

TDEs as potential candidates of high-energy neutrinos

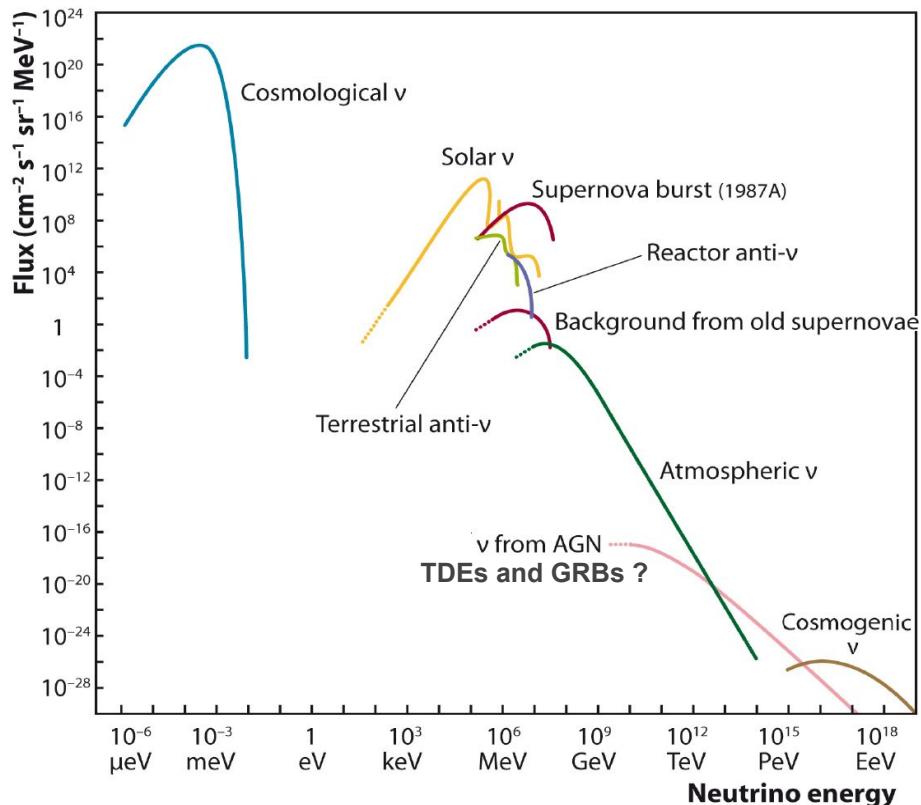


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Astrophysical neutrinos

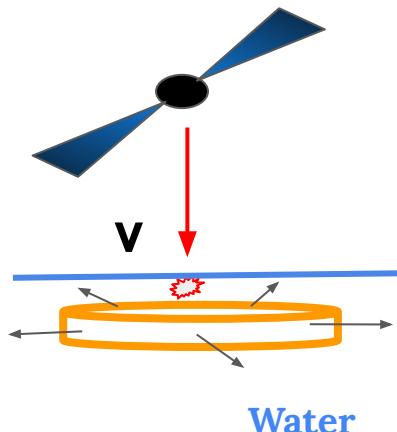
- Neutrinos: neutral, stable, and weakly interacting particles
- Unknown astrophysical origins and mechanism of production high-energy neutrinos
- Their interaction products are detected via Cherenkov radiation measured in PMTs



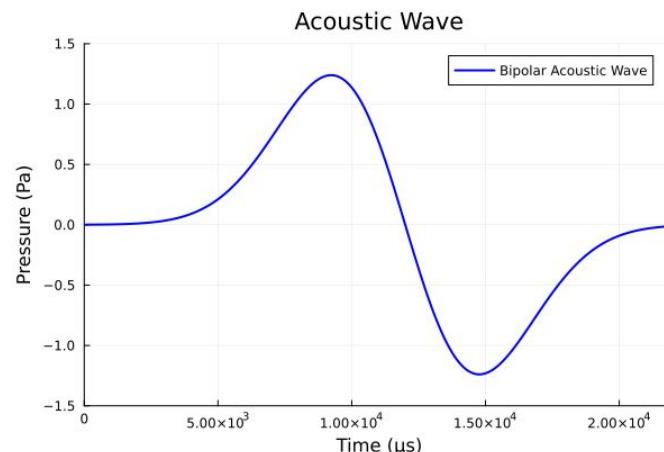
Innovative Acoustics Detection of UHE neutrino

