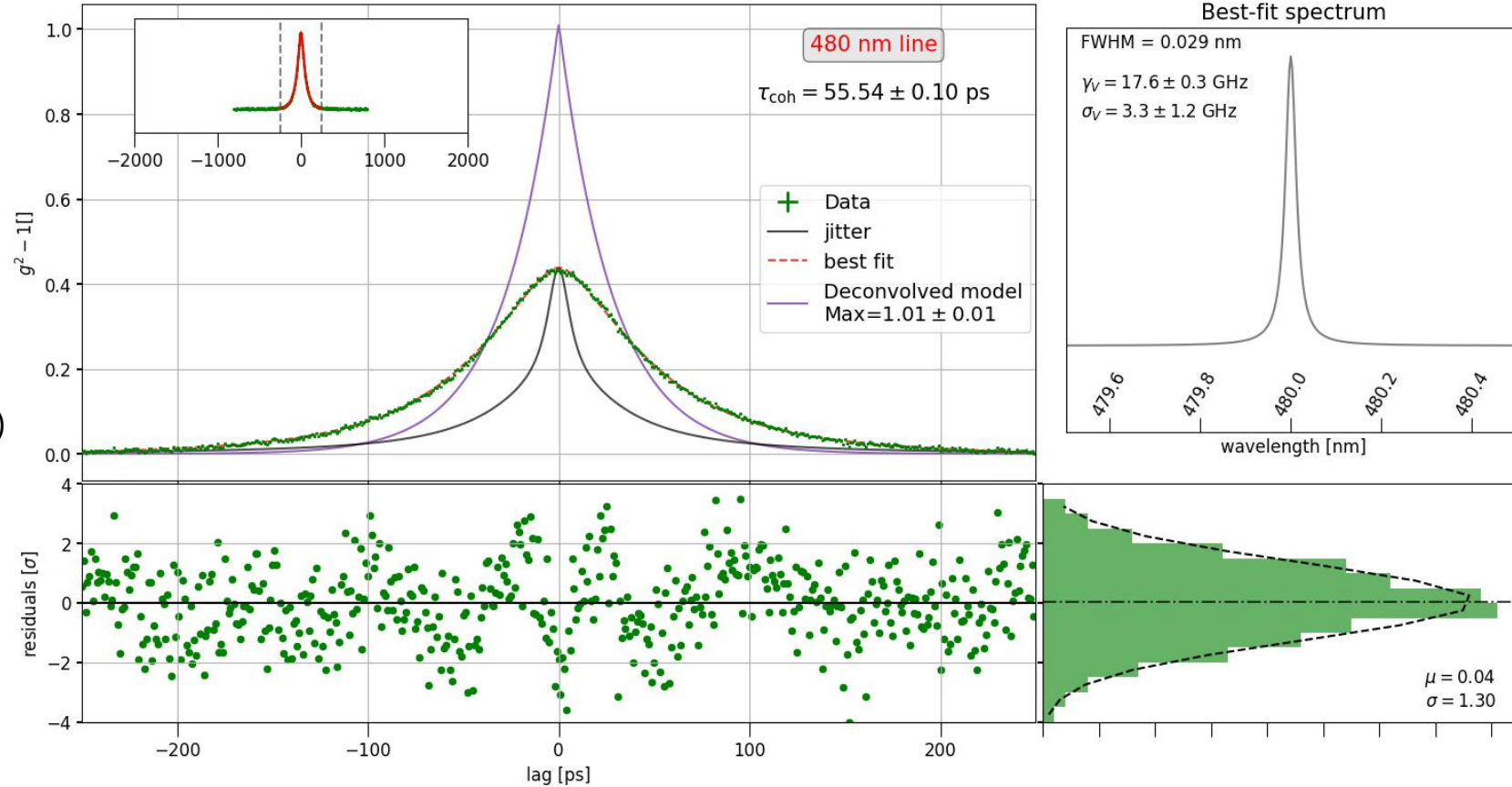
7

- Time resolution (full system) : **10.4ps RMS**
- SNR grows as $\sqrt{\mu}$
- No systematics up to 7e6 cts/ps bin



