

High pressure operation of MPGD in CAPA

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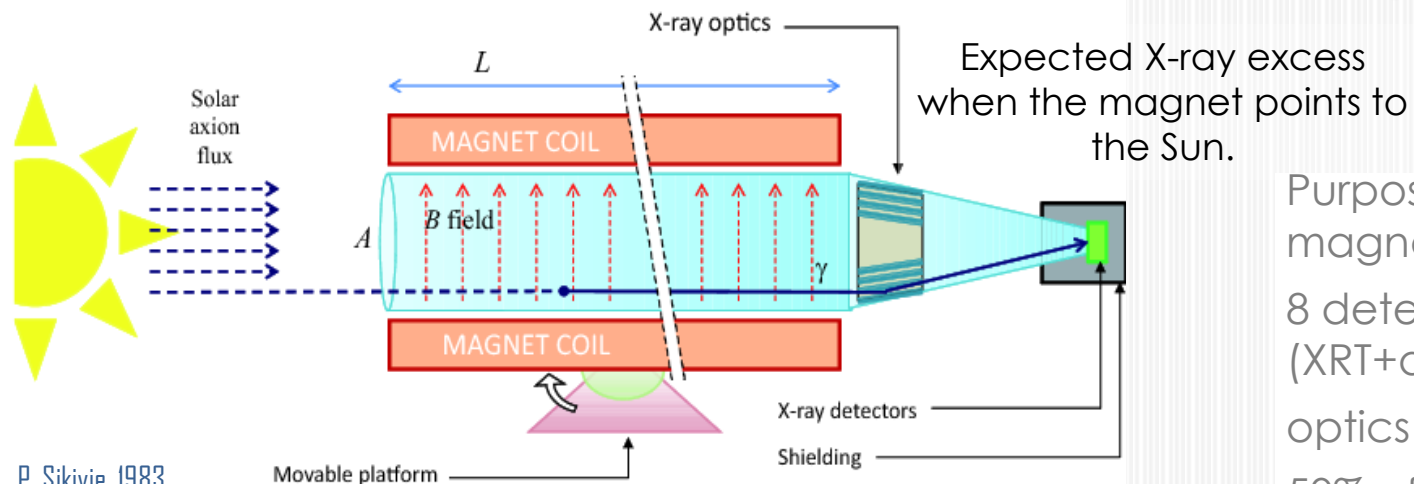
CAPA: Centre for Astroparticle and High Energy Physics

Institute

- Research Centre of the University of Zaragoza
- Research Lines:
 - Dark matter, axion physics and neutrino physics
 - Low background techniques and detector development
 - Lattice gauge theory and field theory applications
 - Standard Model extensions and quantum gravity
 - Observational astrophysics and cosmology
- Close Collaboration (experimental) with:
 - CERN, CEA/Saclay, DESY, LSC, CEFA, ITA, LLNL, NASA, ESA
 - RD-51. DRD1
- Expertise in
 - low-background techniques
 - Gaseous TPC
 - Micromegas detectors
 - **Software related to simulations and data analysis: REST-for-Physics** (<https://rest-for-physics.github.io/>)
- Applications (mainly) to rare-event searches
...and general R&D in gas readouts
 - Active surface, resistivity, segmentation,...

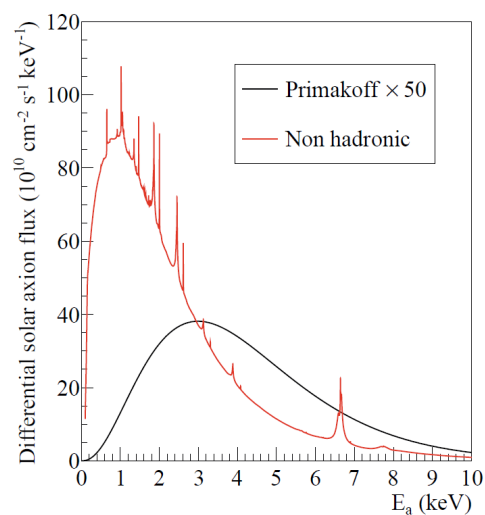


IAXO and BabyIAXO



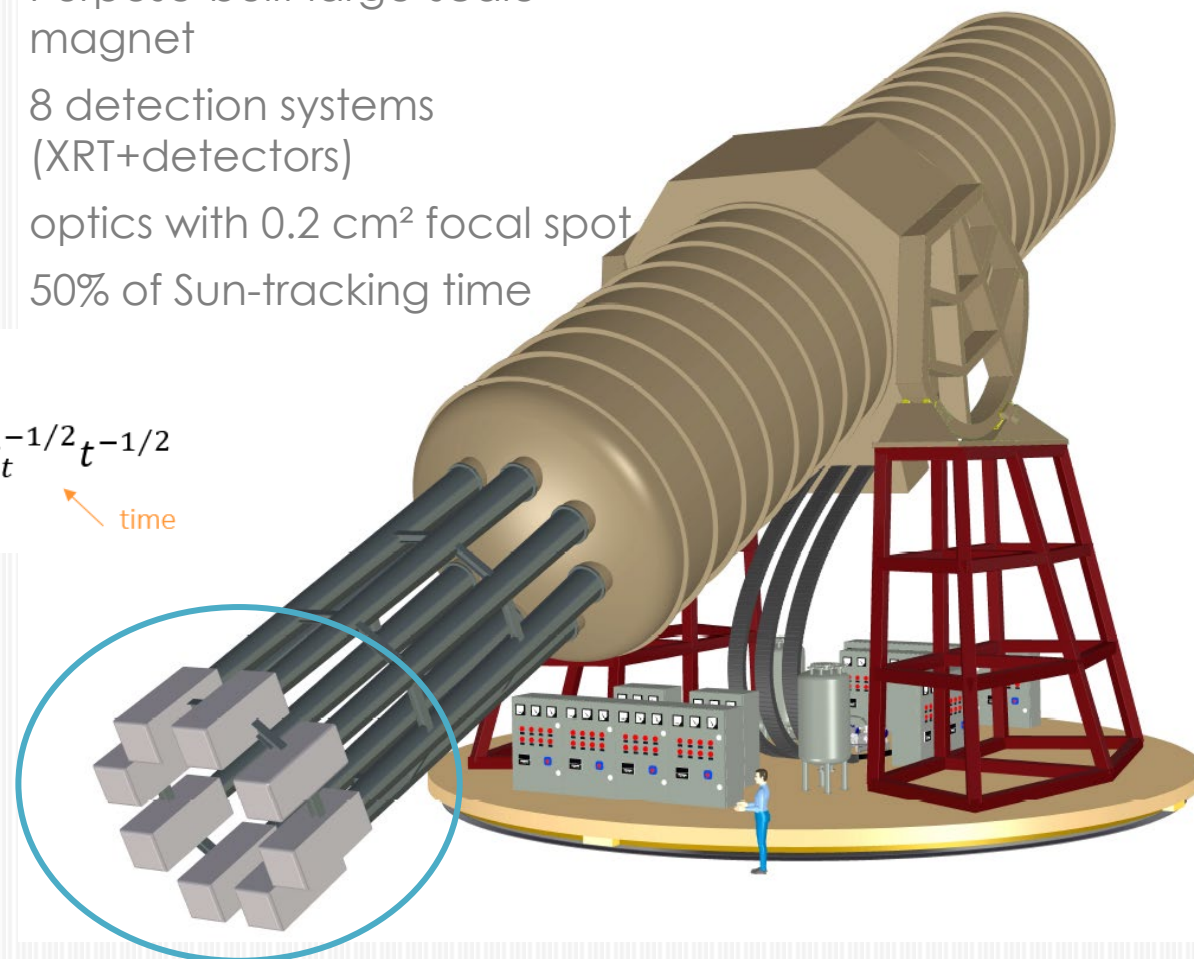
P. Sikivie, 1983
K. van Bibber, G. Raffelt et al. (1989)

Purpose-built large-scale magnet
8 detection systems (XRT+detectors)
optics with 0.2 cm² focal spot
50% of Sun-tracking time



$$g_{a\gamma}^4 \sim \overset{\text{magnet}}{B^2 L^2 A} \quad \overset{\text{optics}}{\varepsilon_o \alpha^{-1/2}} \quad \overset{\text{detectors}}{\varepsilon_d b^{-1/2}} \quad \overset{\text{time}}{\varepsilon_t^{-1/2} t^{-1/2}}$$

Figure of Merit (FoM)

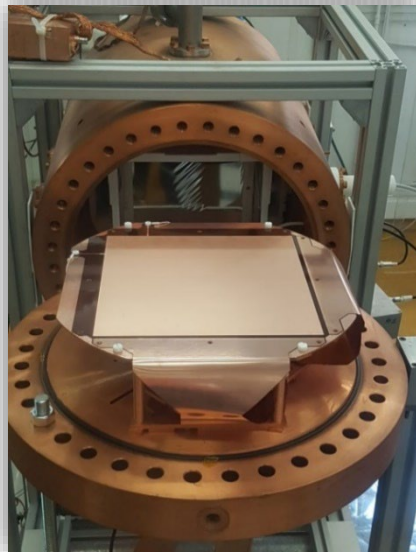
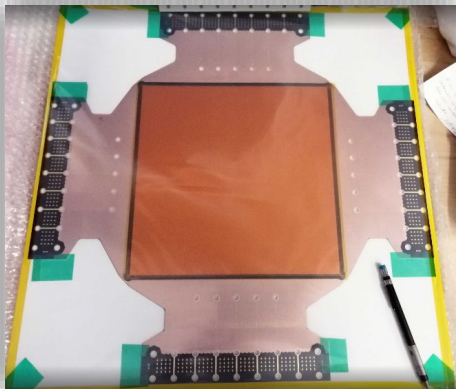
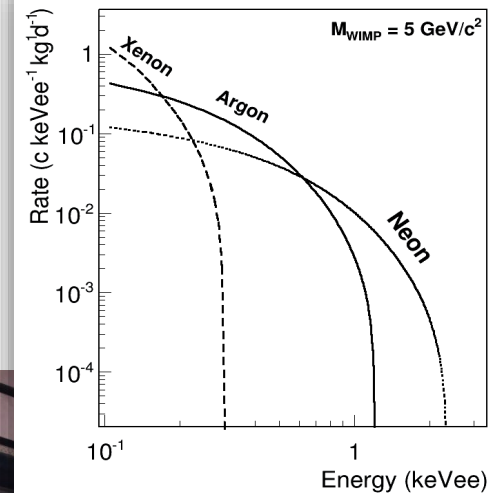


TREX-DM (TPC for Rare Event eXperiments-Dark Matter)

