

# Radiative model of magnetically-arrested flows and its application to M87\*

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# MHD solutions

3D GRMHD, single-fluid simulations performed with  
HARM\_COOL (Sapountzis & Janiuk 2019, Janiuk & James 2022)

resolution:  $288 \times 256 \times 256$  (R- $\theta$ - $\phi$ )

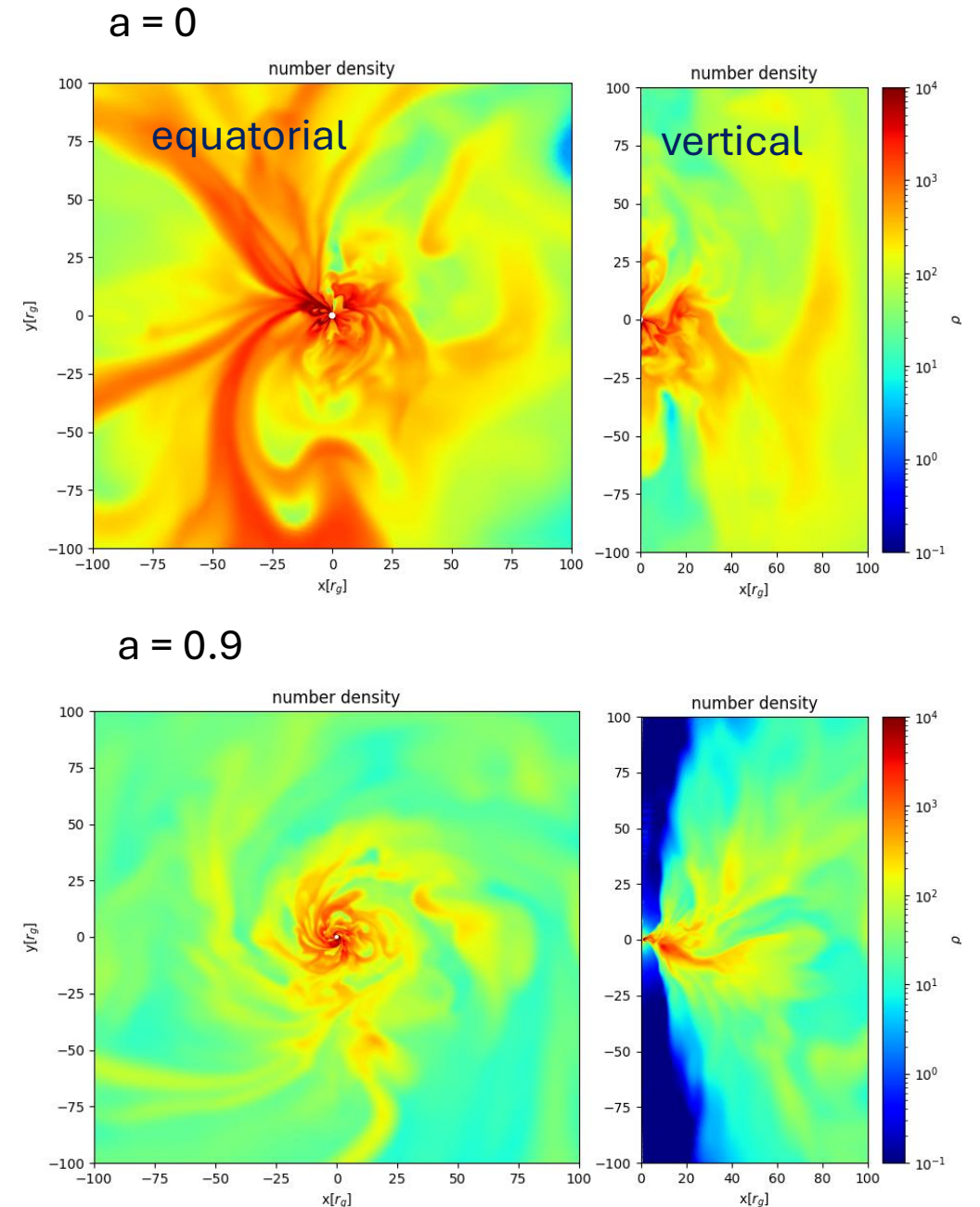
time: 50000 GM/c<sup>3</sup>

all solutions in the MAD state,  $\phi_{\text{BH}} \sim 20$  (MAD-ness parameter)