$$g^{(2)} = 1 + |\mathcal{F}(f)|^2 \cdot |V(d)|^2$$

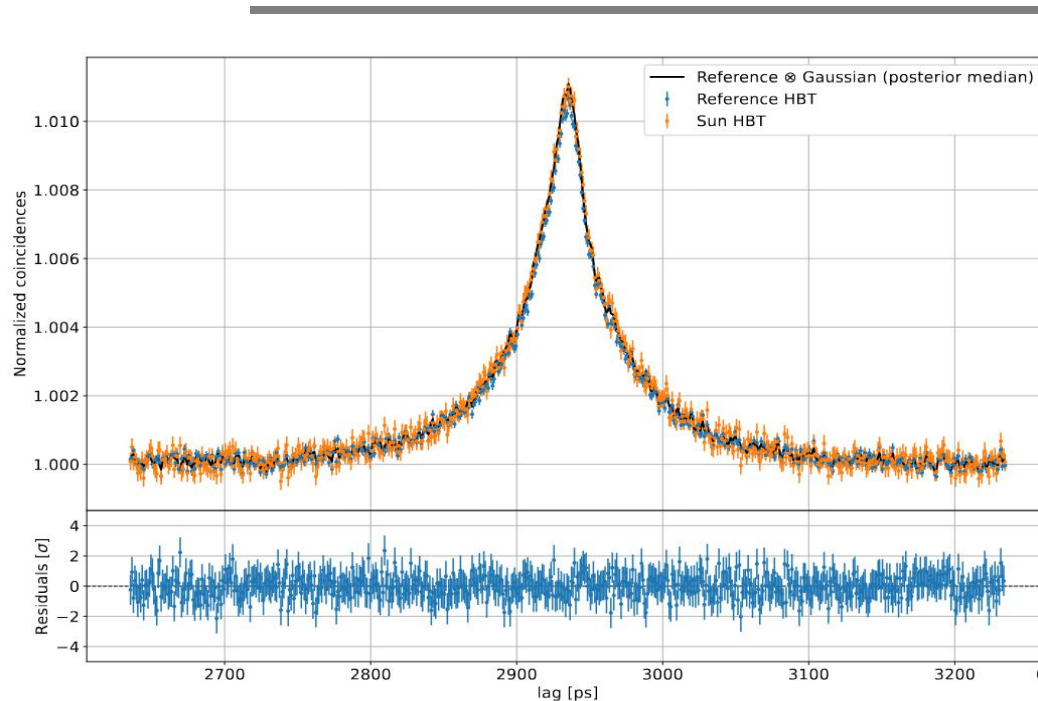
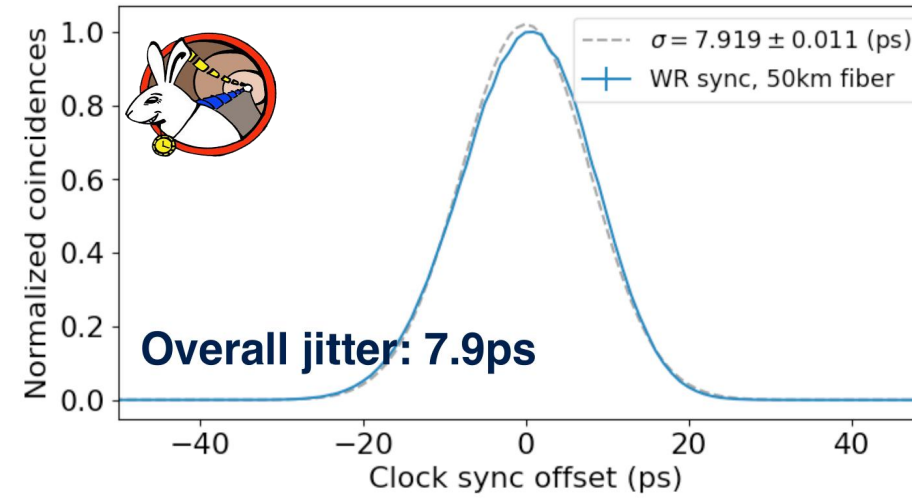
- Narrow line : higher coherence time
 - SM Amplitude recovered
 - Linewidth reconstructed (30pm)
 - Residuals excess (lineshape, lamp)



Current status & results: Timing

9

- Spool in thermal chamber :0.2°C/h cycles
- Monitor master-slave 10MHz sync
 - 7.9ps time sync jitter