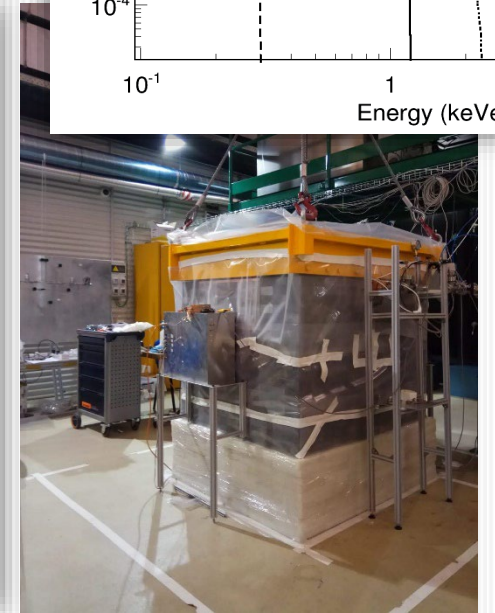
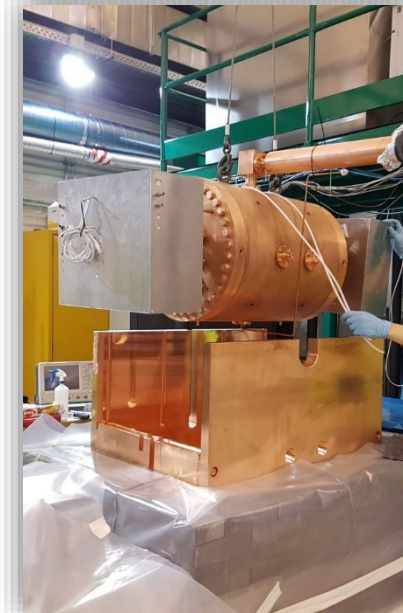
- A Micromegas TPC for **light WIMPs** at the LSC
- ~20 l of pressurized gas (~0.16 kg Ne at 10 bar)
- microbulk Micromegas and AGET-based electronics.
- **Goals:** low energy threshold (< 1 keV)
low background level (~1/keV/kg/day)
- NOT focused in directionality → operation at high P

Potential to be sensitive
to low-mass WIMPs
(0.1-10 GeV/c²) beyond
current bounds

The lighter the target, the
more sensitive to low mass



H=0.5 m, D=0.5 m
Central cathode



Experiments and R&D

Requirements/Similarities

- Relatively high gain
 - Energy range of interest: 0.1-20keV
- Low background
 - Radioactivity issues
 - Background discrimination techniques
 - Good detector response characterization
 - Event topology (appropriately segmented anode, software efforts,...)
- Low energy threshold
 - Sensitivity enhancement (appropriate trigger, DAQ electronics, ...)

Differences

- Active area
- Operating pressure
 - High pressure (up to 10bar) to accumulate target mass
 - Energy threshold differences
- Operating gas optimization
 - Target differs because of Physics
 - Low-mass WIMPs: Ne, Ar with 10%Iso
 - Solar axions: Xe-based mixture

Low-Background levels

...

Low-backgrounds...

- Background components of our experiments:

- Cosmic rays : γ , neutrons
- Radiation and particles from environment
 - Radioactivity in materials (primordial, cosmogenic, anthropogenic):
 - ^{232}Th , ^{238}U , ^{40}K , ^{60}Co , ^{210}Pb , ... ^{39}Ar !



- **Radon ^{222}Rn**

α, γ + Outgassing, daughters plate-out on surfaces

- Neutrons:

- radiogenic: fission, (α, n) reactions in rock,
- cosmogenic: induced by muons

(n, γ) capture; inelastic scattering $\rightarrow \gamma, \dots$
elastic scattering $\rightarrow$ WIMP-like recoils!

Mitigation:

Underground operation
Passive and active shielding
Radiopurity control
Background discrimination techniques

➤ Attempts of creating single public repositories of radiopurity measurements

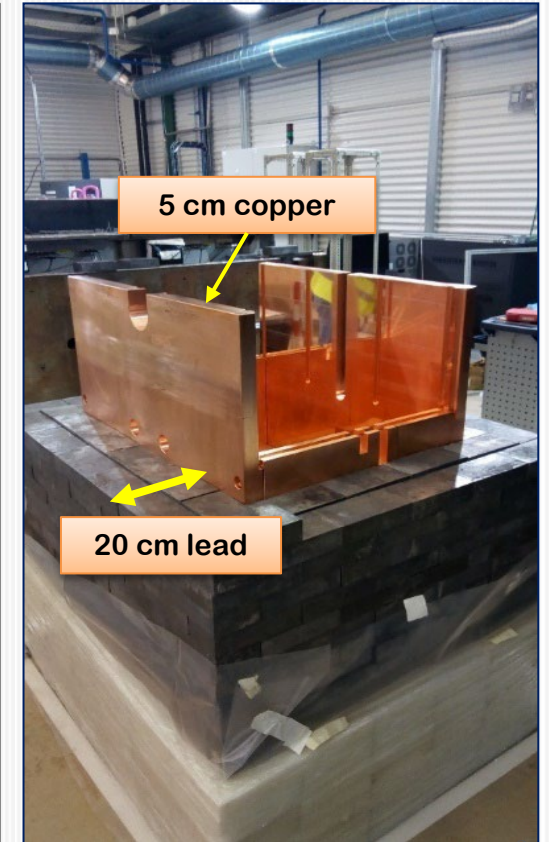
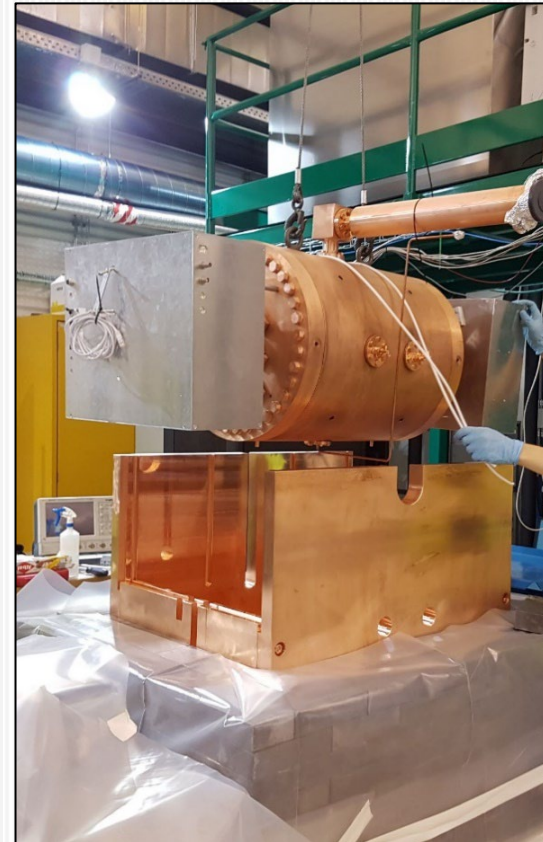
Radiopurity.org: <https://www.radiopurity.org>

Adapted from : S. Cebrian, *Symposium on Large TPCs*, 13th December 2023

J.C. Loach et al., *Nucl. Instrum. Meth. A* 839 (2016)

Background levels: Shielding

- 5 cm copper + 20 cm lead
- DAQ electronics outside the shielding
- N₂ or Rn-free air is flushed into the plastic cover
- Neutron shielding in place (80%)
polyethylene ceiling + water tanks



Radiopure electronics

- Usually readout electronics outside the shielding
 - Longer cables, higher noise/threshold
- Development of radiopure FE boards
- Complemented with non-radiopure BE

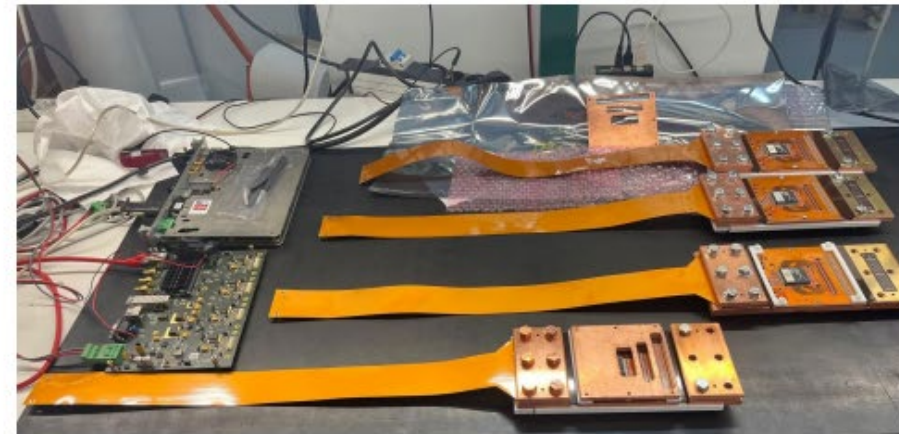
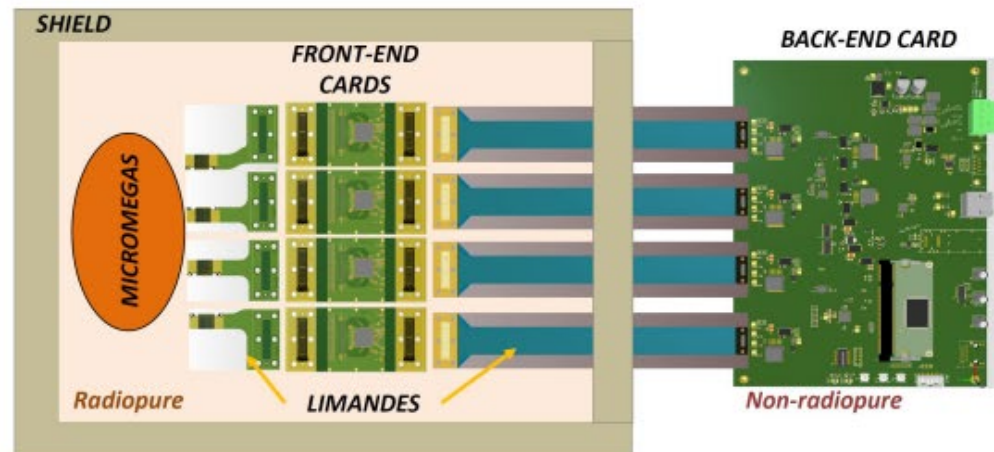