scaled to  $M = 6.5 \times 10^9 \text{ Msun}$

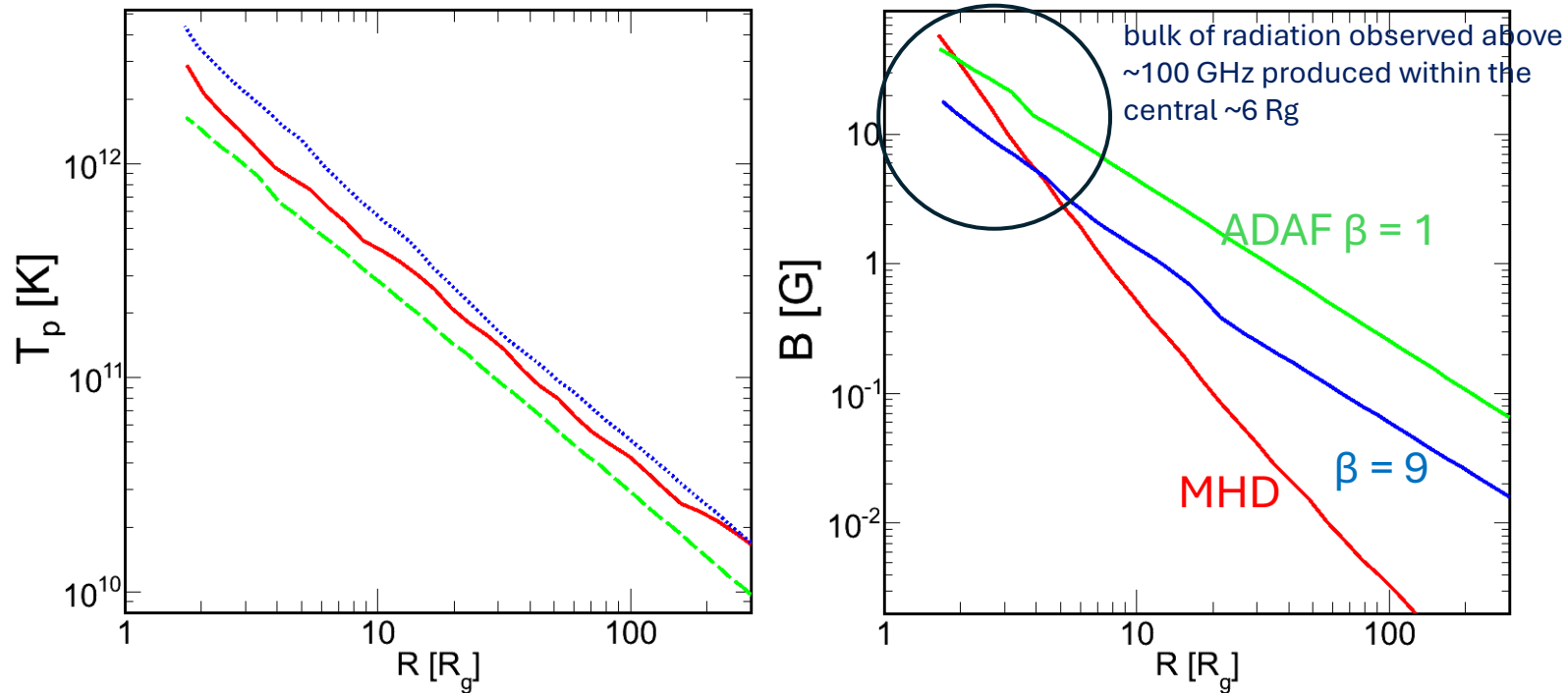
for  $a=0.9$  strong jets (with a power  $\sim 3 \times 10^{43} \text{ erg s}^{-1}$  for the fitted  
accre. rate  $\sim 0.01 \text{ Msun yr}^{-1}$ )

for  $a=0$  no jets, also quasi-spherical structure in the central part



# MHD solutions vs ADAF theory

Radial profiles – very similar in both models for all parameters except for  $B$ , much flatter in analytic models – due to assumption of constant  $\beta = P_{\text{gas}}/P_{\text{mag}}$



