

3D geometry and magnetic connections of erupting black hole jet

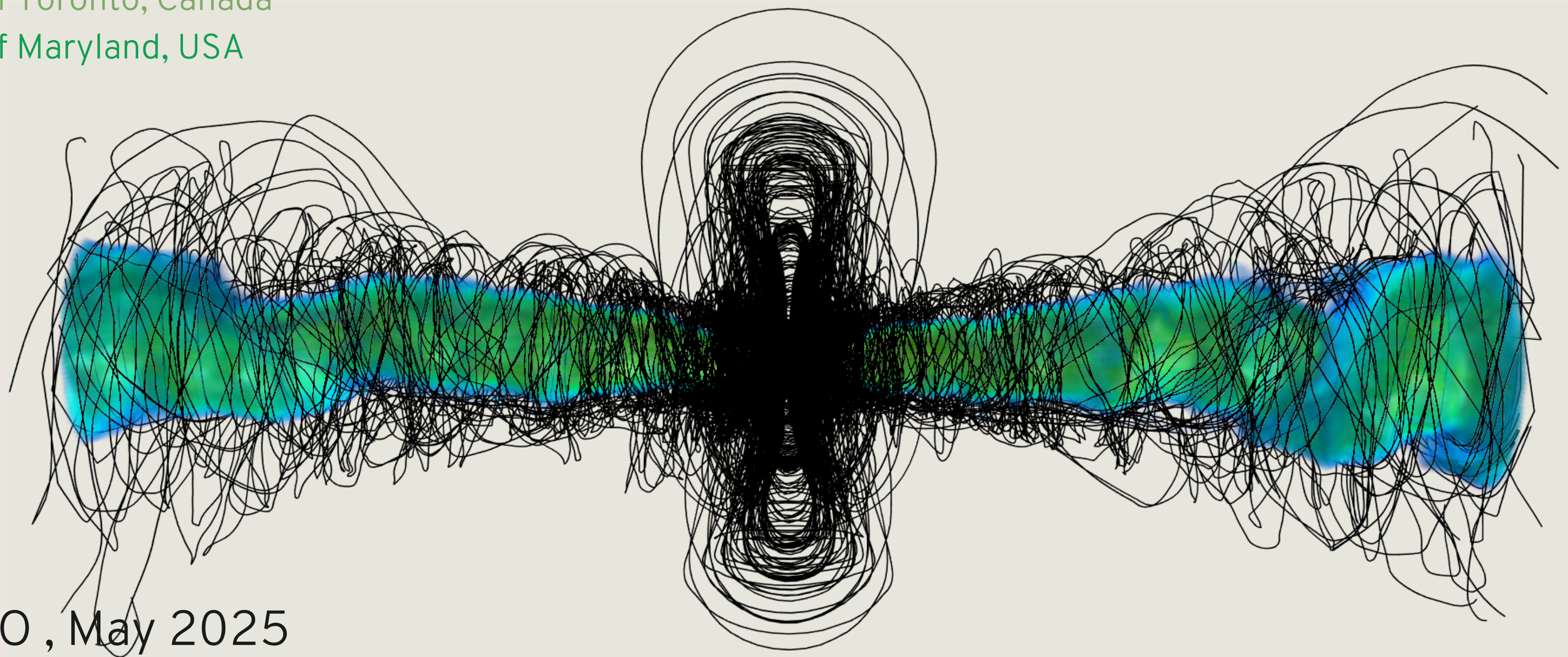
M. Kapusta, K. Nalewajko, B. Ripperda, S. Philippov