Phenomenon

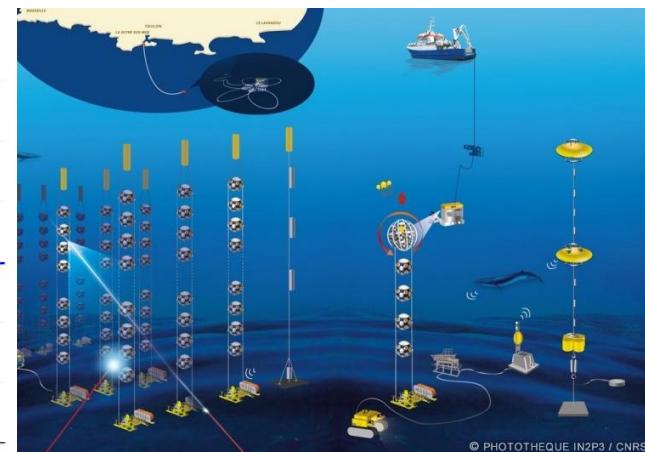
- Thermo-acoustic effect:
- Energy deposited by UHE ν heats water up



Pulse



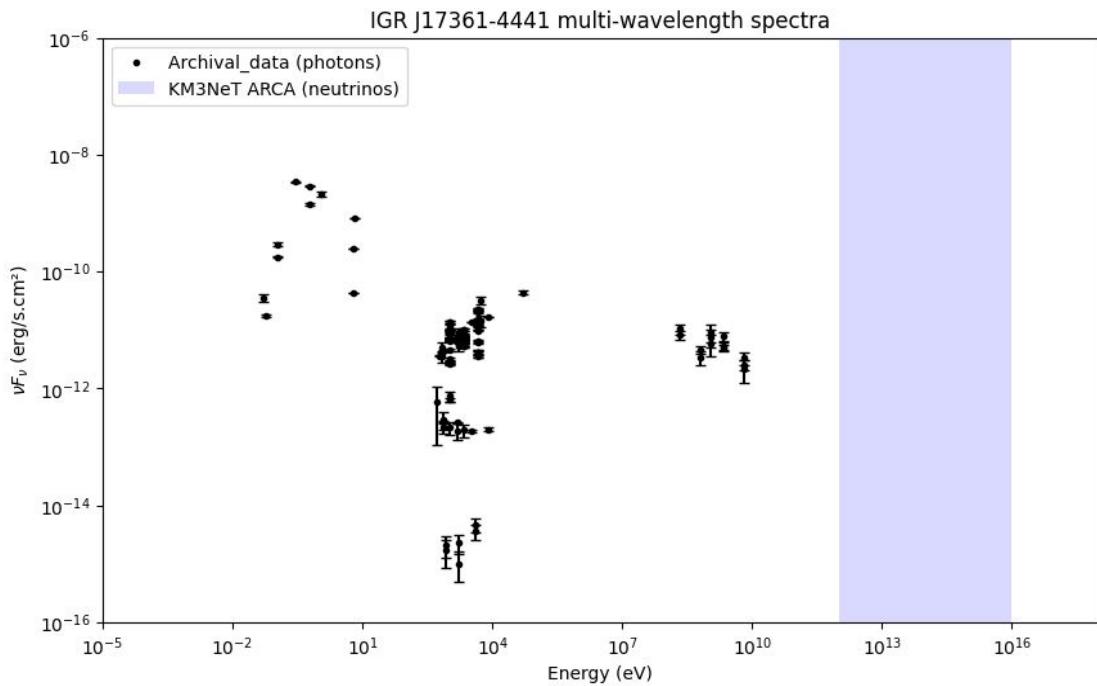
KM3NeT