Red: angle-averaged radial profiles in the GRMHD solutions

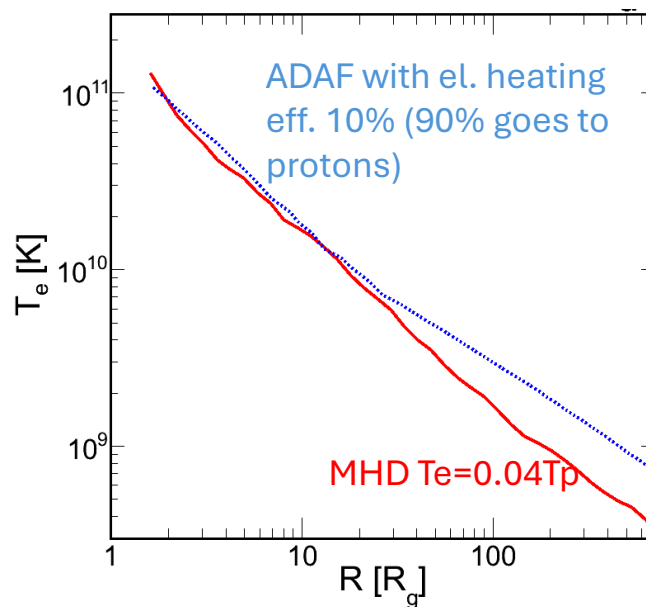
Blue and green: ADAF (two-temperature flow) model of Shapiro et al. (1976), in the GR form Abramowicz et al. (1996) Gammie & Popham (1998), Manmoto (2000)

Both MHD and ADAF models for the same  $a=0.9$ , accr. rate =  $0.1 \text{ Msun/year}$ ,  $M = 6.5 \times 10^9 \text{ Msun}$

# Thermal synchrotron spectra

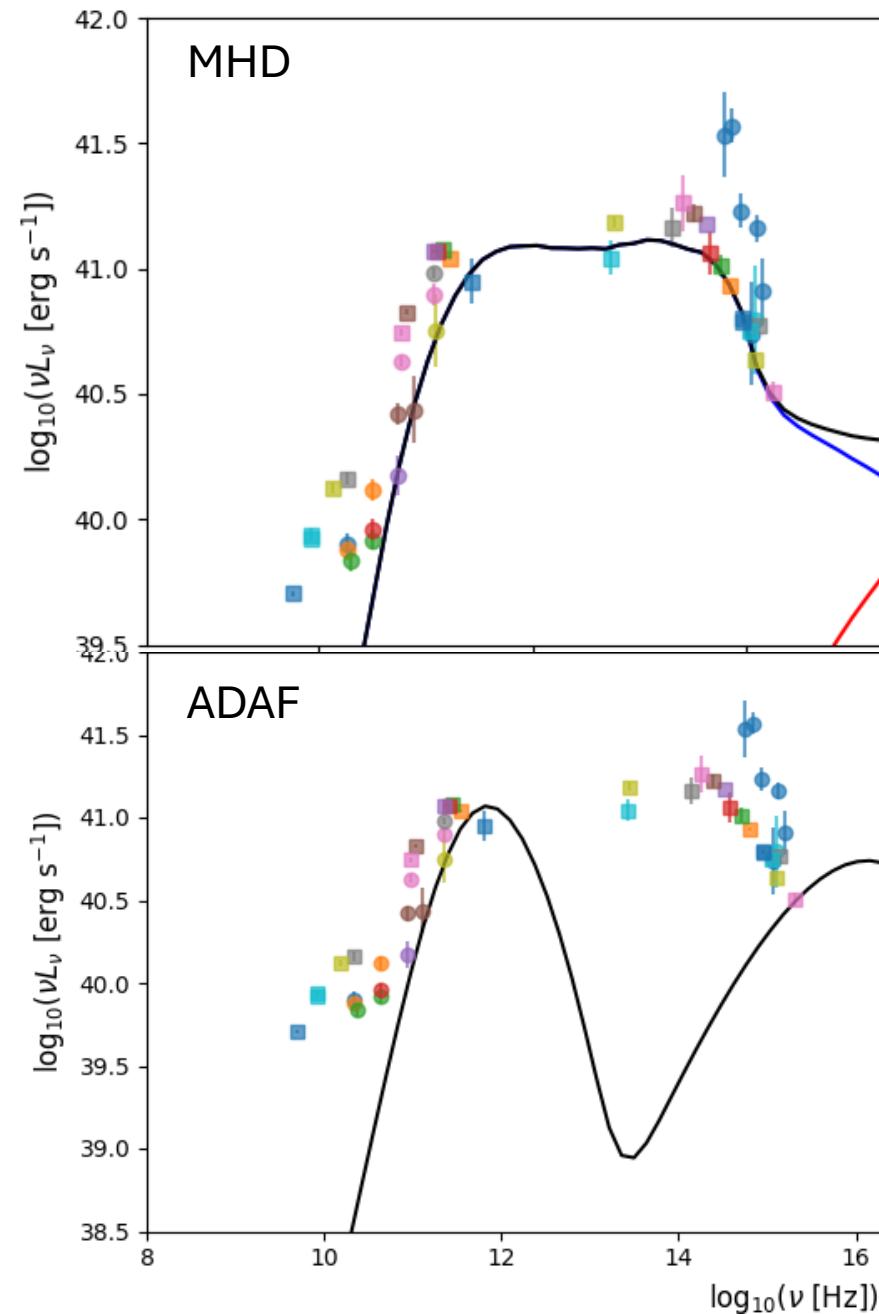
For MHD and ADAF spectra computed with the same GR Monte Carlo code