University of Warsaw, Poland

Nicolaus Copernicus Astronomical Center, Poland

University of Toronto, Canada

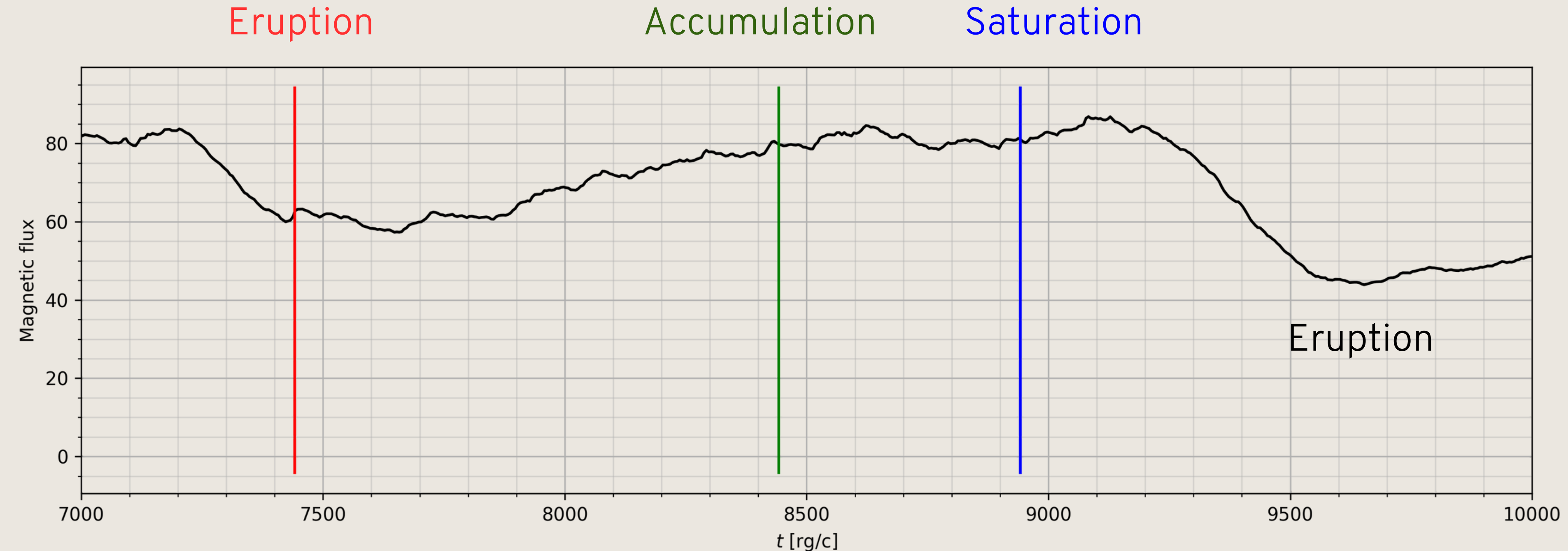
University of Maryland, USA



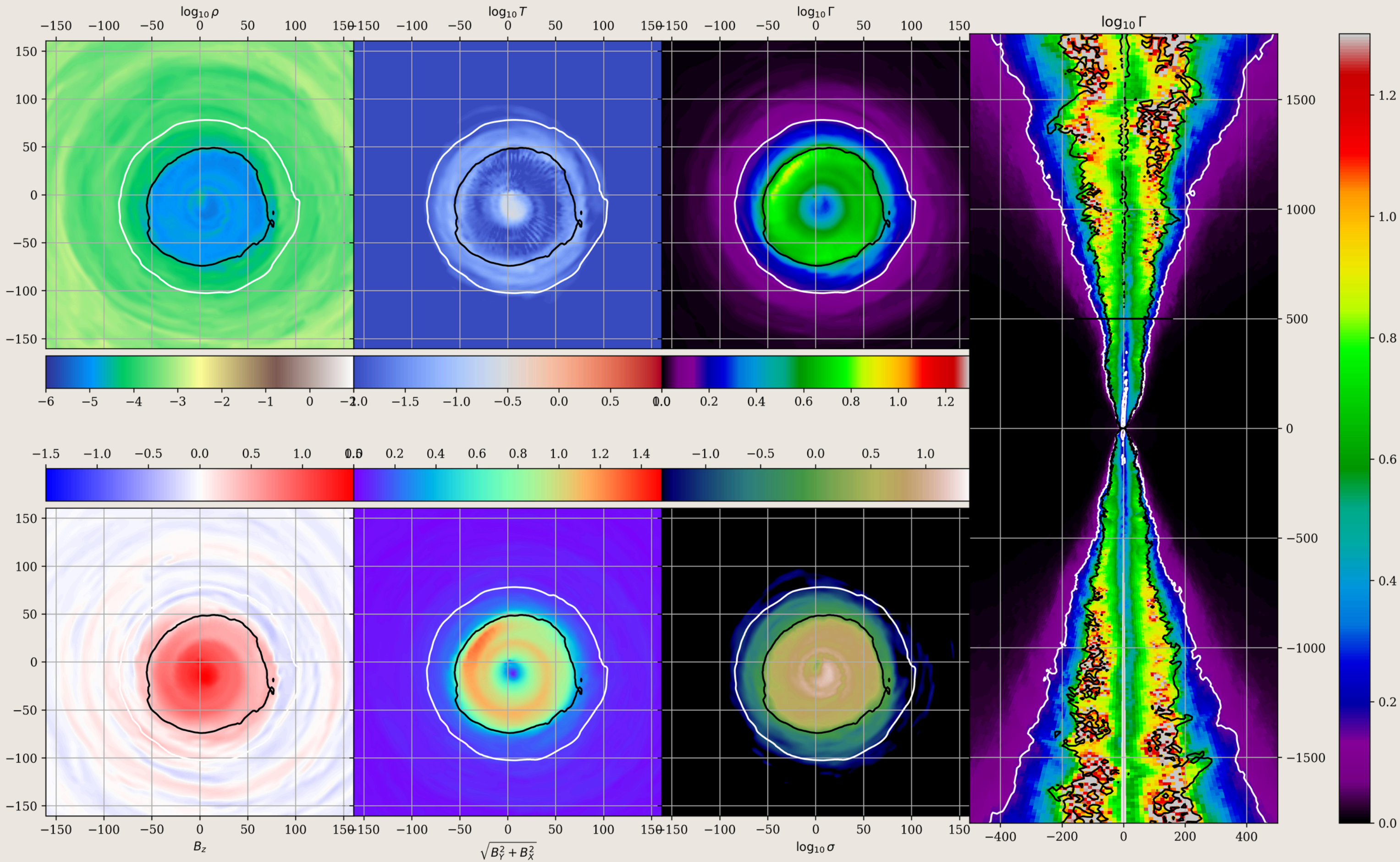
ReFCO , May 2025

Analysis of extreme-resolution 3D GRMHD simulation