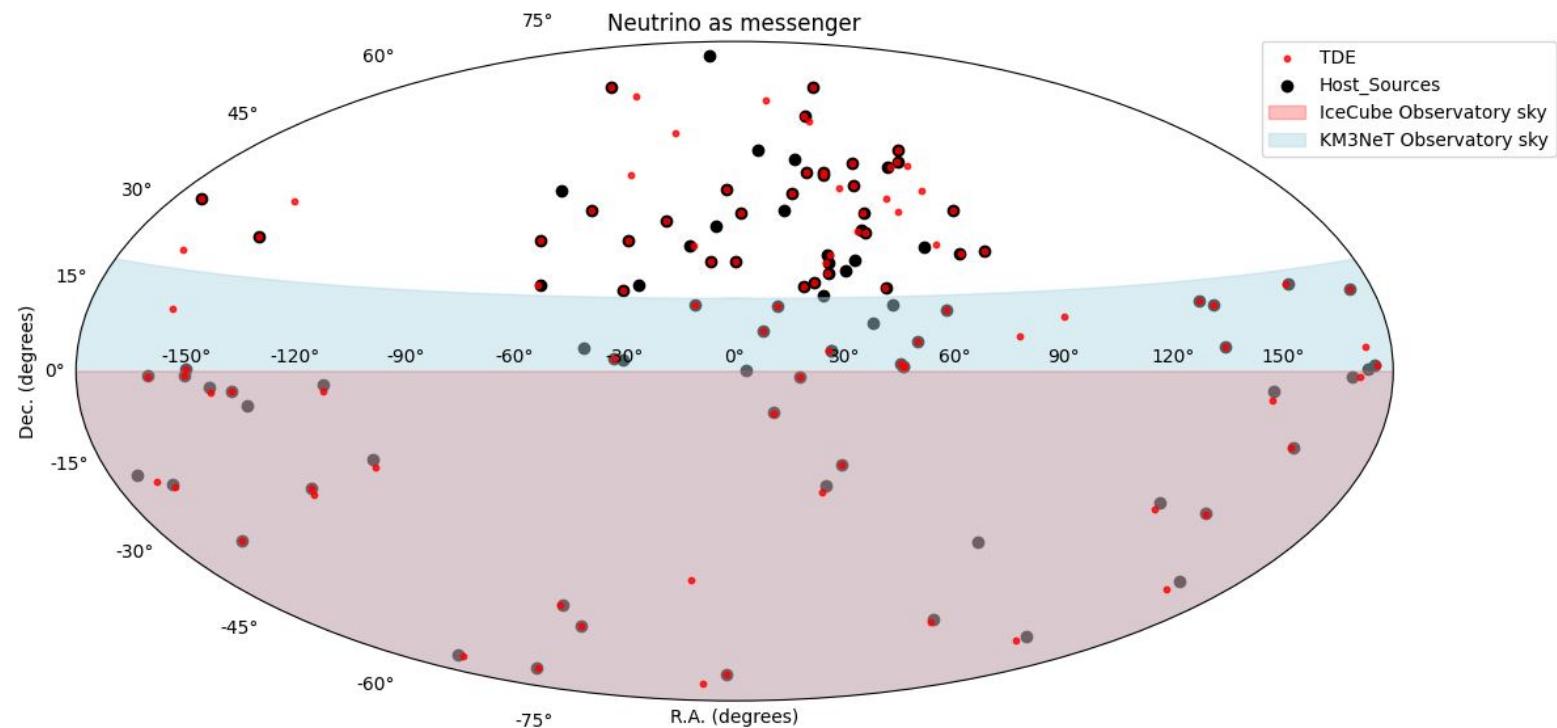
- Angular resolution < 0.1° (Optical detection) Eur. Phys. J. C 84, 885 (2024)
- Hydrophone sensitivity: -173 dB relative 1V/μPa up to 70kHz

