

Zjazd CAMK 2026

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GW Group @ CAMK



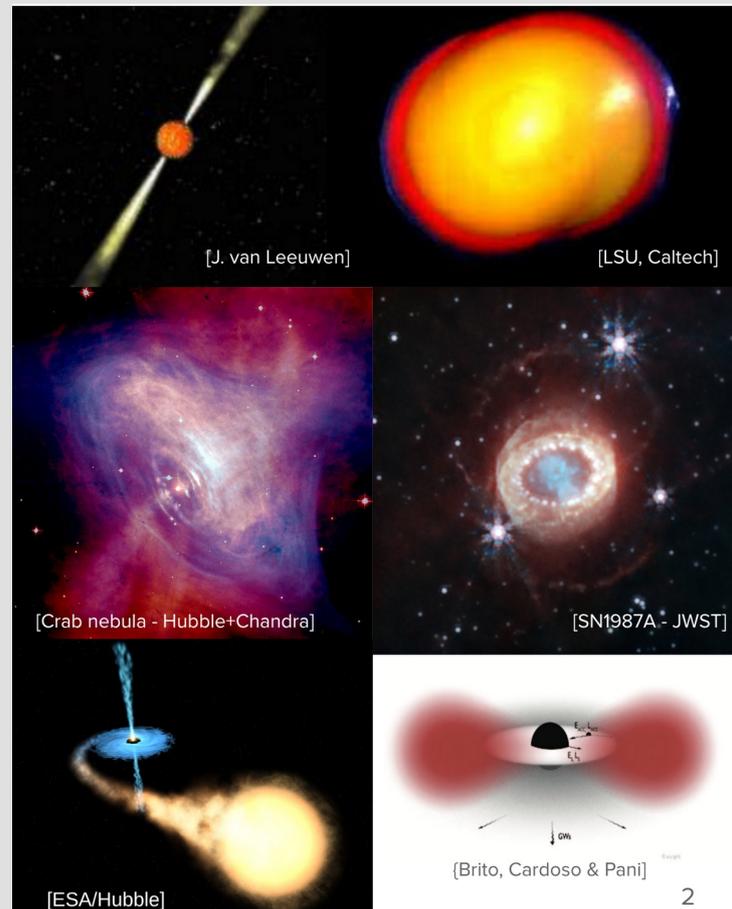
CW group science aims

Search for continuous (long-duration) GW signals;
fiducial source is an asymmetric rotating neutron star.

- **Targeted searches:** known frequency evolution, known sky location (e.g., pulsars with “mountains”)
- **Directed searches:** known location (e.g., supernova remnants, Galactic Center, LMXBs), unknown frequency evolution, (some) unknown binary parameters, post-glitch (r-modes), post-merger...
- **All-sky searches:** unknown location and frequency evolution, binary parameters (model agnostic)

But also

- **“New physics”:** black hole boson clouds, inspirals of light primordial black holes, dark matter (including direct detection)



TDFstat pipeline

How to detect continuous GW signal which is orders of magnitude weaker than the noise?



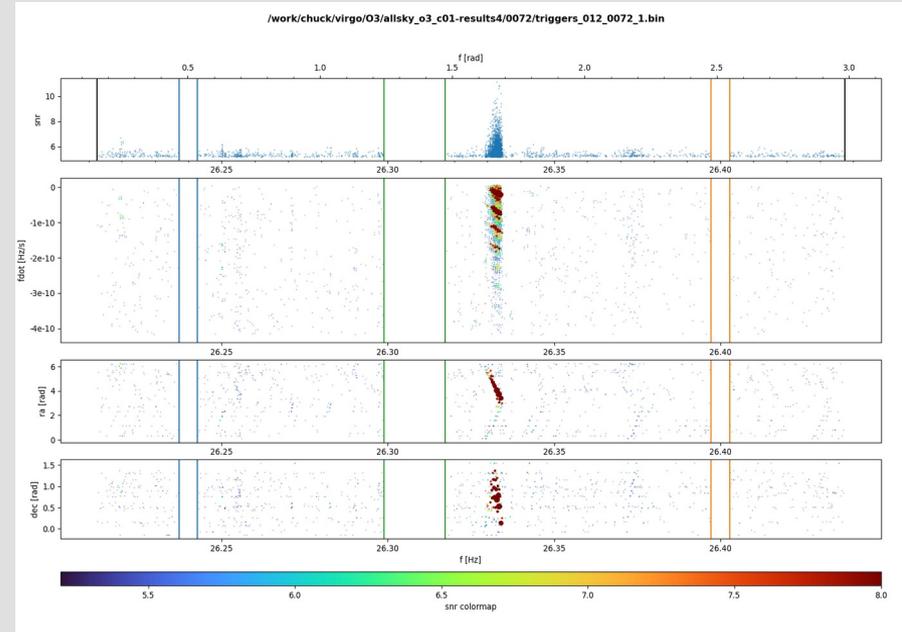
Likelihood of matching model with the data