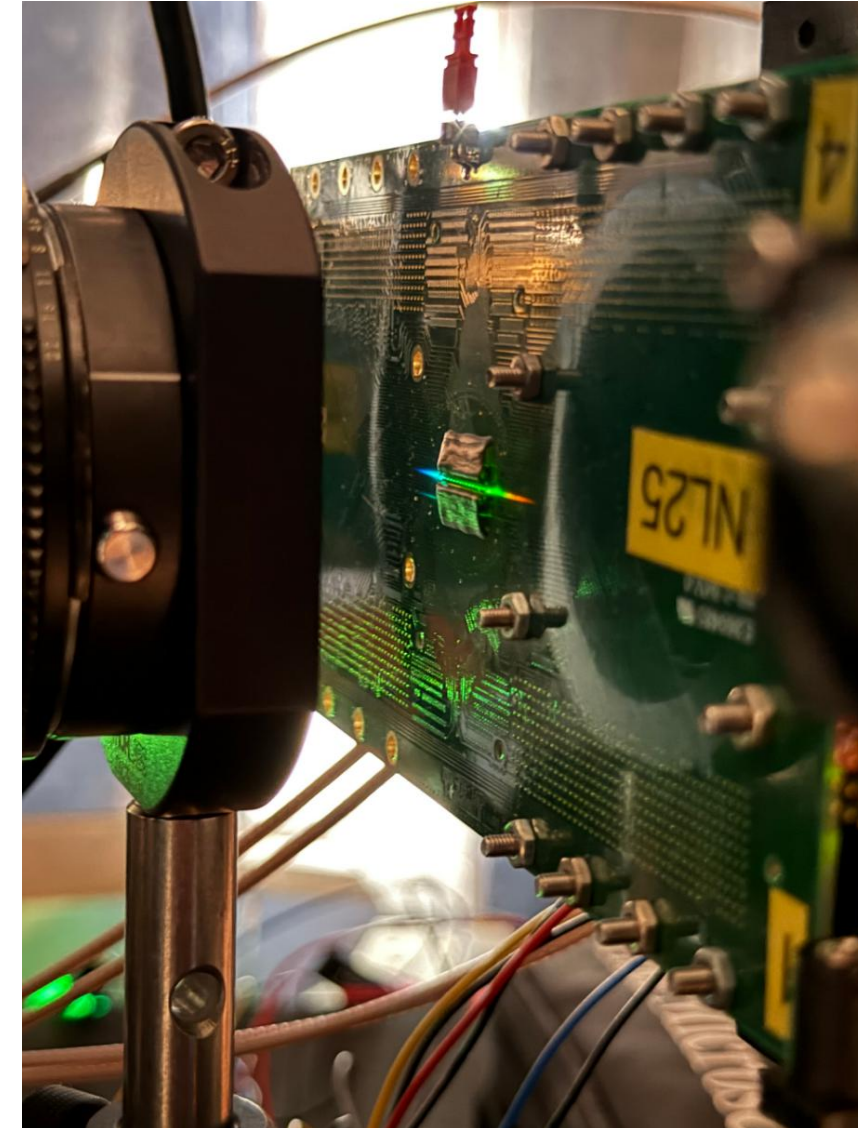
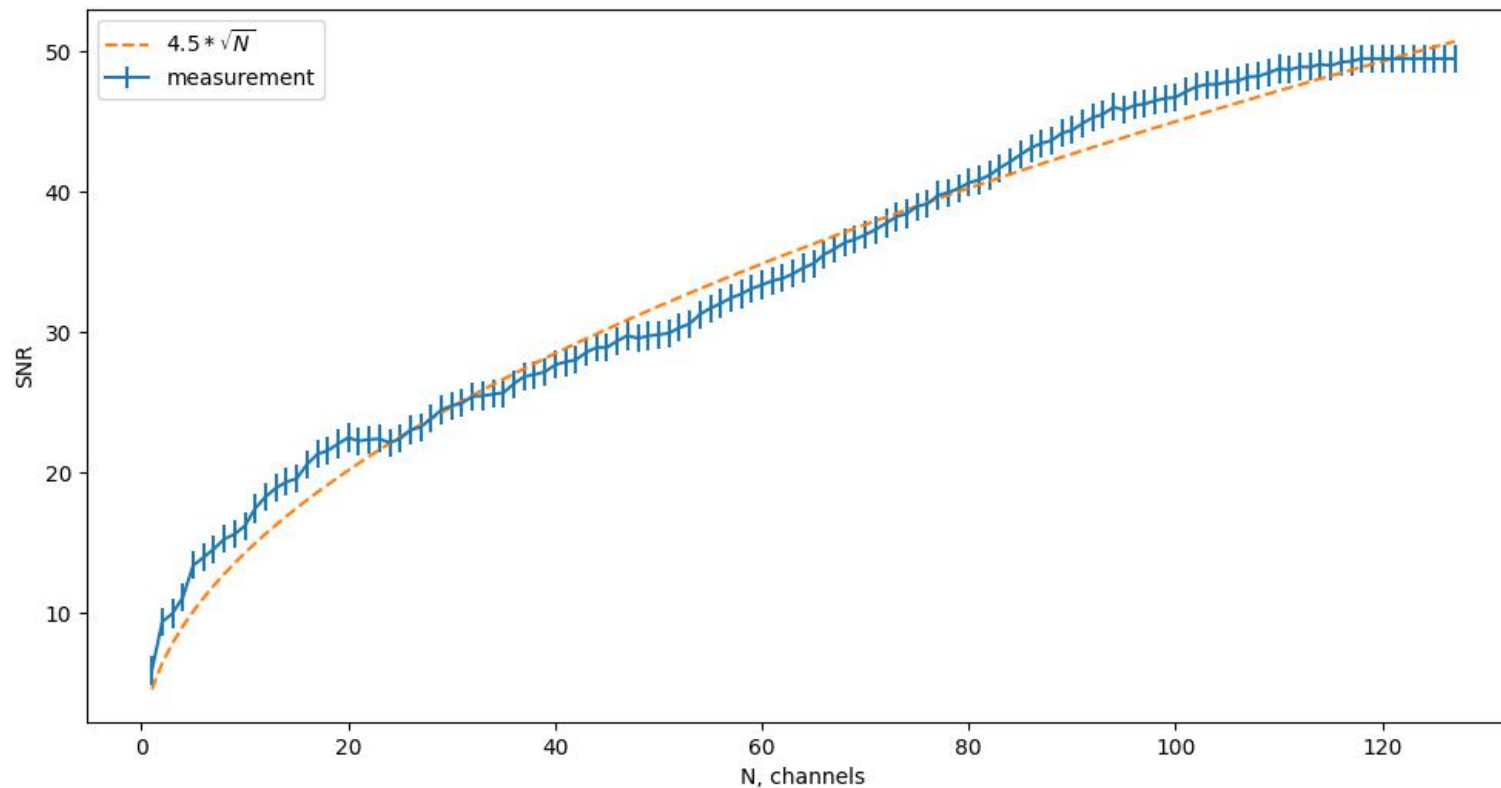