Study of origins via Tidally Disrupted Events

- TDE: massive black hole and star system reaching Roche limit
- SED analysis for estimating the accretion efficiency
- Likelihood analysis for 102 raw TDE samples



Study of origins via Tidally Disrupted Events

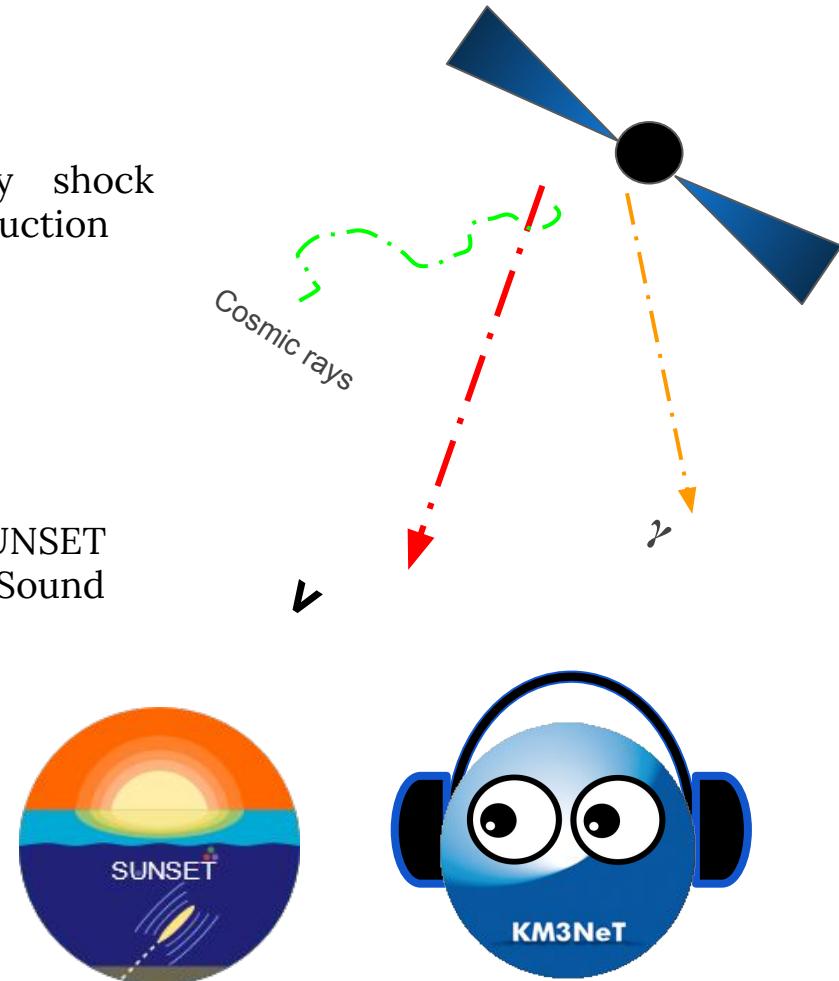


Based on data in catalog: Aaron Goldtooth et al 2023 PASP 135 034101

Summary and conclusion

- TDE pristine class of transients to study shock interaction, jet formation, and HE neutrinos production
 - ◆ GW can be detected by LIGO and DECIGO
 - ◆ HE neutrinos with KM3NeT