Institut de Ciències del Cosmos
UNIVERSITAT DE BARCELONA

EXCELENCIA
MARIA
DE MAEZTU
2020-2023



Front End, 4 cards with STAGE and flat cables to connect to Back End



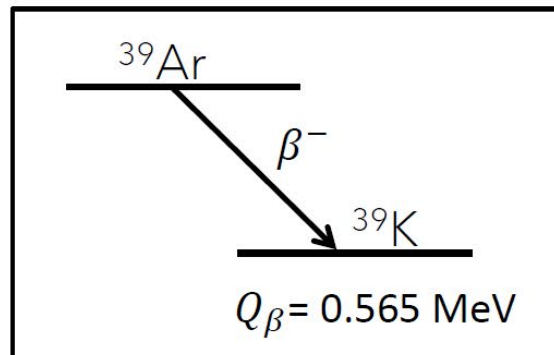
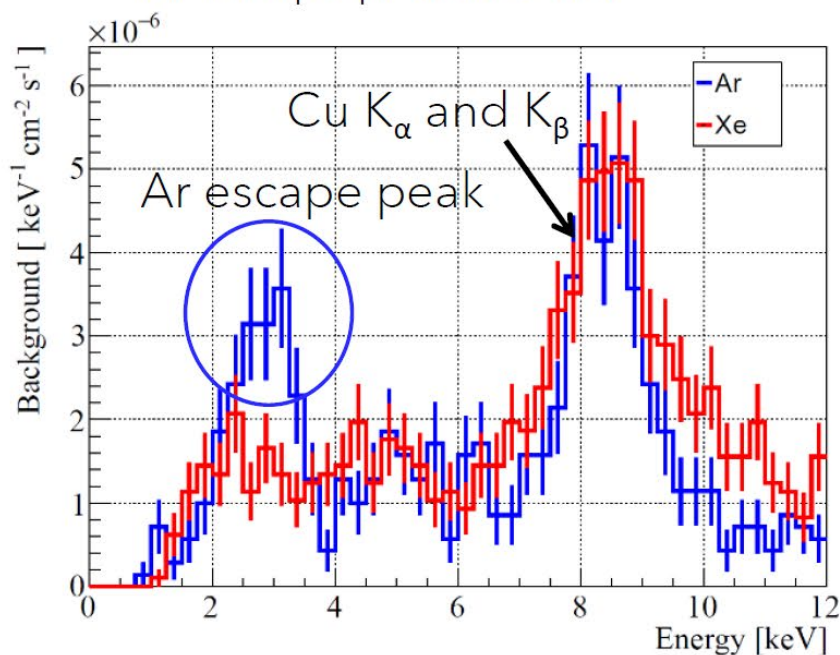
Being installed in the lab prototype during this week

^{39}Ar and Rn: (Baby)IAXO prototype in LSC

Goals

- Study of intrinsic background
- Optimization of the performance with different gas mixtures:
Ar+Isobutane and Xe+Ne+Isobutane
 - From simulations: intrinsic background in Ar limited by ^{39}Ar
 - Ar escape peak at 3 keV

Requires recirculation



Slide from M. Jiménez

[Altenmüller, K., et al. "New upper limit on the axion-photon coupling with an extended CAST run with a Xe-based micromegas detector." *Physical Review Letters* 133.22 \(2024\): 221005.](#)

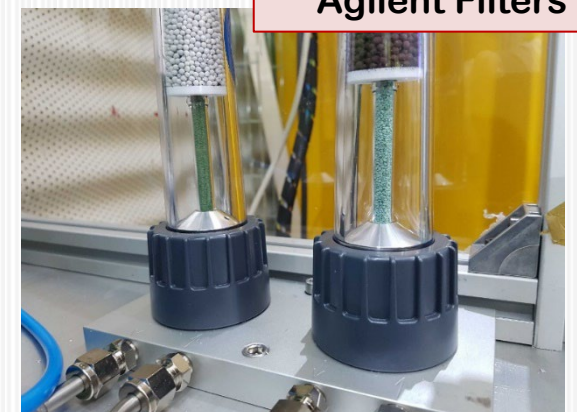
Purifiers: Rn in TREX-DM

- Recirculation:
gas through purifiers, which emanate Rn
- Several options used:
 - Agilent (for O₂ and **H₂O**)
 - SAES MicroTorr (O₂ and H₂O)
 - 5Å molecular sieve
 - Custom-made O₂+H₂O+Rn filter (developed by U. Birmingham)
 - Activated carbon filters
- For the moment, best results with open-loop low-flow
- On-going search for possible solutions:
 - Currently preparing tests with new material (silver-zeolites, Ag-ETS-10) to reproduce results in:

<https://doi.org/10.48550/arXiv.2505.07979>

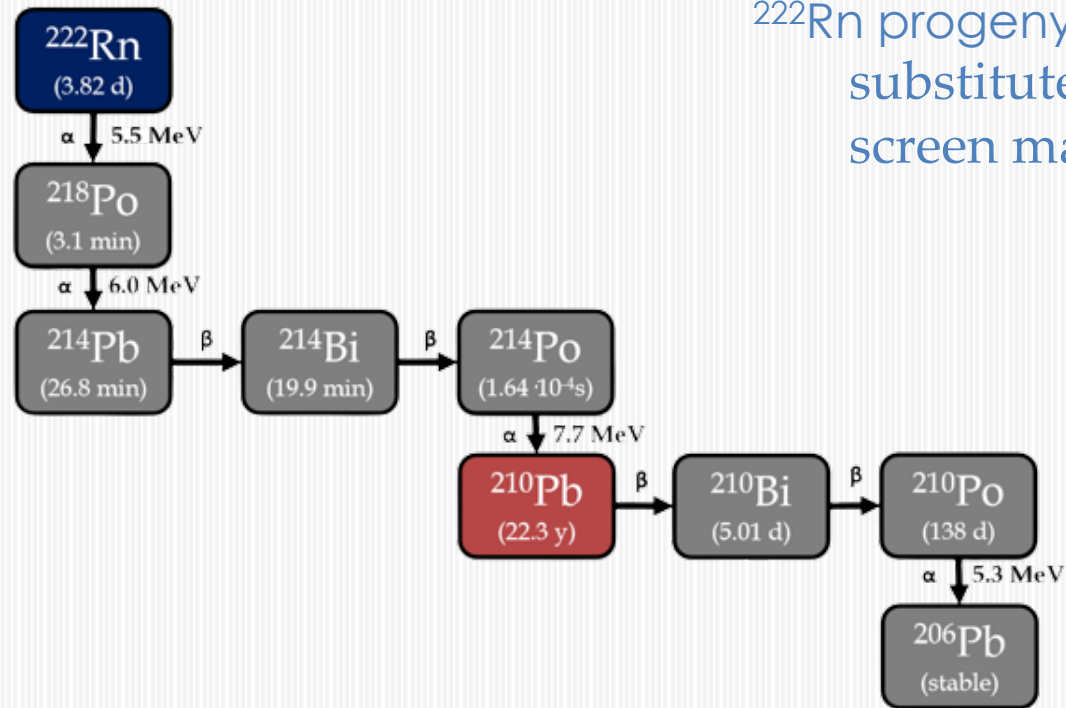


Ag-ETS-10

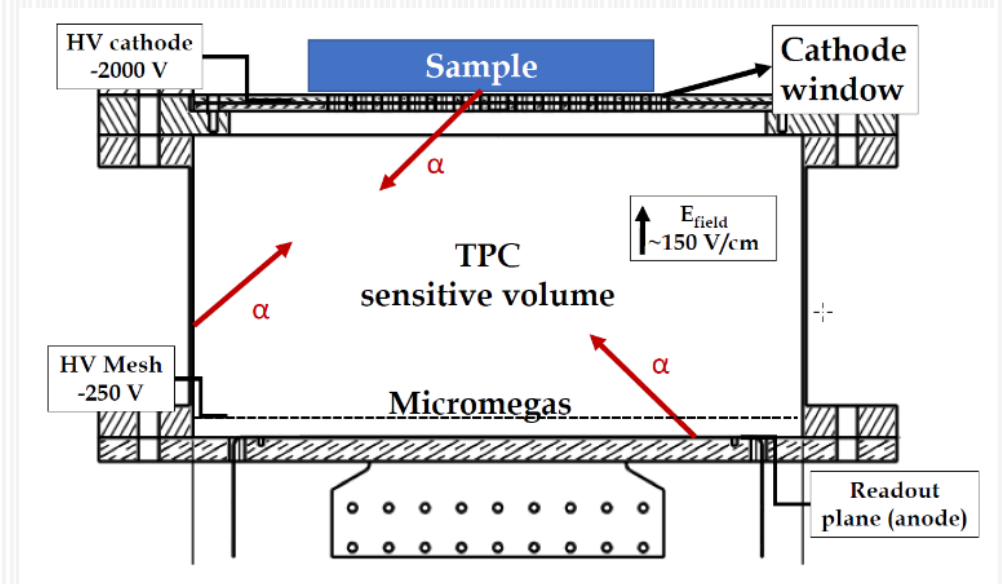


Background: Rn levels

- Background level
 - Switched from sealed mode to open loop: 600dru to 100dru (dru = $\text{keV}^{-1} \text{kg}^{-1} \text{day}^{-1}$)



^{222}Rn progeny contamination on mylar cathode surface
substitute cathode
screen material for alpha surface contamination