For the MHD model we assumed  $T_e = 0.04 T_p$ , for which the average  $T_e$  is very similar to ADAF at this accr. rate (also very similar to radiative GRMHD simulation of Chael et al. 2019)



data for 2018 multiwavelength campaign of M87

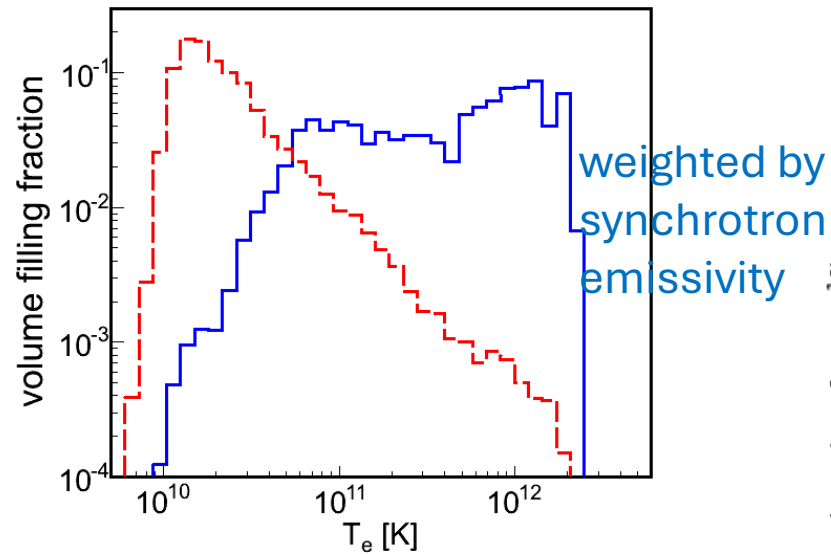
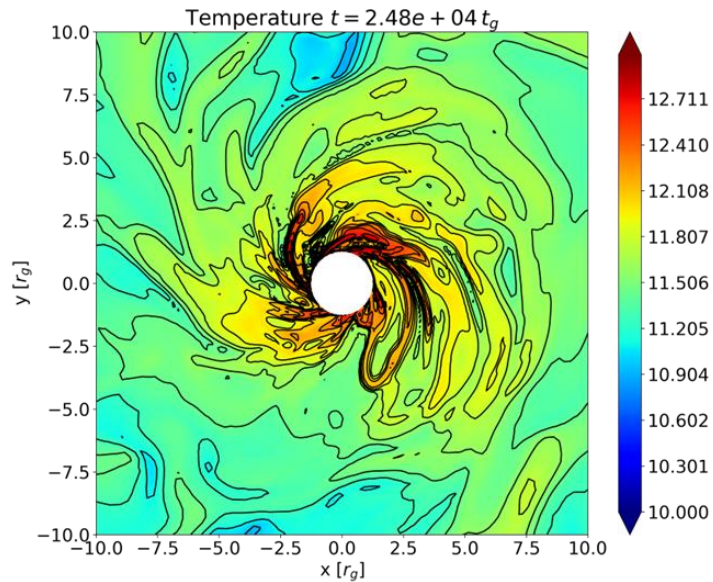
VERA	HST	HST/ACS-HRC F220W	HST/NIC2 F110W	ALMA 252 GHz
EAVN/KaVA	Swift-UVOT	HST/STIS 2360 AA	HST/NIC3 F166N	ALMA 221 GHz
VLBA	Chandra + NuSTAR	HST/ACS-HRC F250W	HST/NIC3 F222M	ALMA 108 GHz
Global VLBI	Fermi-LAT	HST/ACS-HRC F330W	Gemini 10.8 micron	ALMA 94 GHz
GMVA+ALMA	H.E.S.S.	HST/ACS-HRC F475W	Keck 11.7 micron	VLA A-conf. 22 GHz
KVN	VERITAS	HST/ACS-HRC F606W	ALMA 635 GHz	VLA A-conf. 15 GHz
ALMA	MAGIC	HST/ACS-HRC F814W	ALMA 350 GHz	VLA A-conf. 8.4 GHz
SMA	HST/STIS 1465 AA	HST/ACS-WFC F850LP	ALMA 286 GHz	VLA A-conf. 5 GHz
EHT				