- Data from the H-AMR code (Ripperda et al. 2022), resolution (5376 x 2304 x 2304)
- Cycles of magnetic field eruption: just after **eruption E**, **accumulation A**, **saturation S**.

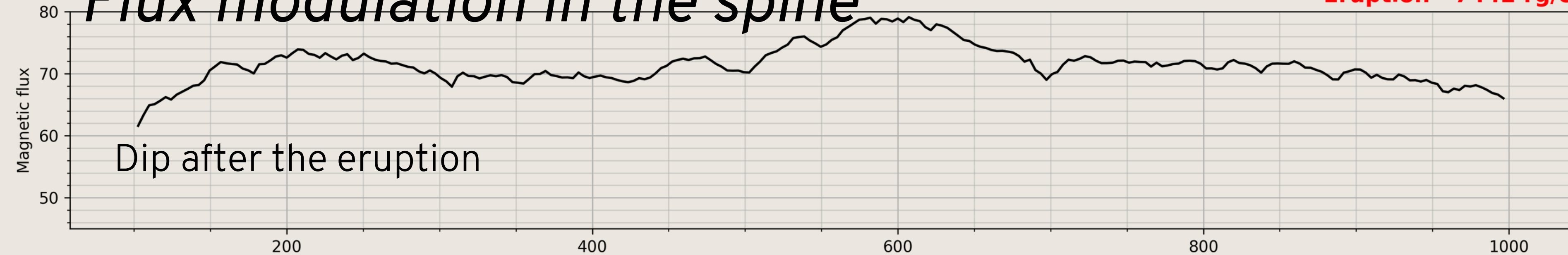


Spine ($\sigma > 1$), Sheath ($\Gamma > \text{sqrt}(2)$)