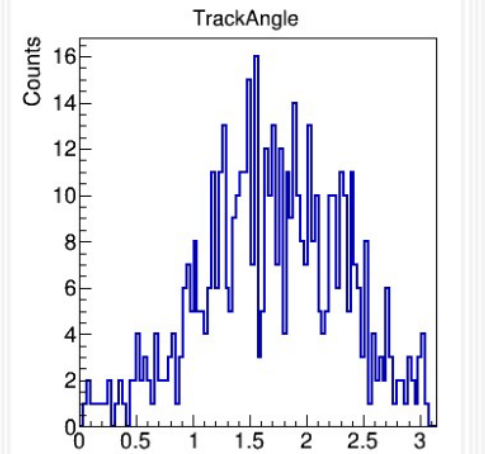
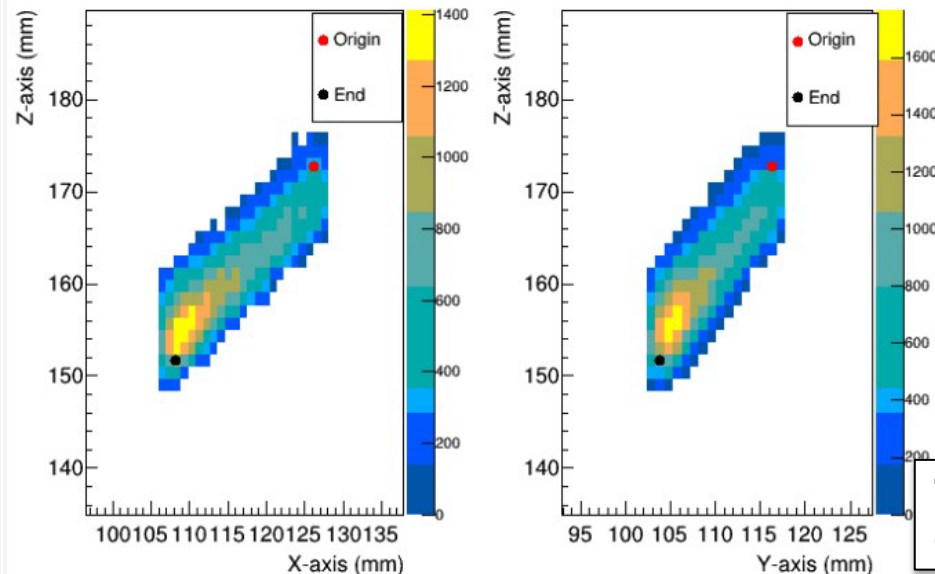
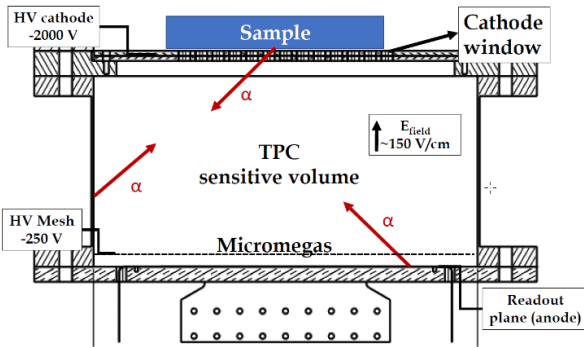


AlphaCAMM (Alpha CAMera Micromegas)

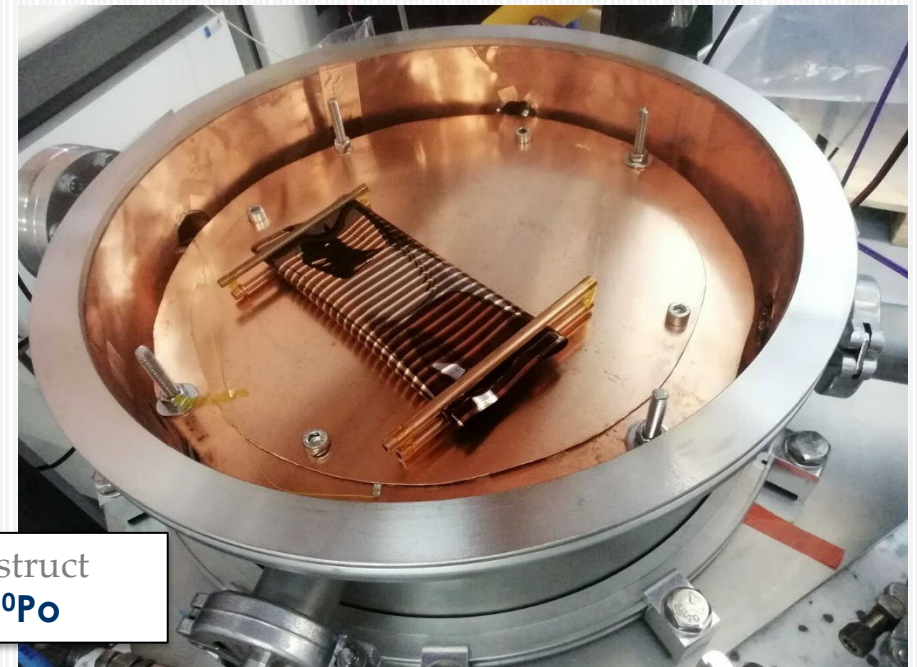
- ‘Spin-off’ of TRES-DM: gaseous chamber with a segmented mM (25cm x 25cm) to measure ^{210}Pb surface contamination of flat samples down to 100 nBq/cm^2

K. Altenmuller et al, 2022 JINST 17 P08035

Measurements of materials of interest already started:
aluminised mylar, ‘clean’ copper wires, copper-clad kapton, copper strips on kapton...

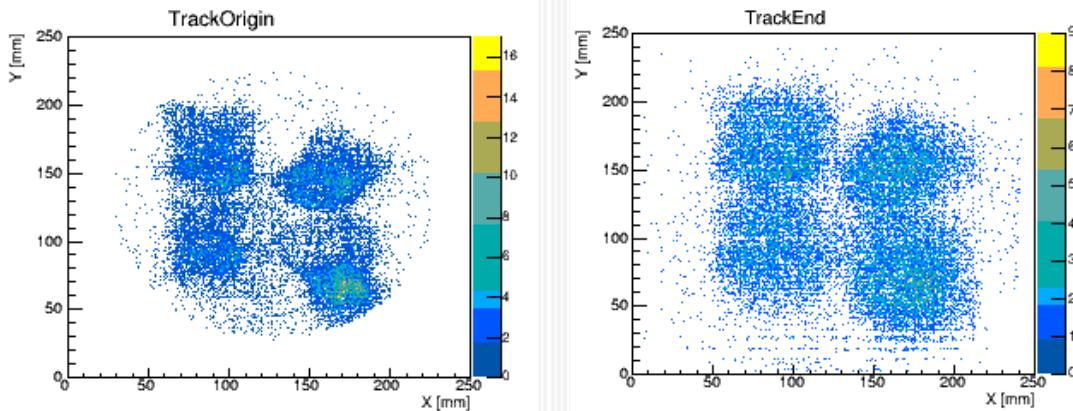
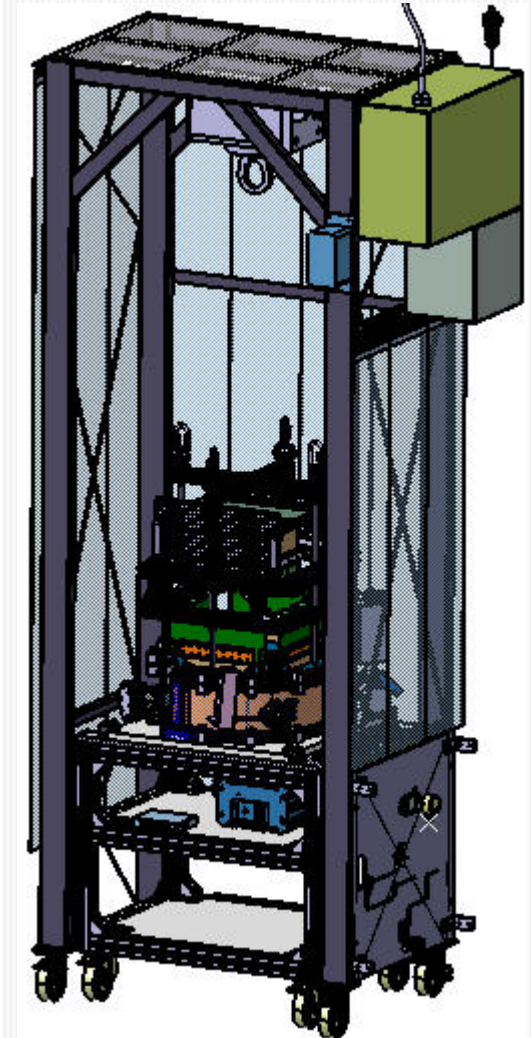
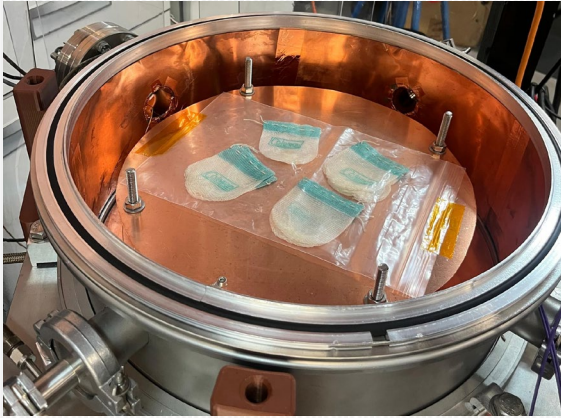


Topological information to reconstruct origin and end of a tracks from ^{210}Po



AlphaCAMM (Alpha CAMera Micromegas)