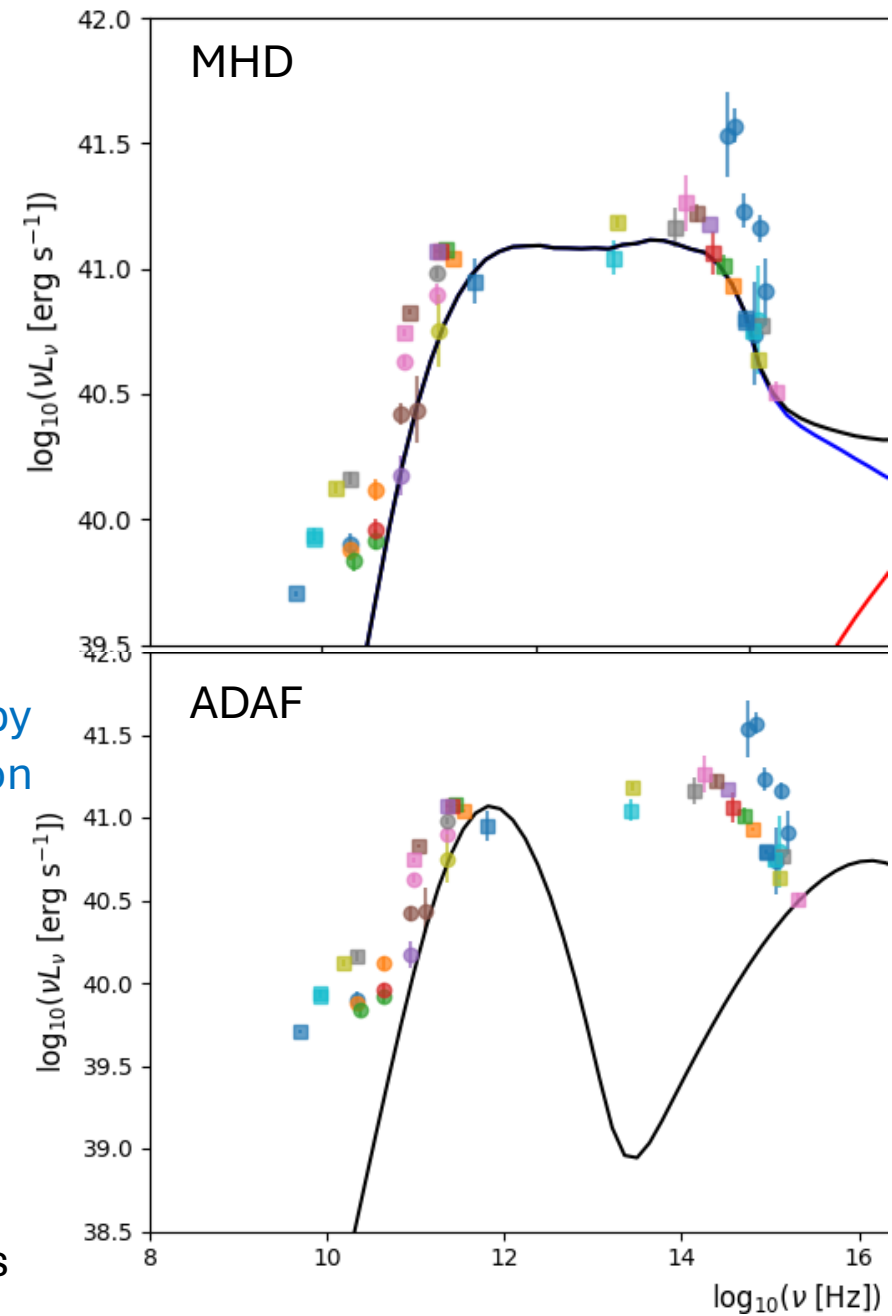
# Thermal synchrotron spectra

Much broader in the (MAD) MHD model and remarkably well matches the mm to UV data of M87

In the MAD state the emitting region characterised by a large range of plasma parameters:



The analytic models are essentially 1D, so single value of  $T_e$  at each radius



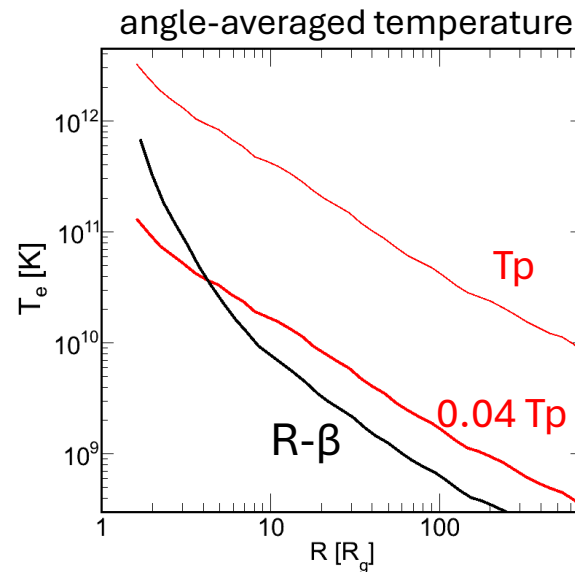
# R-β prescription

$$\frac{T_i}{T_e} = \frac{1}{1 + \beta^2} R_l + \frac{\beta^2}{1 + \beta^2} R_h$$

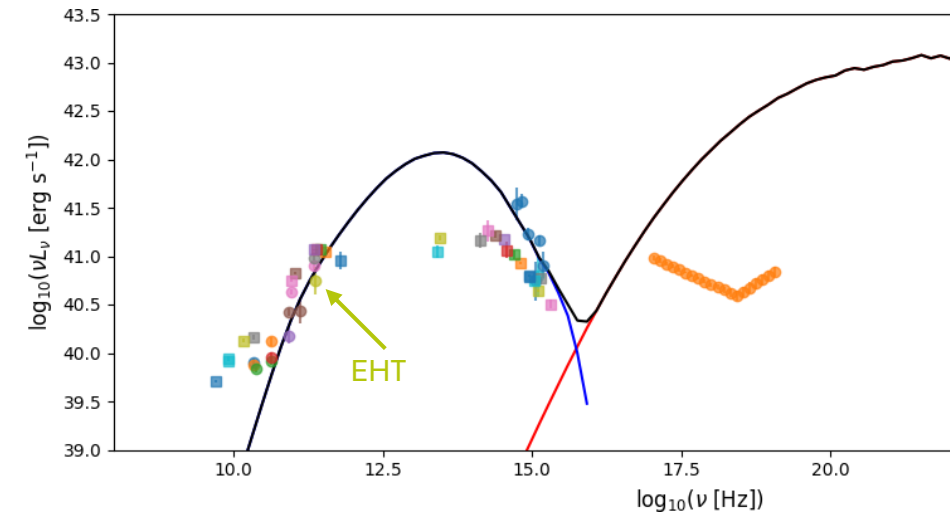
Mościbrodzka et al. (2016)