- Models : 30ps broadening
- Measure reference (led) VS sun
 - Measured broadening : 0.65 ps

Current status & results: Multiplexing

10

- Disperse the light on 2x128 channels
 - Spectrometer made in Geneva
 - Older version of detectors (~50ps)



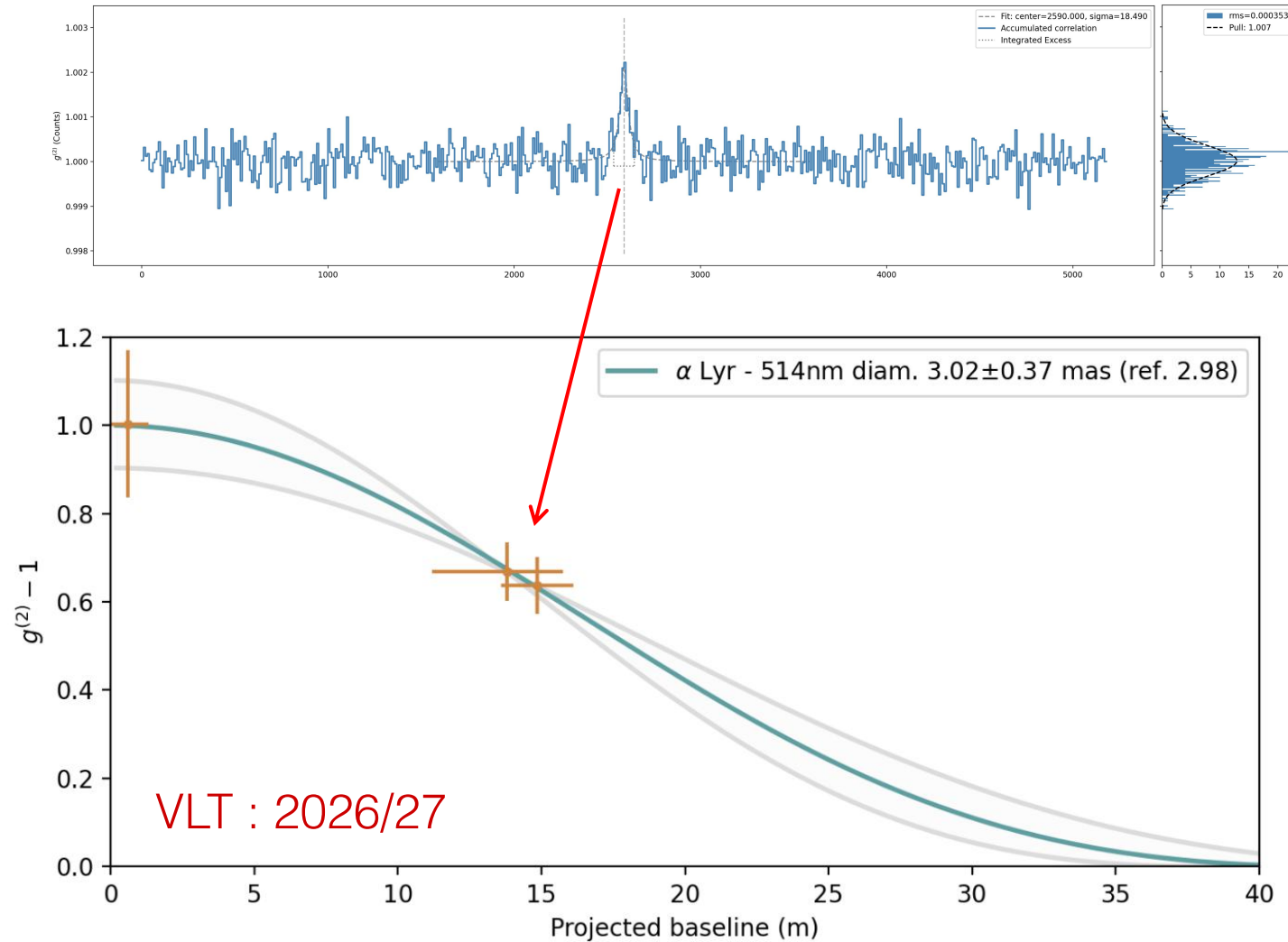
Current status & results: Vega

11

- Skinakas (Crete) + Calern (France)
- Bad weather for the full week
 - 4h of usable data