- Stand-alone facility for measurements
 - Open to anyone interested

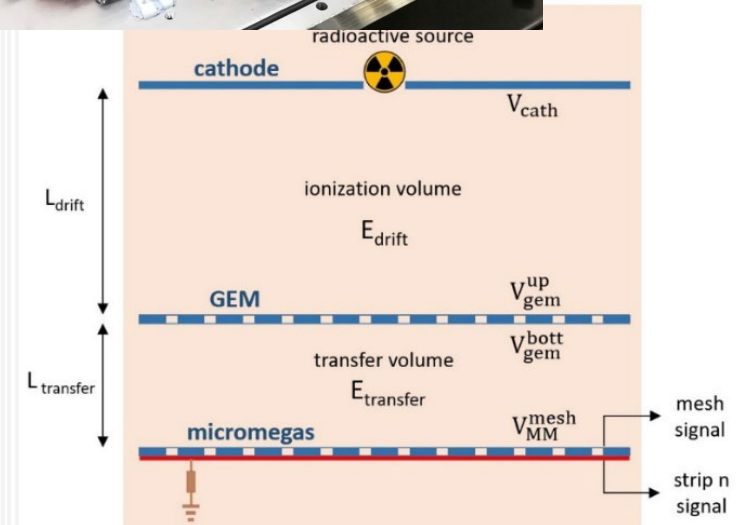
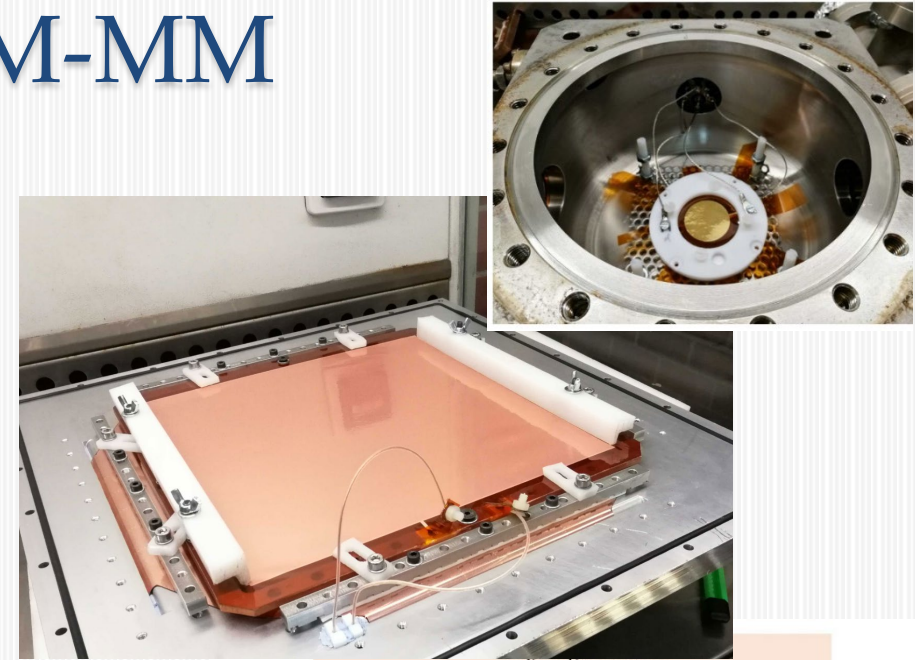
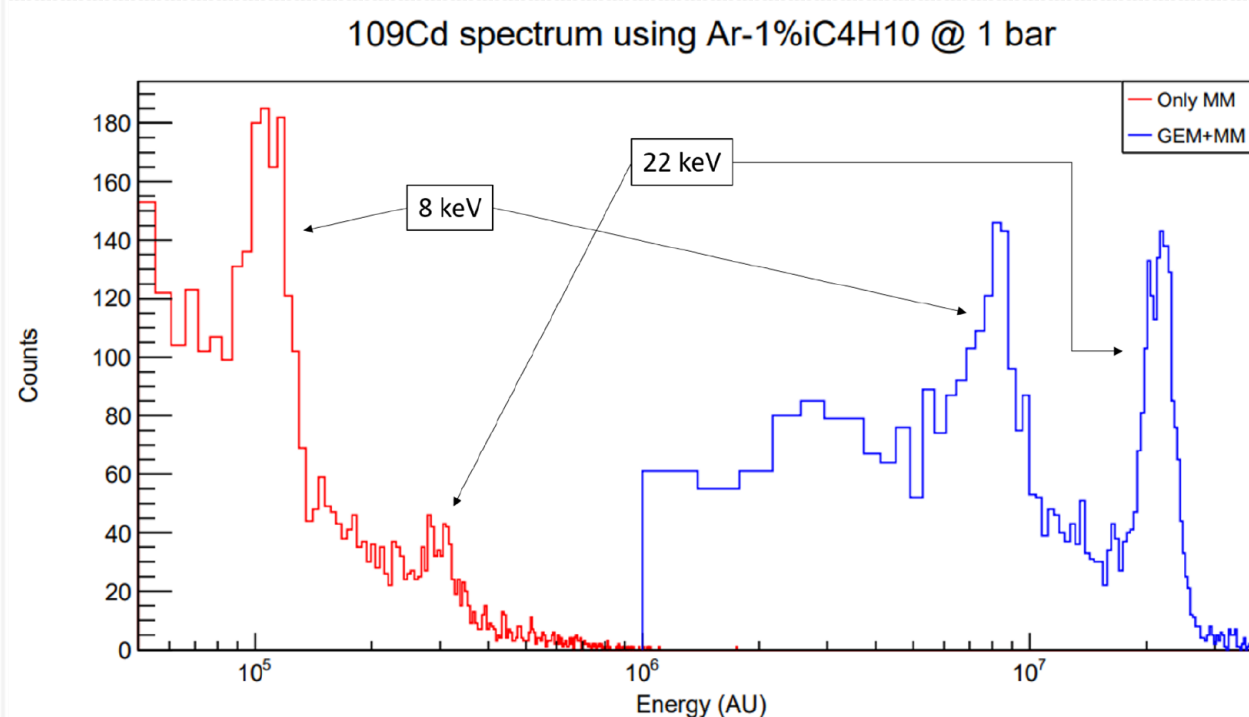


Low-energy threshold

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Lower threshold: GEM-MM

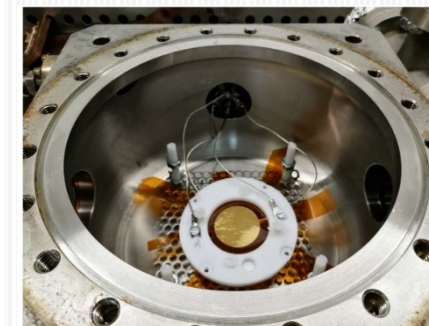
- Energy threshold
 - preamplification volume with a GEM on top of a mM factors would allow very low energy threshold (even single e-)
 - Big microbulk mM @1bar (x80)
 - Small microbulk mM @1-10bar (x90 to x20)



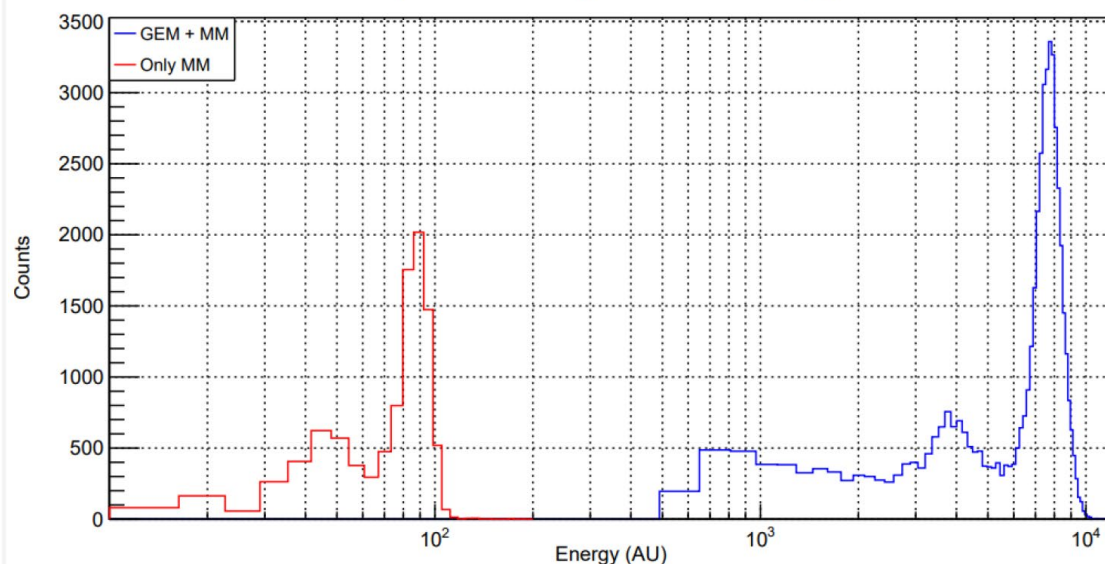
Open Res Europe 2025, 5:53
arXiv: 2412.19864 [physics.ins-det]

GEM-MM operation: different pressures

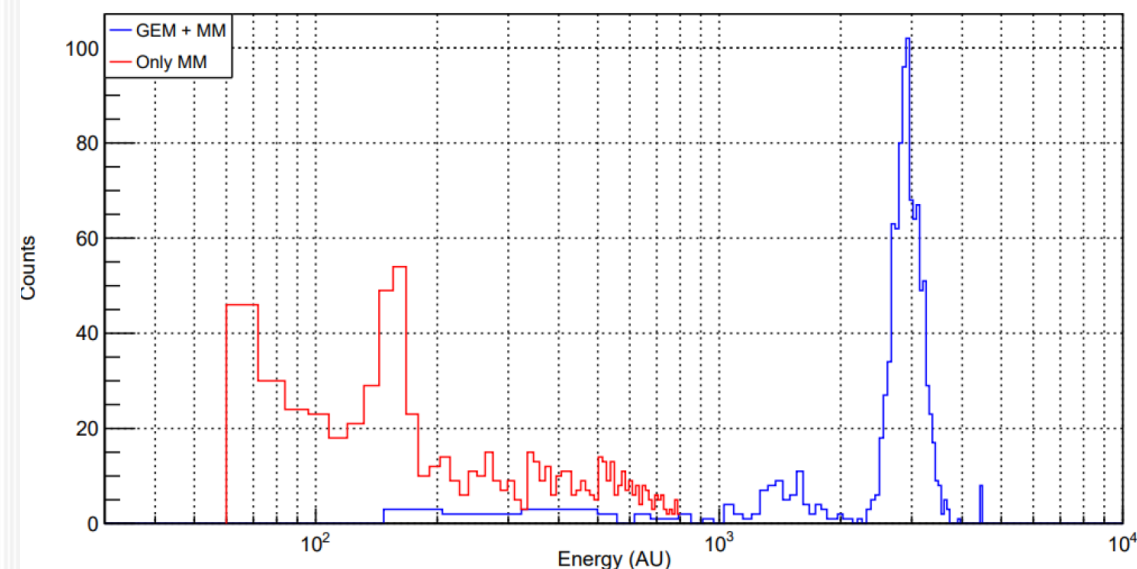
Pressure (bar)	V_{mesh} (V) (GEM + MM system)	V_{GEM} (V) (GEM + MM system)	Preamp. factor	V_{mesh} (V) (only-MM system)	GEM effective extra gain factor
1	305	310	-	315	90
4	390	410	70	400	50
10	535	550	21	540	19
1	290	285	85	293	80



55Fe spectrum using Ar-1%iC4H10 @ 1 bar



55Fe spectrum using Ar-1%iC4H10 @ 10 bar



Open Res Europe 2025, 5:53
arXiv: 2412.19864 [physics.ins-det]

Low energy, volume calibrations

Gas source: distribution in all volumen, homogeneous calibration of the readout plane.