In MADs  $\beta \sim 1$  in the inner flow, implying  $T_e \sim T_p$  for  $R_{low} \sim 1$  (often applied in image analysis)

For  $R_{low} = 1$ ,  $R_{high} = 80$ :



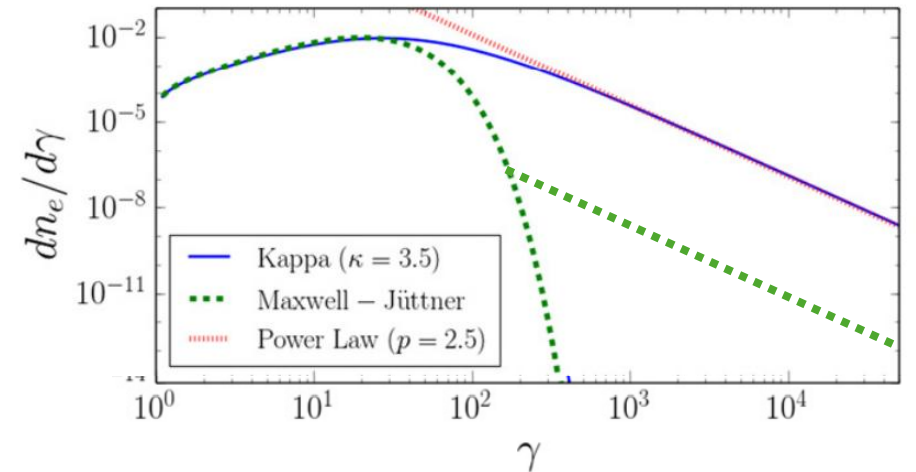
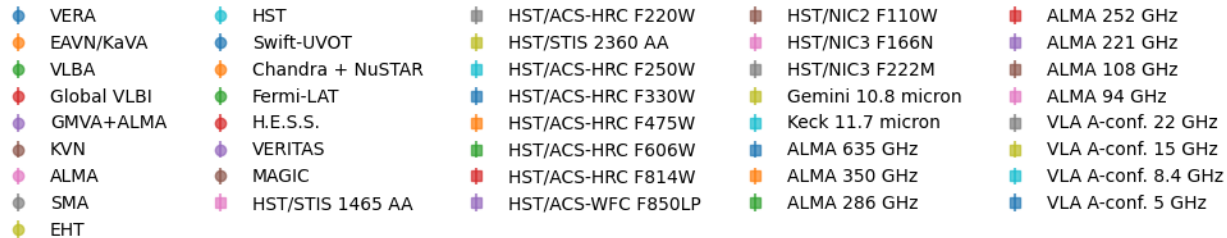
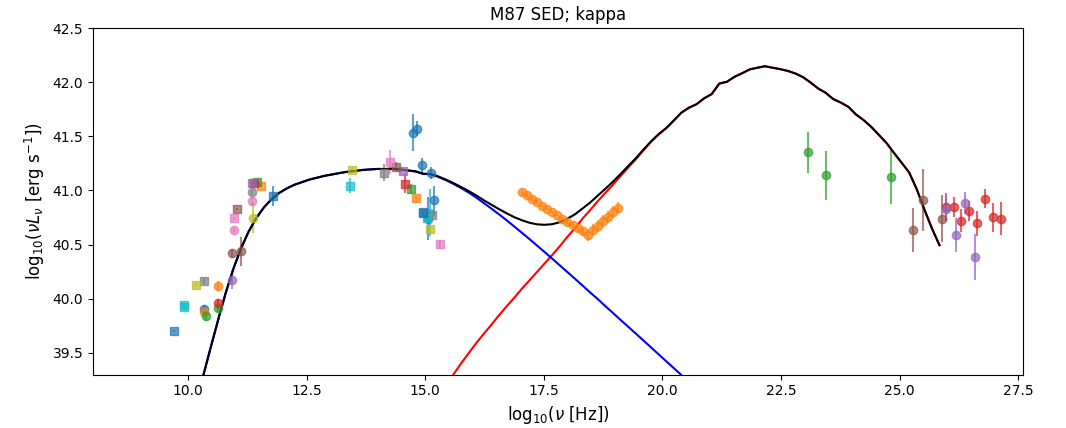
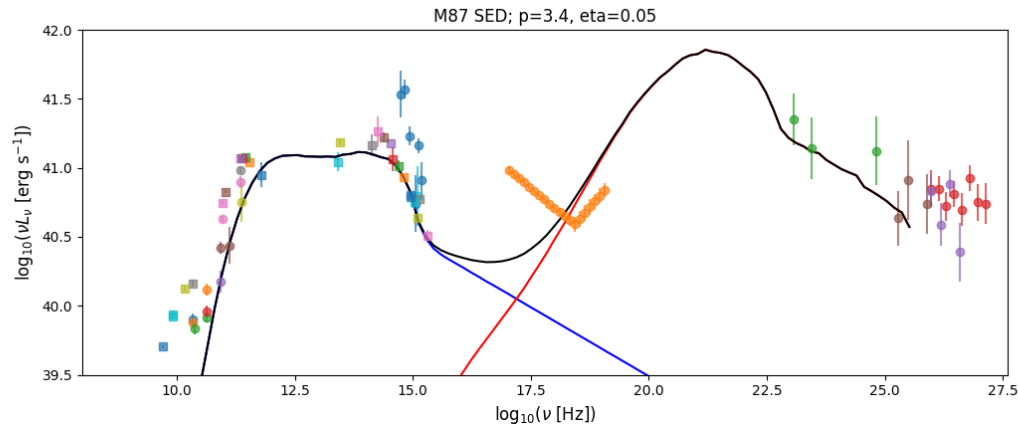
$\tau \sim 10^{-4}$ , but Compton amplification =  $\tau 16 \Theta_e^2 > 1$