C2PU, Calern

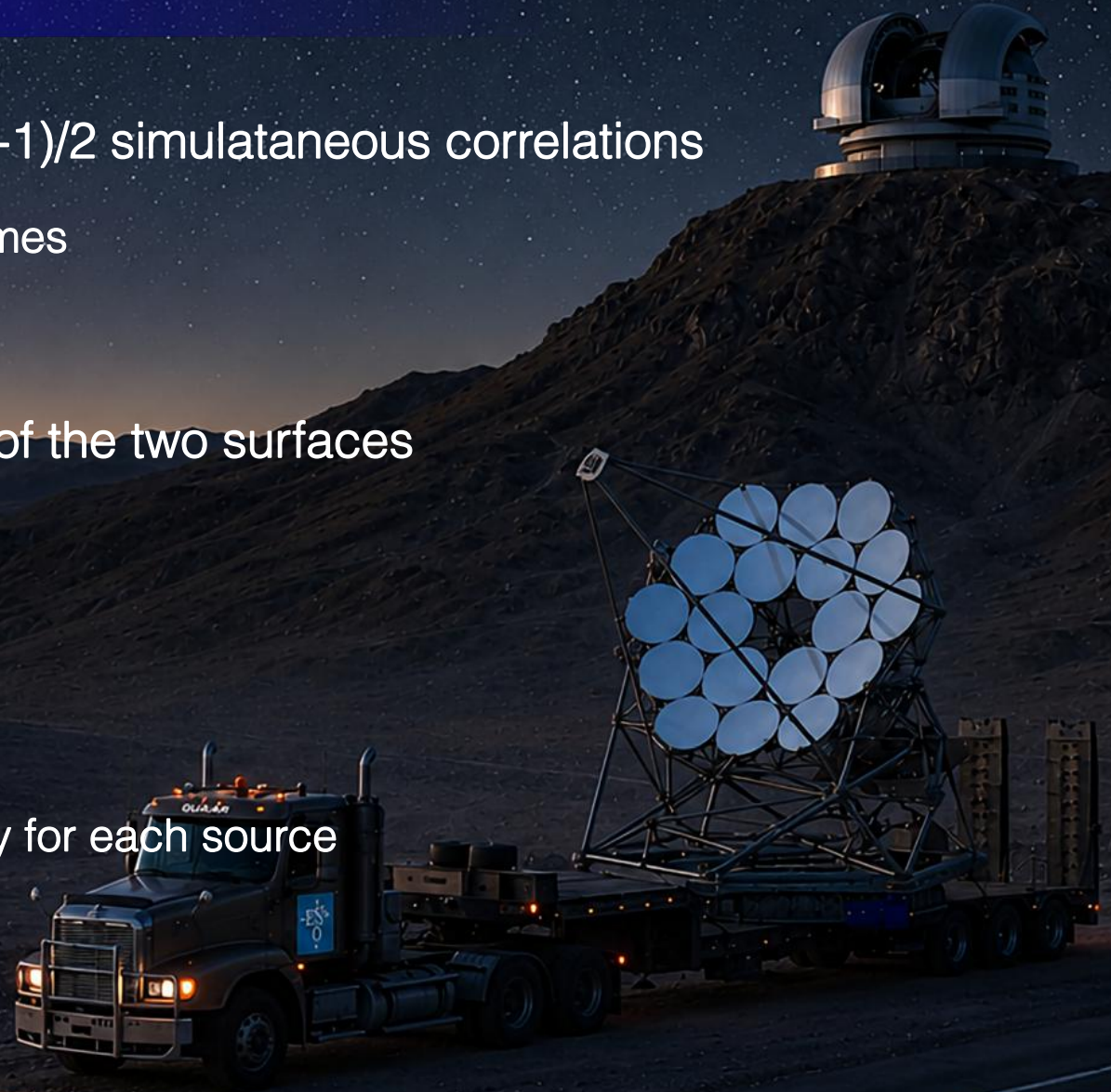




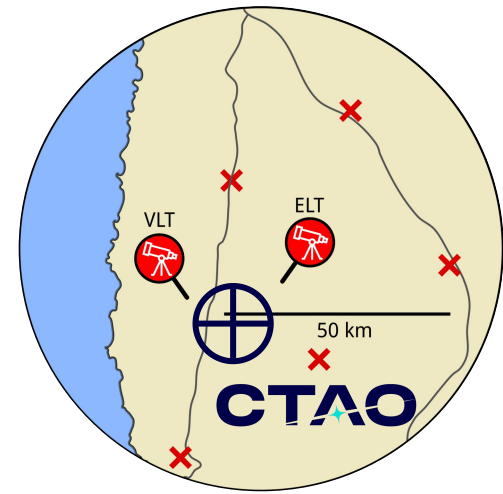