Flux modulation in the spine

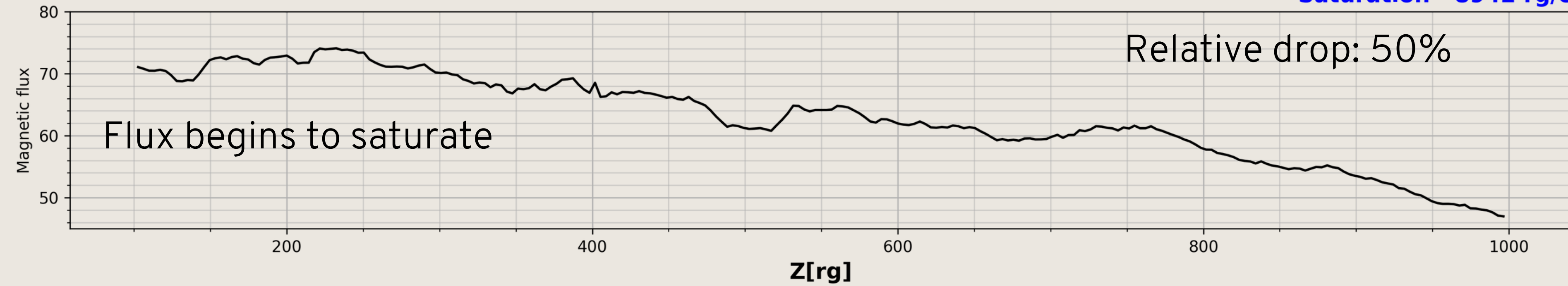
Eruption - 7442 rg/c



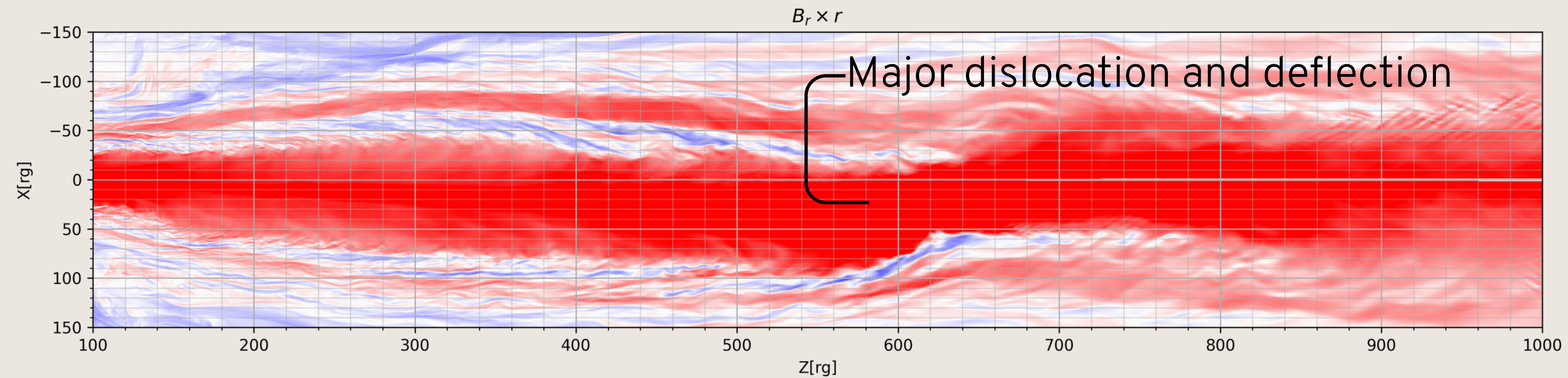
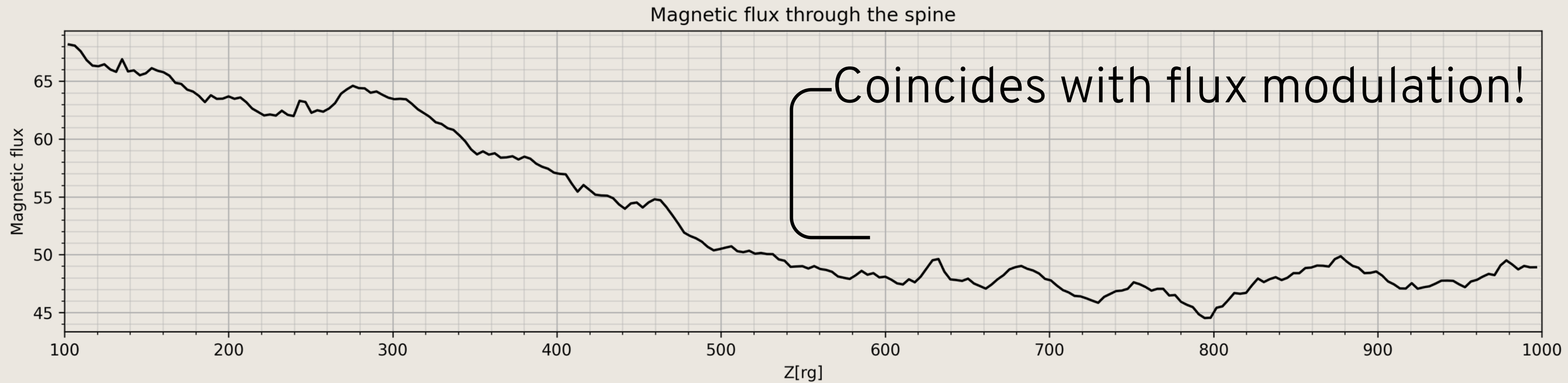
Accumulation - 8442 rg/c