bolometric luminosity  $\sim \dot{M} c^2$

# Hybrid model

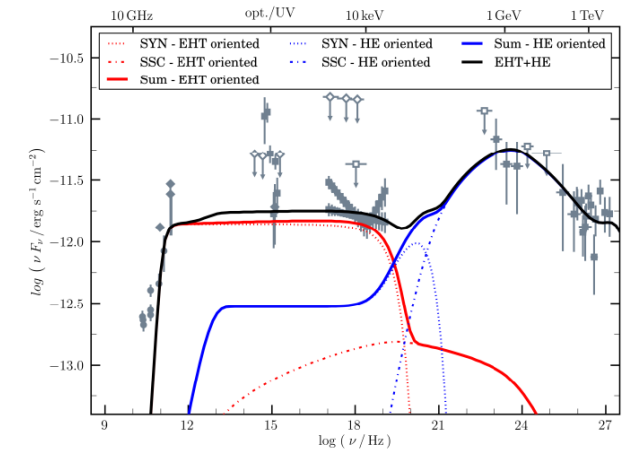
A 5% energy content of nonthermal electrons allows to explain the broadband spectrum (except for soft X-rays), VHE gamma-ray underpredicted likely due to less efficient cooling of electrons in the KN regime not taken into account (yet)



# Summary

Inner parts of MAD flows characterised by a wide spread of the flow parameters, in particular  $\Gamma_e$ , resulting in a broad thermal synchrotron spectrum, notably consistent with mm-UV SED of M87

Inclusion of a weak nonthermal tail allows to a large extent to explain the broadband SED, in a model simpler and more consistent with modeling of EHT images than multi-zone models, e.g.



EHT collab. (2025)

Work in progress, electron energy balance and polarization maps under implementation