^{37}Ar : 2.8 keV (90%), 0.27keV (9%)

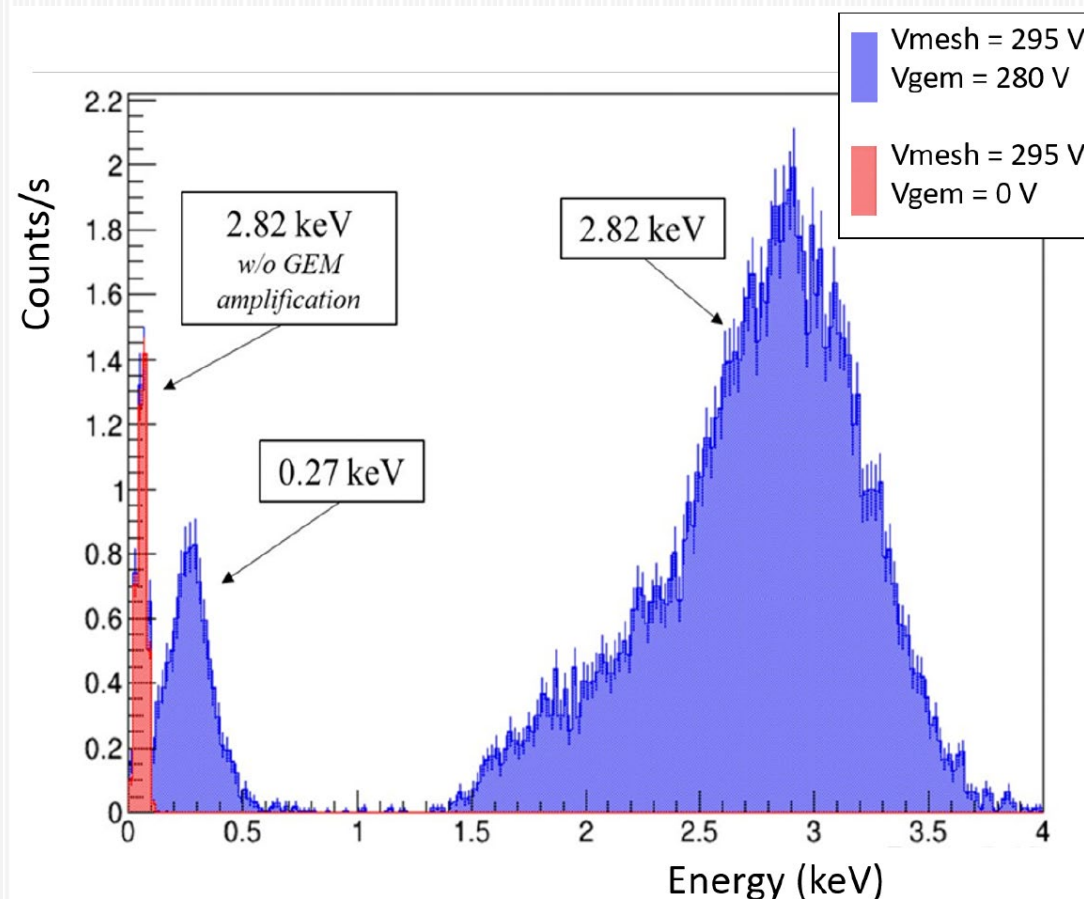
- Irradiated CaO powder at HiSPANoS

^{40}CaO



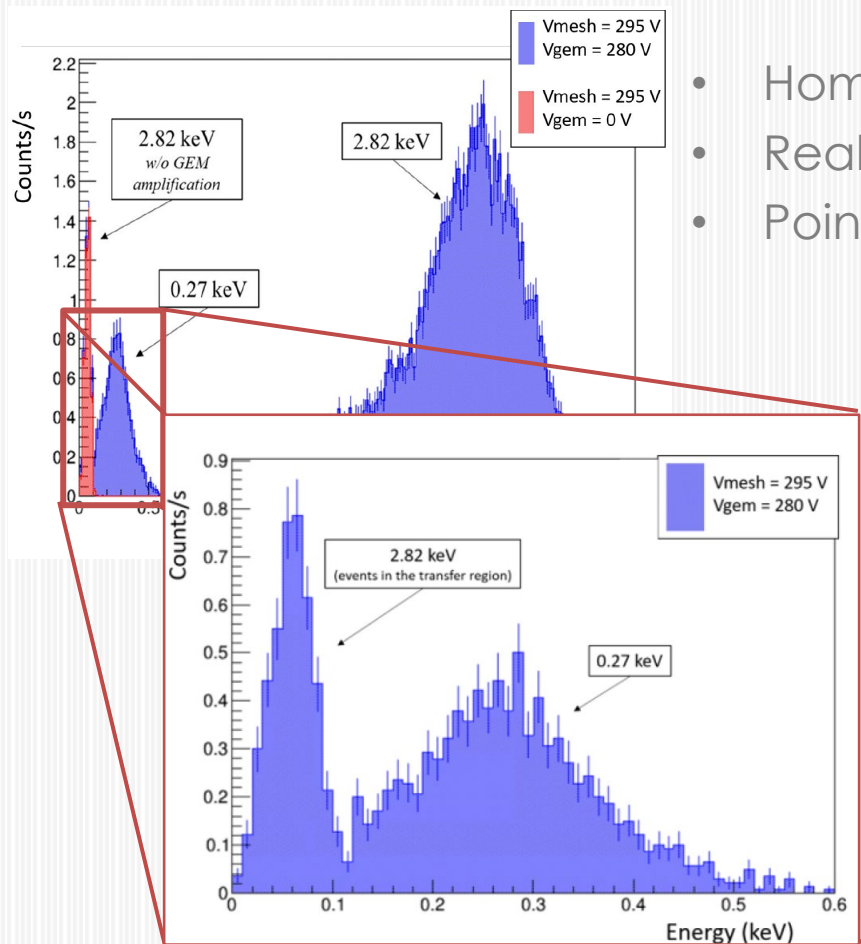
publication in preparation

PhD thesis of Ó. Pérez Lázaro, [arXiv:2507.02172](https://arxiv.org/abs/2507.02172) [physics.ins-det]

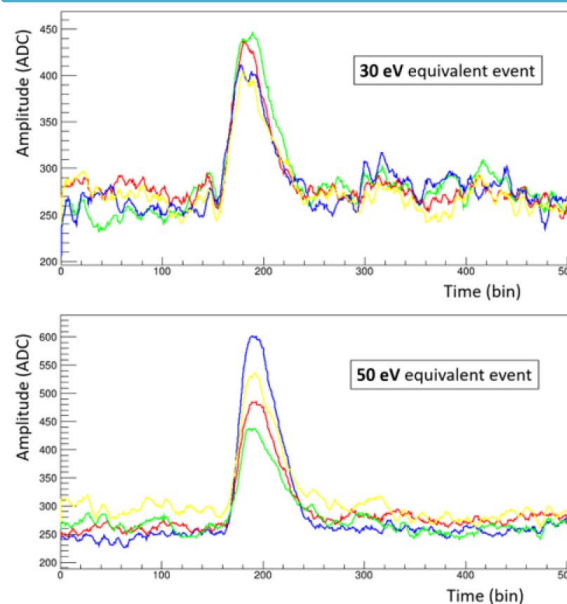


37Ar calibrations

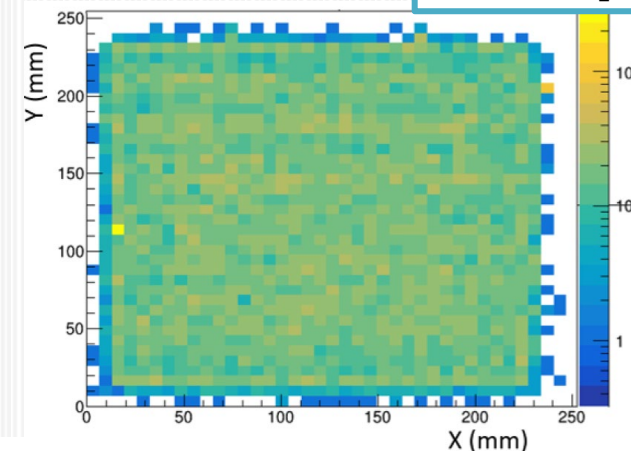
- Homogeneous illumination and response
- Real signals of sub-100eV events
- Point to $O(10)$ eV threshold