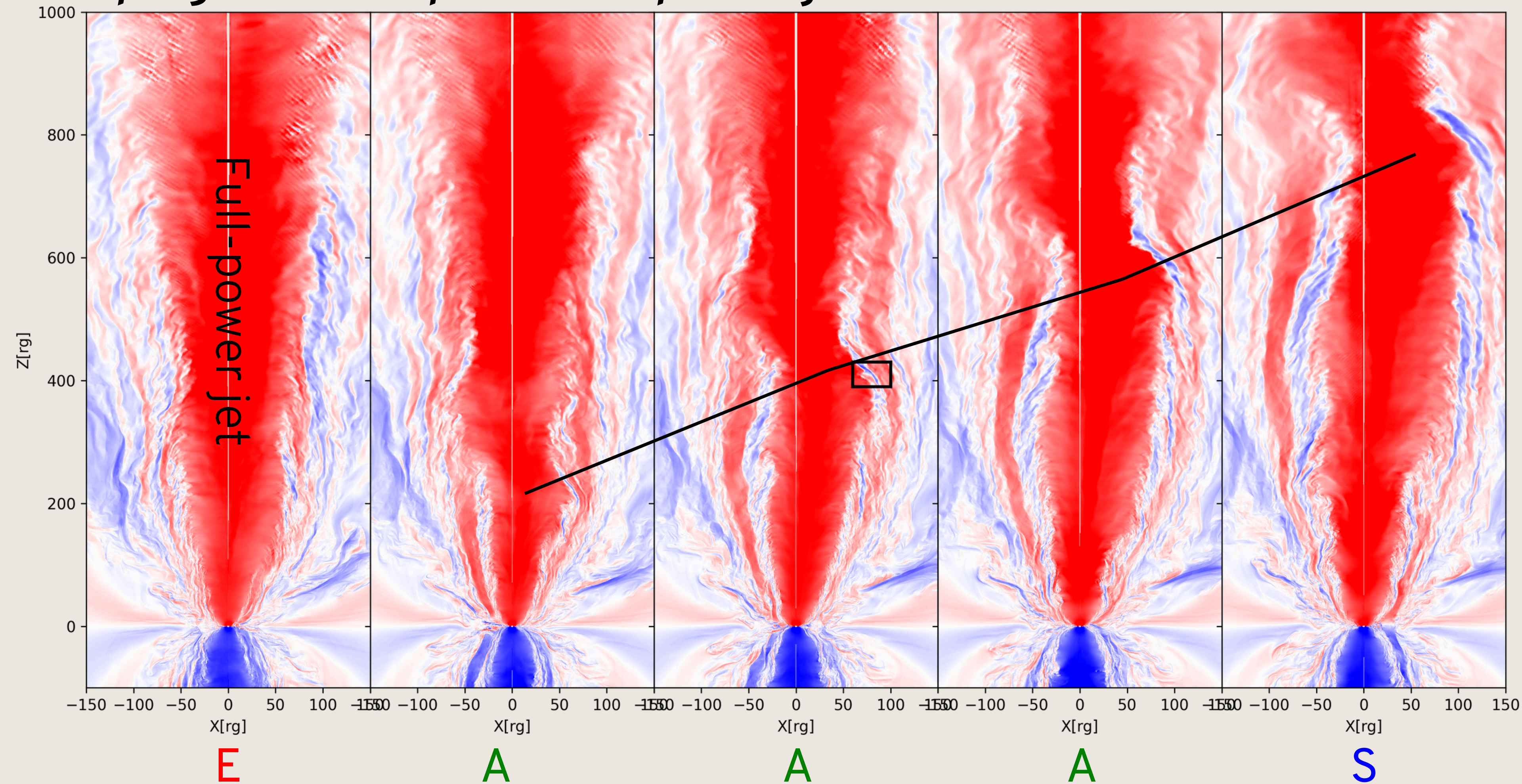
Saturation - 8942 rg/c



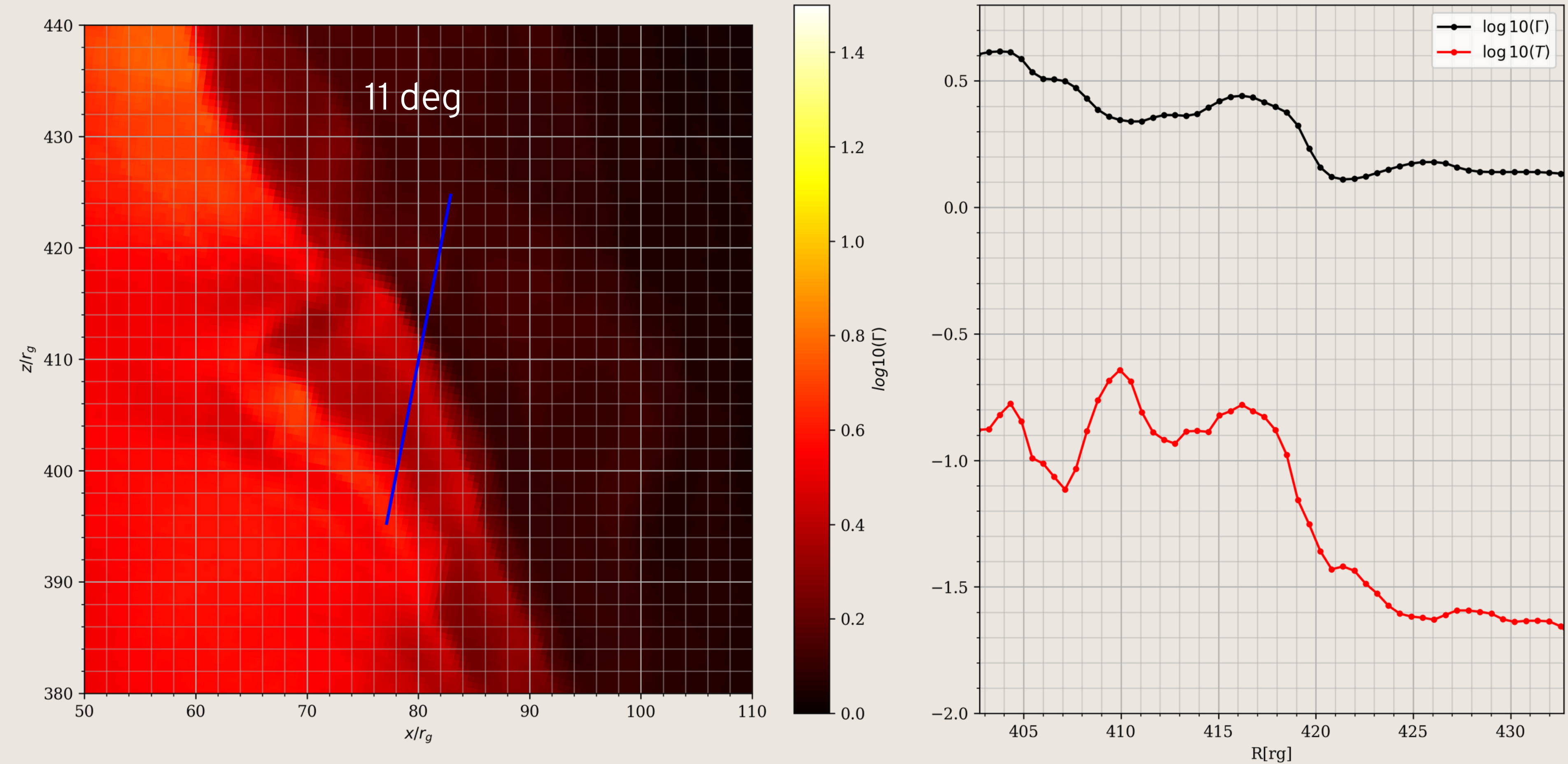
Disruption- accumulation state - 8442 rg/c



Propagation of post-eruption jet distortion



Qblique forward shock - accumulation state 8442 rg/c



Conclusions

- Magnetic field eruptions in MAD systems induce a temporary reduction of magnetic flux through the jet.
- A distortion propagates along the jet spine, which can be associated with the recovery of magnetic and Poynting fluxes.

Thank you for your attention!