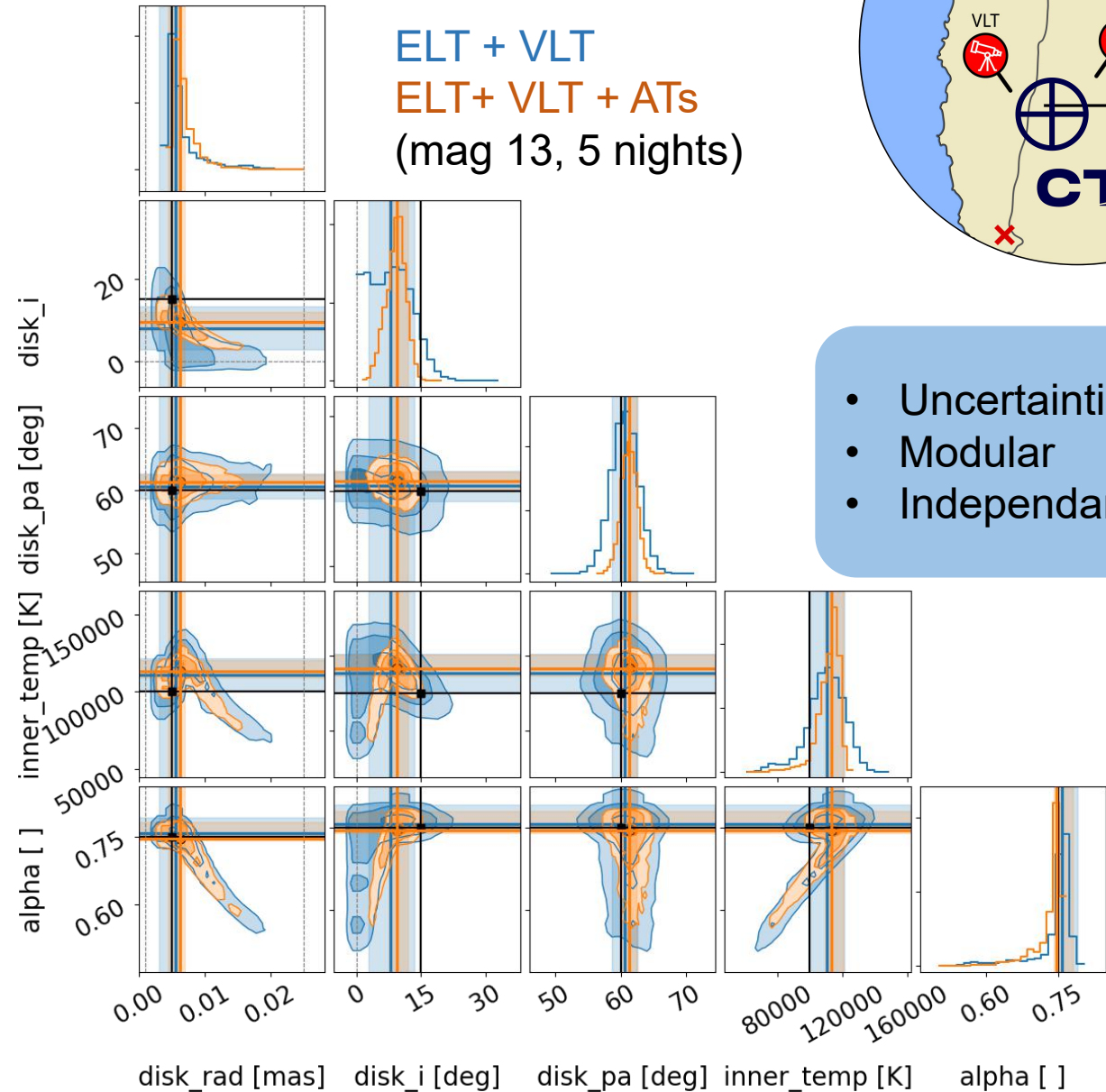
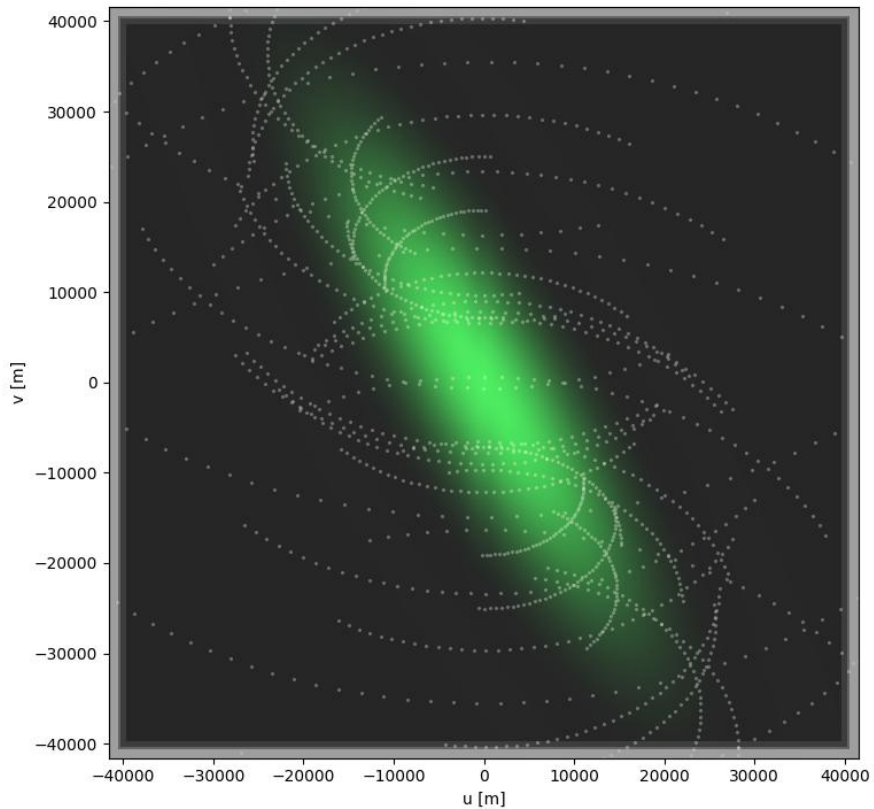
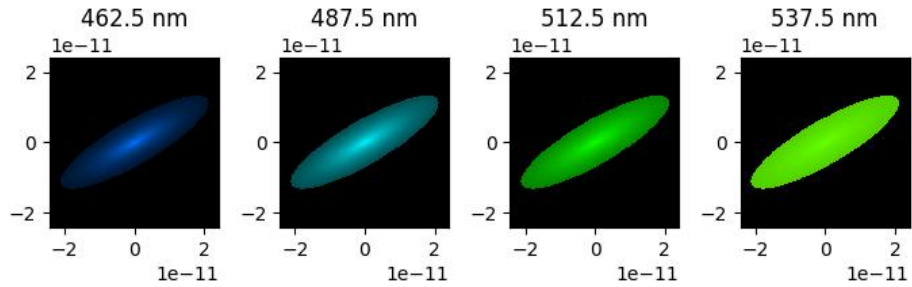
Auxiliary Telescopes (ATs)