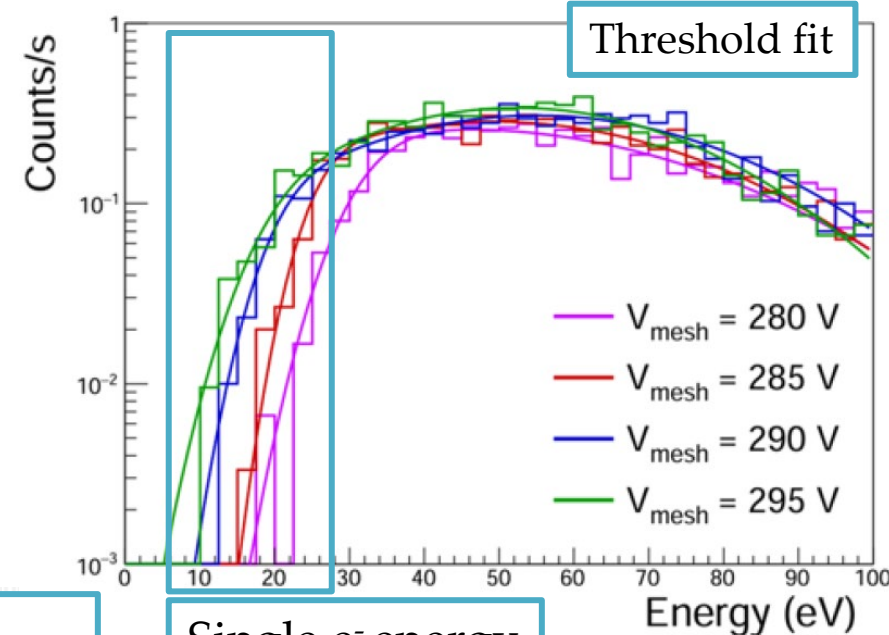
Examples of low-energy events



2D hitmap



Threshold fit



Single e^- energy

a window to uncharted territory for this kind of detectors

^{37}Ar calibration, nominal gain

Threshold still clearly $\sim 200\text{eV}$

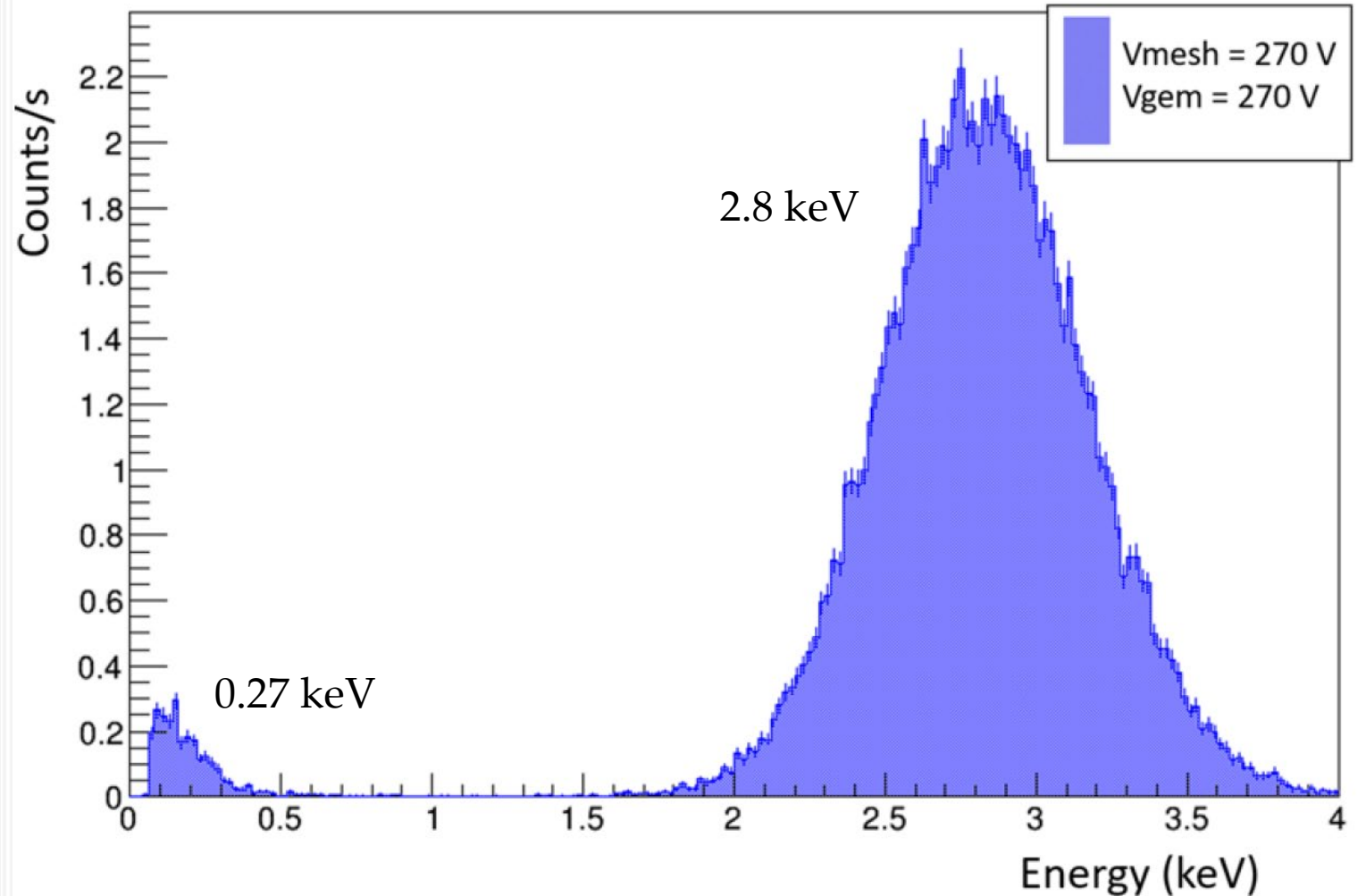
Some loss of linearity...?

Need for multiple low/energy calibration points

Also of interest:

^{83}mKr :

30keV , 7.6keV , 2.1keV , 1.8keV



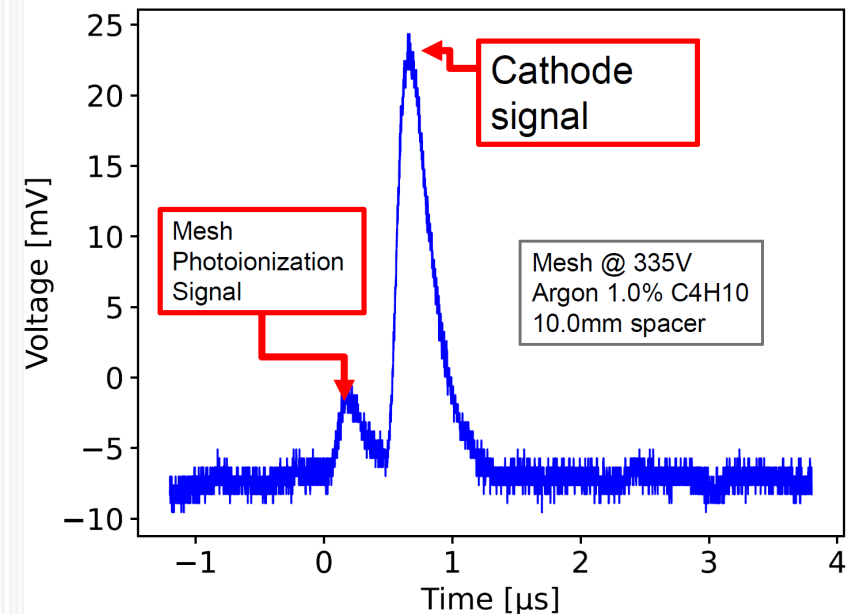
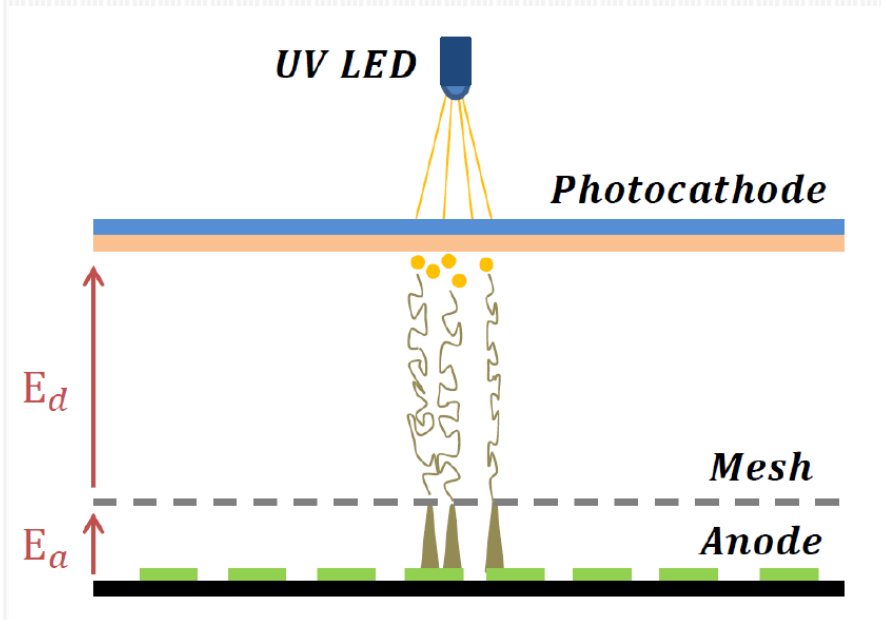
General R&D

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General R&D

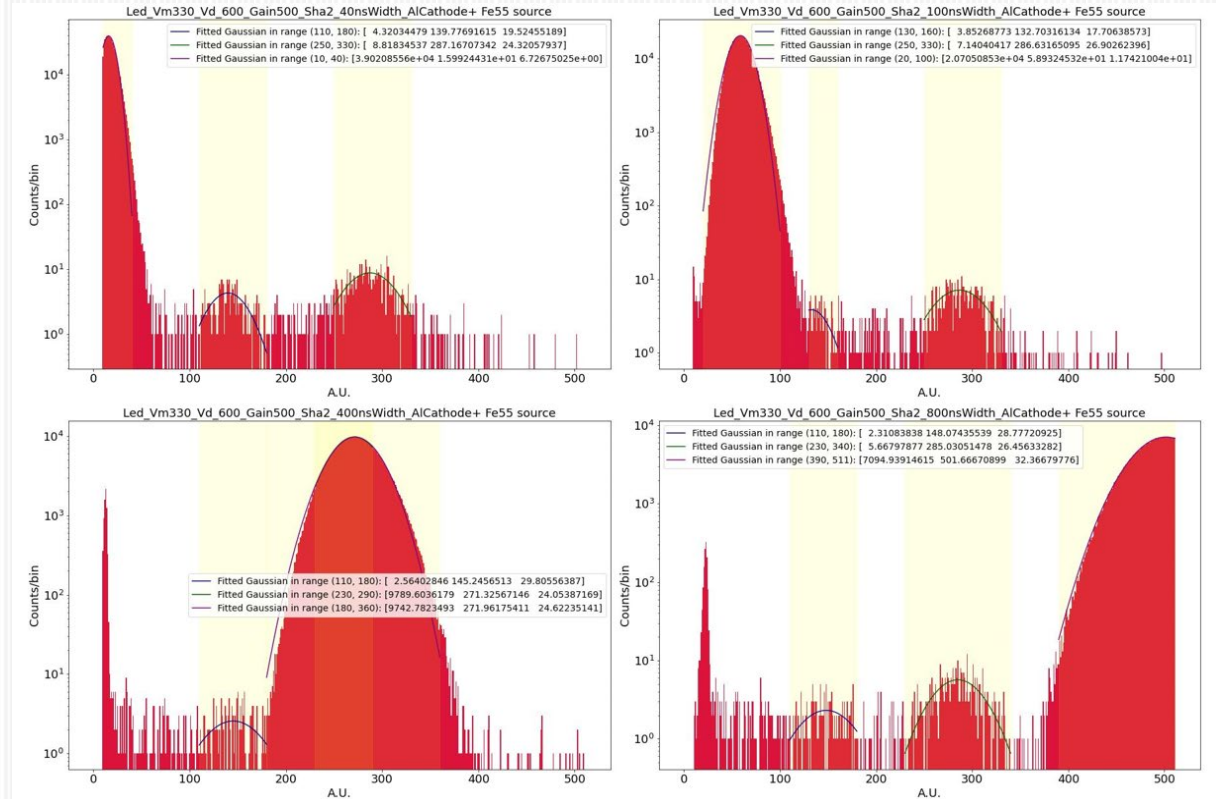
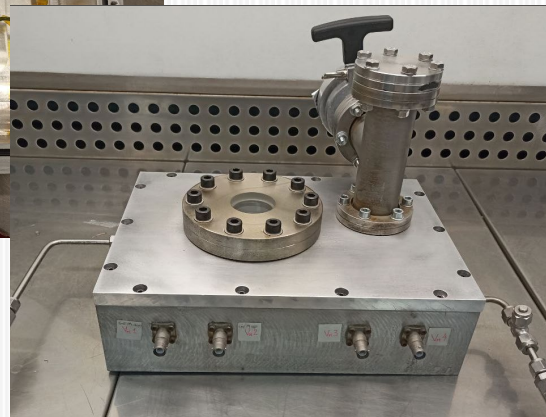
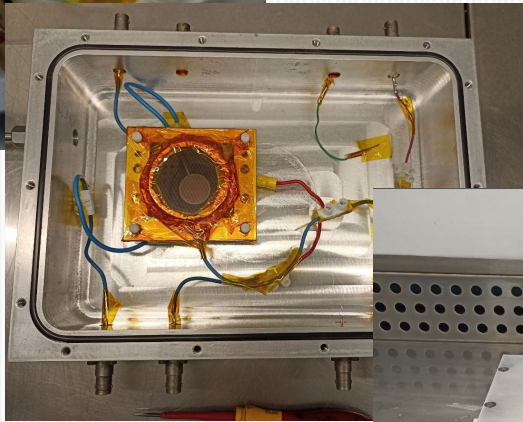
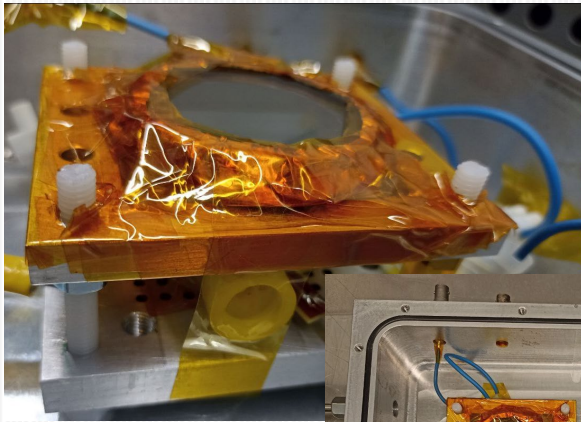
- A facility for gas drift velocity characterisation
 - T0 from the source
 - ToA of photocathode electrons

