Global simulations of relativistic jets from accreting black holes 2024 report of Krzysztof Nalewajko

- A&A, 686, A56
- 2025, submitted to A&A
- new Ph.D. student: Wen Xuan Sia (stipend: CAMK -> OPUS 27)
- master student: Mateusz Kapusta
- new collaborators: Bart Ripperda, Alexander Philippov
- new grant NCN OPUS 27
- 2 invited conference talks (SCEECS*/NYU, Purdue), 4 seminars (Grenoble, UMK, CAMK, St. Louis)
- popular talk in CNK planetarium
- supercomputing allocation: 1.5 MSU at Ares (Cyfronet ACK)
- secretary of the Scientific Council

K. Nalewajko, M. Kapusta, A. Janiuk "Chaotic magnetic disconnections trigger flux eruptions in accretion flows channeled onto magnetically saturated Kerr black holes", 2024, A&A, 692, A37 D. A. Kann, N. E. White, et al. "Fires in the deep: The luminosity distribution of early-time gamma-ray-burst afterglows in light of the Gamow Explorer sensitivity requirements", 2024,

K. Nalewajko "Tension of toroidal magnetic field in reconnection plasmoids and relativistic jets",

* Simons Collaboration on Extreme Electrodynamics of Compact Sources



Tension of toroidal magnetic field in reconnection plasmoids and relativistic jets

merger of 2 plasmoids

magnetic tension enhances energy density in plasmoids and jets, combined to enhance luminosity of blazar flares submitted to A&A

PIC simulations of relativistic reconnection energy density: magnetic vs plasma

relaxed monster plasmoid





proposed lateral structure of relativistic jets



main spine

paraboloidal diverging $B_{\rm p}, {\rm v}_{\rm p}$ Γ, B_{ϕ} decreasing with R bulk of energy flux





Chaotic magnetic disconnections trigger flux eruptions in accretion flows channeled onto magnetically saturated Kerr black holes

with Mateusz Kapusta and Agnieszka Janiuk; A&A, 692, A37; https://users.camk.edu.pl/knalew/aa50490-24/

prograde (a=0.9)



retrograde (a=-0.9)



Extreme-resolution simulation of relativistic jet



with Mateusz Kapusta (OAUW; master project), Bart Ripperda (Toronto) and Alexander Philippov (Maryland)

data from GRMHD simulation performed with H-AMR, presented in Ripperda+22 numerical resolution 5376 x 2304 x 2304 (r, θ, ϕ) ; no AMR relativistic jet from magnetically saturated accretion (MAD) effects of magnetic flux eruptions large samples of integrated field lines

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