12

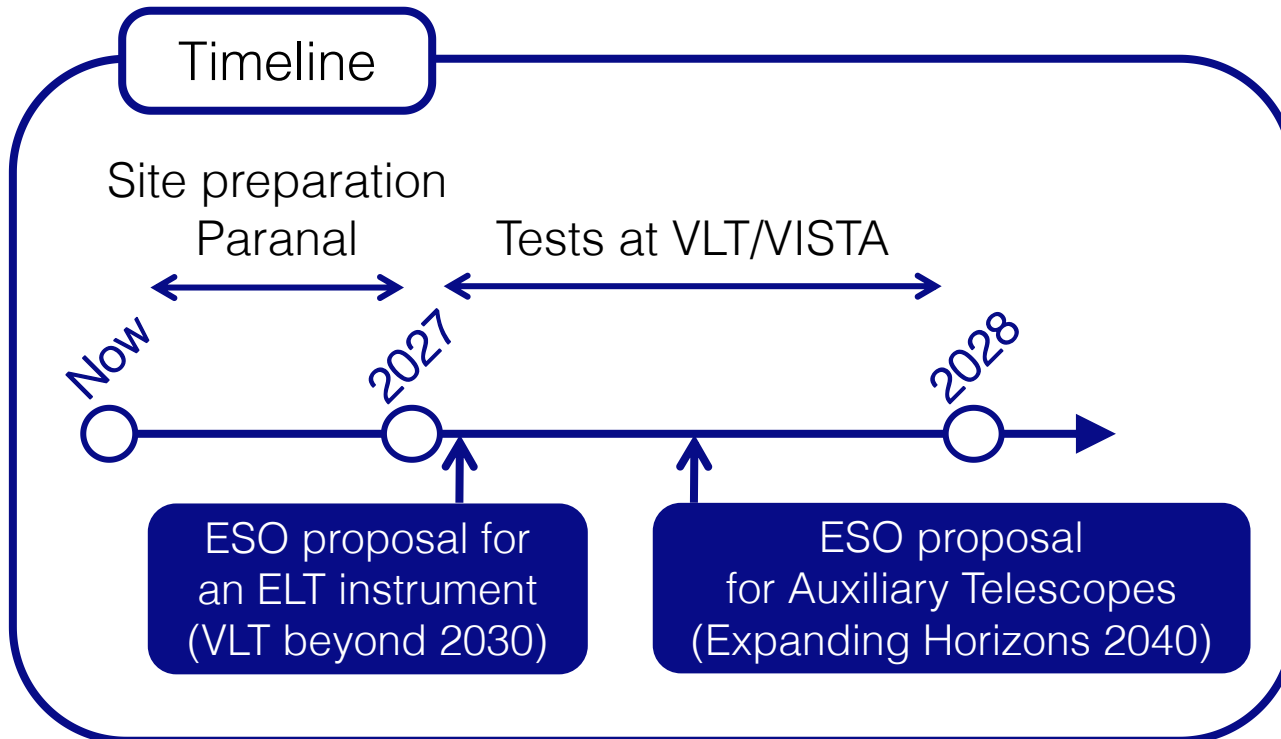
- Intensity interferometry allows to make $N(N-1)/2$ simultaneous correlations
 - Photon timestamps can be used multiple times
 - The UV coverage increases
- The SNR depends on the geometric mean of the two surfaces
 - Limits the SNR loss of small telescopes
- Small telescopes have 2 main benefits
 - Cost : Target $< 1\text{M}\$/\text{unit}$
 - Modularity : Can be moved & placed ideally for each source



* AI generated



- QUASAR aims at reaching microarcsecond resolution in the visible
 - Chip is being tested
 - First results expected next year
 - Low cost & modular auxiliary telescopes can improve the uncertainties a lot



Richard Feynman jumped up and said 'It can't work!'

In his inimitable style, Hanbury responded, 'Yes, I know. We were told so. But we built it anyway, and it did work'.

Late that night, Feynman phoned and woke Hanbury up to say 'you are right'. He also wrote a letter in which he magnanimously admitted his mistake and acknowledged the importance of this phenomenon that, at first sight, appears counter-intuitive, even to quantum theorists.