At CEA/Saclay



General R&D

- Variable energy calibration method: study the energy threshold limits
 - Down to single electron calibration

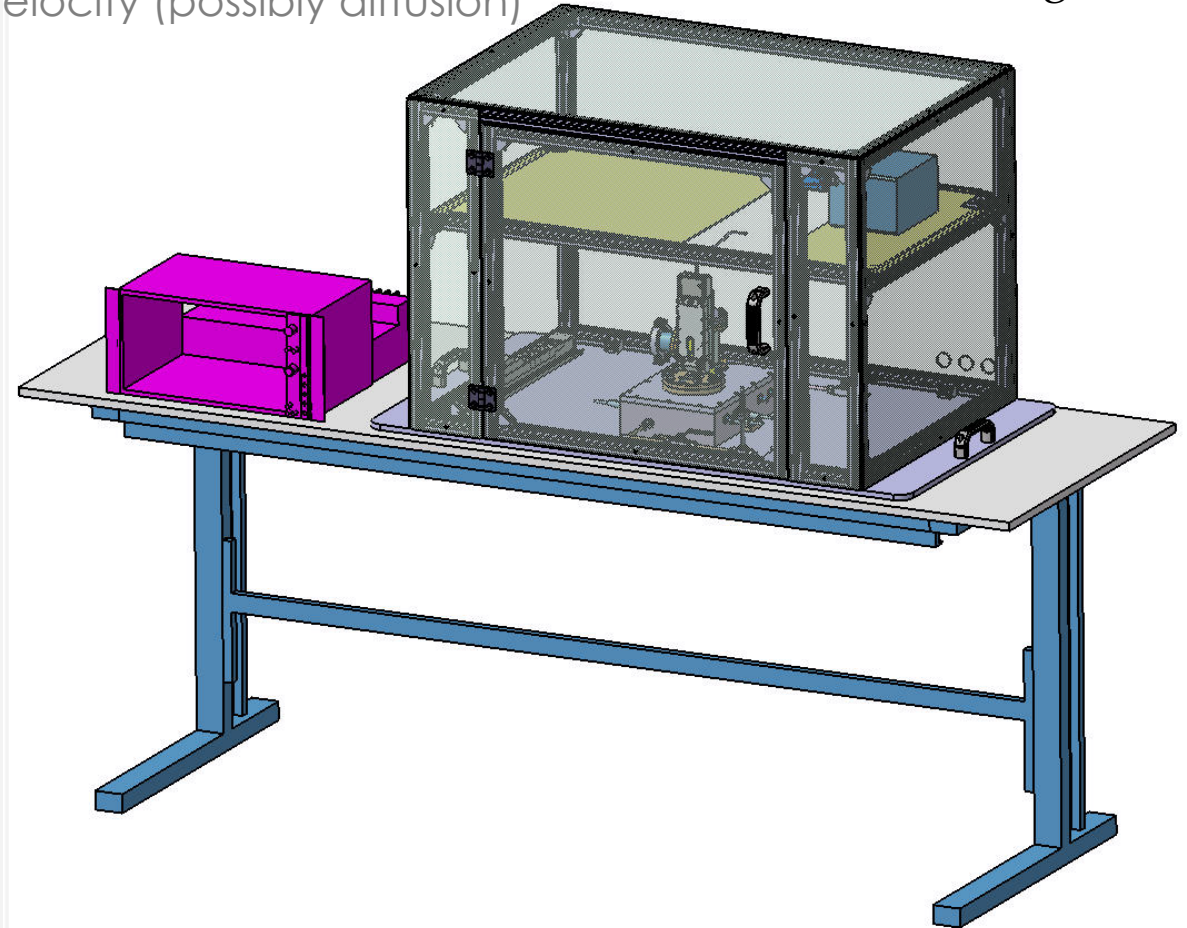
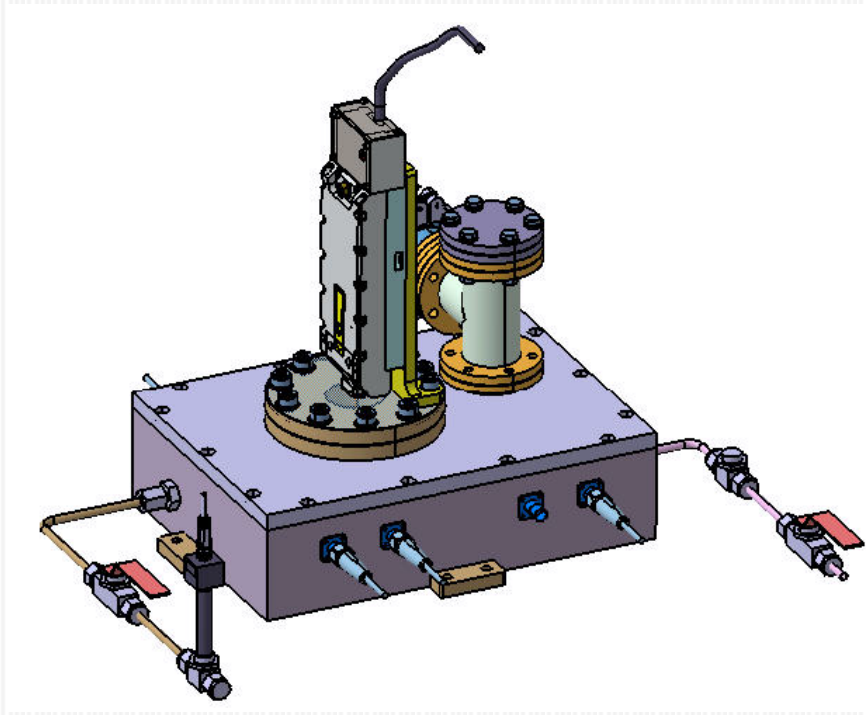


PhD thesis of D. Díez Ibáñez

General R&D

- A facility for gas studies
for energy calibration, spatial resolution and gas drift velocity (possibly diffusion)
 - Laser purchased
 - Cabin material on the way

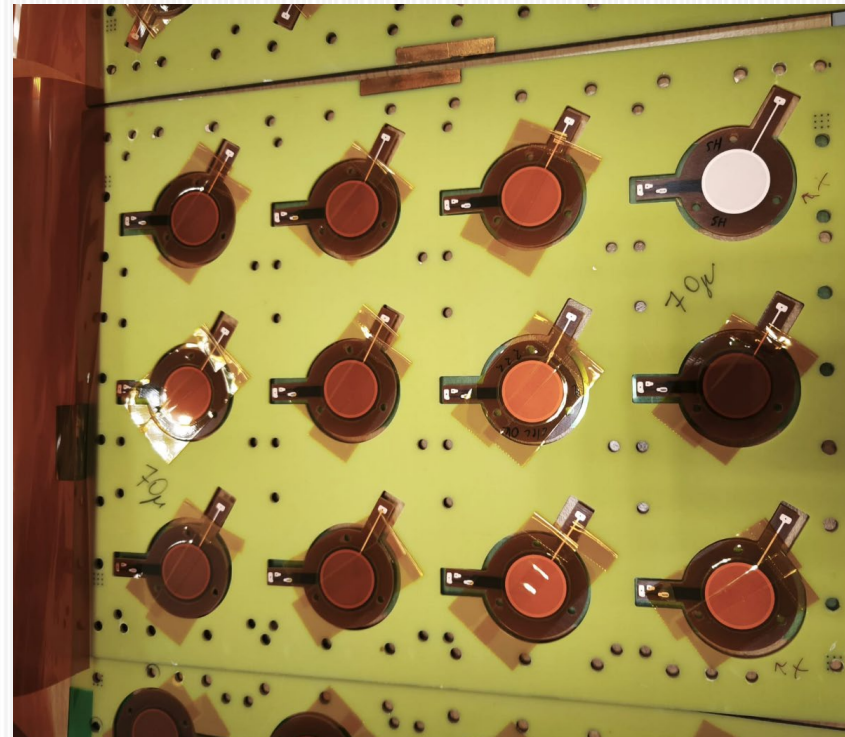
Offered at the discussions for
the nextAIDA through WG3



General R&D

- Resistive microbulk Micromegas
 - Should allow safer operation at higher gains
 - For the moment, radioactivity measurements are not Good
 - Active R&D with MPT@CERN
 - Batch of small detectors to characterize should arrive in the next days

At MPT/CERN
tests at 500V ok
checks at 700V pending



Comments

Although driven by our primary interests in baby(IAXO) and TREX-DM

- Baby(IAXO) optimizes all parameters for the most sensitive Helioscope
- TREX-DM offers a technology that can be very sensitive to low-mass WIMPs

Open detector challenges being addressed: Mostly encompassed in DRD1

- Improve the radiopurity of detectors of this type in order to reduce their intrinsic background
- Improve the energy threshold of detection (by improving operational gain and/or EM noise)
- Improve general robustness of detector, technical design choices, novel gas mixtures and operation points, etc. to facilitate future implementations: scaling-up to larger detector sizes, handling large readout granularity, etc
- **Busy past has shown the way to... a busy future ahead**
- **Continuous R&D gives birth to 'spin-off' projects of great interest**
 - AlphaCAMM for Alpha Surface contamination
 - Laser setup for gas characterisation even at HP

Dziękuję
...