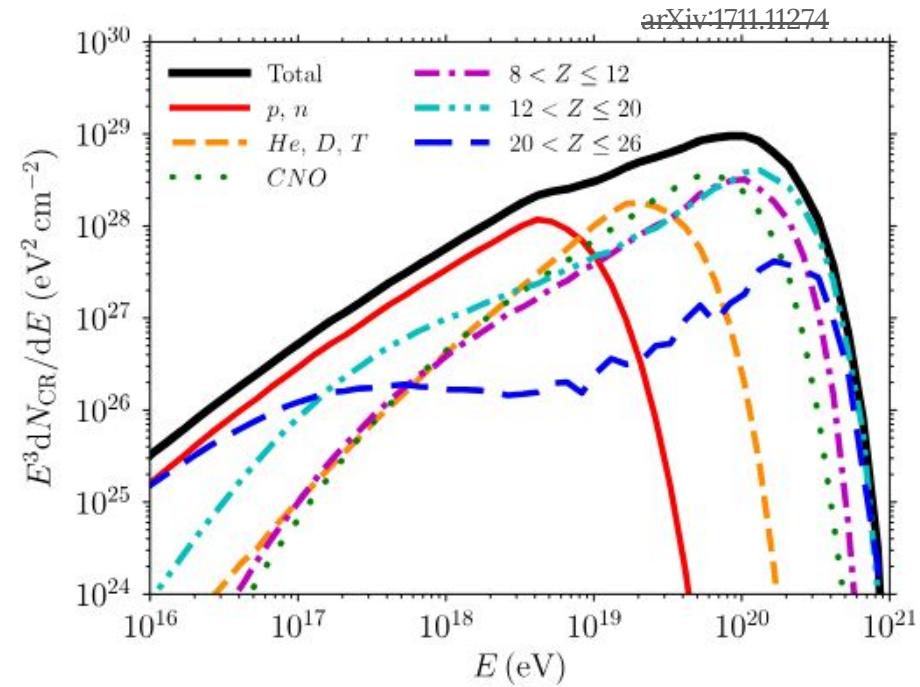
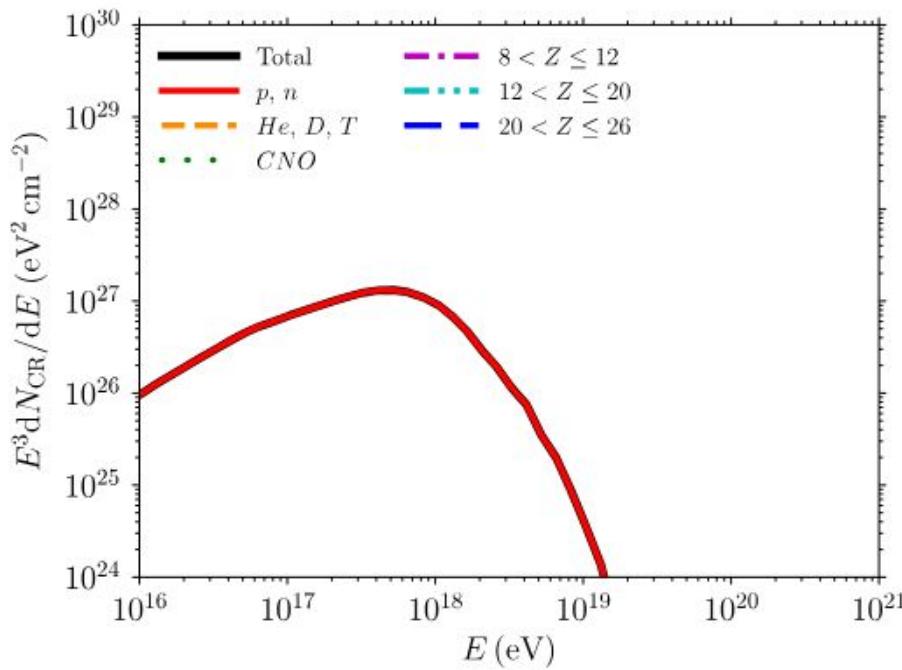
- Developing dedicated acoustic software: SUNSET (Simulation of Underwater Neutrino-induced Sound Emission and Transport)
 - ◆ Signal generation and propagation
 - ◆ Noise Analysis