F-statistic (Jaranowski, Królak, Schutz 1998)



All-sky, blind search pipeline (f, f', α, δ):
<https://github.com/Polgraw/TDFstat>



TDFstat pipeline

TDFstat development

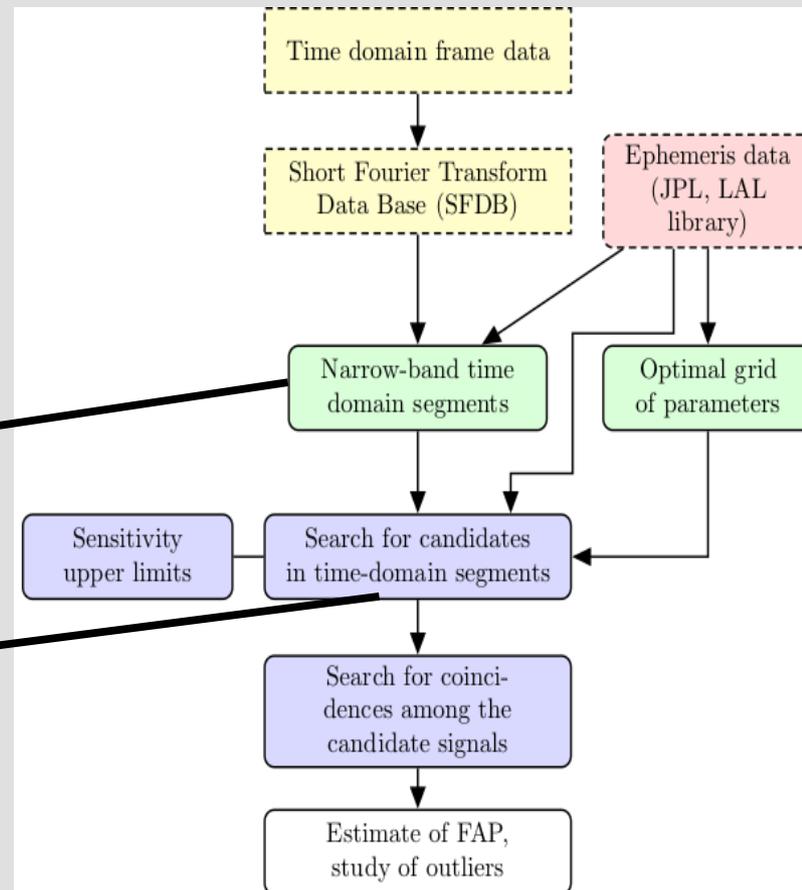
Drivers: LVK deadlines, various searches (spotlight/directed),
group activities: ML analysis, GW lensing

Directions: optimization, semi-automatic analysis, new applications

- Input data generation
 - extract_band (HDF files with STS data)
 - genseg (rewritten from scratch)

- Search
 - HDF output
 - Fractional grid with ranges

- Scripting in Julia



Virgo-PL: Computing

- IGWN Computing Tier-2 node at Cyfronet
 - 5 mln PLN for hardware + 1 FTE
 - >6000 CPU cores, >2 PB storage
 - Tier-2 site for offline processing, with full, local data cache
 - Part of IGWN computing network
 - Large contribution to the european resources in LVK
- **Support for Polgraw (archive, semi-automatic pipelines)**