Back up

Supplementary material

- Cosmic-ray spectra for one source with pure iron injection
- Most highly super-Eddington TDEs are powered by the complete disruption of massive star



Supplementary material

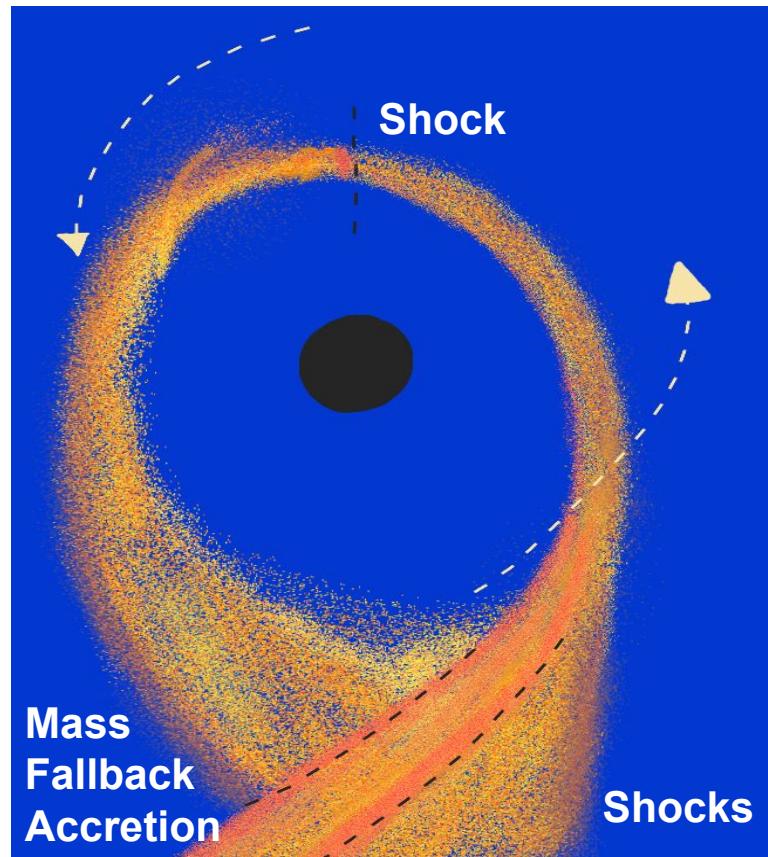
- High-energy neutrino emission is correlated with temporal and spatial emissions across all the multi-messenger
- Tidally disrupted events are one of the potential candidates of high energy neutrinos

Daniel J. Price et al 2024 ApJL 971 L46



Supplementary material

- Multi-messenger properties:
 - Spectral classification by UV - optical color diagram into TDE-H, TDE-H+He, and TDE-He
 - At X-ray and radio energies non-thermal emissions
 - Very high-energy neutrinos of TeV and PeV
 - Gravitational waves candidate up to 10 Hz



References:

- Suvi Gezari, ARA&A, 50, 67 (2012); DOI: 10.1146/annurev-astro-081811-125848
- Spiering, C. (2012) <https://doi.org/10.48550/arXiv.1207.4952>
- Guépin, C., Kotera, K., Barausse, E., Fang, K., & Murase, K. 2018, A&A, 616, A179 DOI: 10.1051/0004-6361/201732392
- Pfister, H., Toscani, M., Wong, T. H. T., Dai, J. L., Lodato, G., & Rossi, E. M. arXiv:2103.05883 (DOI: 10.48550/arXiv.2103.05883)
- Goldtooth, A., Zabludoff, A. I., Wen, S., Jonker, P. G., Stone, N. C., & Cao, Z. March 13, 2023 DOI 10.1088/1538-3873/acb9bc
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- <https://tools.ssdc.asi.it>